



- Physics case
- The n_TOF experimental program on ⁷Be
 - **–** ⁷Be(n,p) measurement (2016)
 - ⁷Be (n,α) measurement (currently ongoing)
- Conclusions



Big Bang Nucleosynthesis (BBN), together with Hubble expansion and Cosmic Microwave Background Radiation is one of the cornerstones for Big Bang Theory.

BBN gives the sequence of nuclear reactions leading to the synthesis of light elements up to Na* in the early stage of Universe (0.01-1000 sec)

At his first formulation, it depended on 3 parameters:

-the baryon-to-photon ratio η,
-the number of species of neutrino ν,
-the lifetime of neutron τ.

Nowadays **BBN is a parameter free theory****, being the **cross-sections** of reactions involved the only input to the theory.

* A.Coc et al., The Astrophysical Journal, 744:158 (2012) **D.N. Schramm and T.S Turner, Rev. Mod. Phys 70 (1998) 303





The Cosmological Lithium Problem

0.26

0.25

0.24

0.23

 10^{-3}

10-4

Х

D/H

BBN successfully predicts the abundances of primordial elements such as ⁴He, D and ³He.

A serious discrepancy (factor 2-4) between the predicted abundance of ⁷Li and the value inferred by measurements (Spite et al, many others.)



baryon-to-photon ratio η

baryon density $\Omega_{\rm b}h^2$ 10^{-2}

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* R.H.Cyburt et al., Journal of Cosmology and Astroparticle Physics 11 (2008) 012

** A.Coc et al., The Astrophysical Journal, 744:158 (2012)



Approximately 95% of primordial ⁷Li is produced from the electron capture decay of ⁷Be ($\underline{T}_{1/2} = 53.2 \text{ d}$).

⁷Be is destroyed via (n,p) and (p,x), (d,x), (³He,x), ... reactions Very small contribution of the (n, α) reactions according to **estimated** cross section.



⁷Be(n,α)

Two beam lines and experimental areas available at n_TOF, EAR1 and EAR2



⁷Be(n,α)

Two beam lines and experimental areas available at n_TOF, EAR1 and EAR2





2 Silicon detectors (5x5 cm² area, 20 μm and 300 μm thickness) to detect 1 MeV and 1.4 MeV protons. <u>Telescope Technique</u>



Purification of sample needed: ISOLDE

- 100 ng of ⁷Be (material from water cooling of SINQ spallation source at **PSI**)
- Offline mass separation required at ISOLDE ≤1 day (starting from 100 GBq of ⁷BeNO₃)
- Implantation on C backing





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The (n,α) reaction produces **two a-particles** emitted back-toback with **several MeV energy** (Q-value=19 MeV)

2 Sandwiches of silicon detector (140 μ m,3x3cm²) with ⁷Be sample in between directly inserted in the neutron beam

Coincidence technique: strong background rejection





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Biggest challenge related to the sample:

•For a reasonable count-rate, a few μg of ⁷Be are needed

•The short half-life of ⁷Be (53.2 d) **very large specific activity** (~1.3 GBq/µg of 478 keV g-rays

•Difficult to find and handle

At n_TOF:

•Up to 10 µg of ⁷Be from cooling water of SINQ spallation target at PSI (Zurich)

•Chemically purified before depositing (on polyethilene and Al backing)

•Complex mechanics and procedure for mounting in the hot-cell







Raw signals from one of the sandwich (det1&det2)





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Amplitude distribution of coincidence events

NFN



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NFN

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• Uncertainties in nuclear data strongly affect the Big Bang Nucleosynthesis calculations for the abundance of ⁷Li and could probably explain (at least partially) the Cosmological Lithium Problem.

• Experimental program aiming at measuring at n_TOF-EAR2 for the first time the ⁷Be(n,p) and ⁷Be(n, α) cross sections in the whole range of interest for BBN, with the aim of reducing uncertainties in nuclear data used in calculations, thus setting stronger constraints to BBN theory and on CLiP.

• Taking advantage of the well suited features of the facility (high instantaneous neutron flux, work sector type A):

- The measurement of ⁷Be(n,p) reaction cross-section is scheduled for 2016 in EAR2
- The measurement of ${}^{7}Be(n,\alpha)$ reaction cross-section is **currently ongoing** and <u>preliminary results are extremely encouraging</u>.