



EXPLORING PARTICLE-ANTIPARTICLE ASYMMETRY IN NEUTRINO OSCILLATION

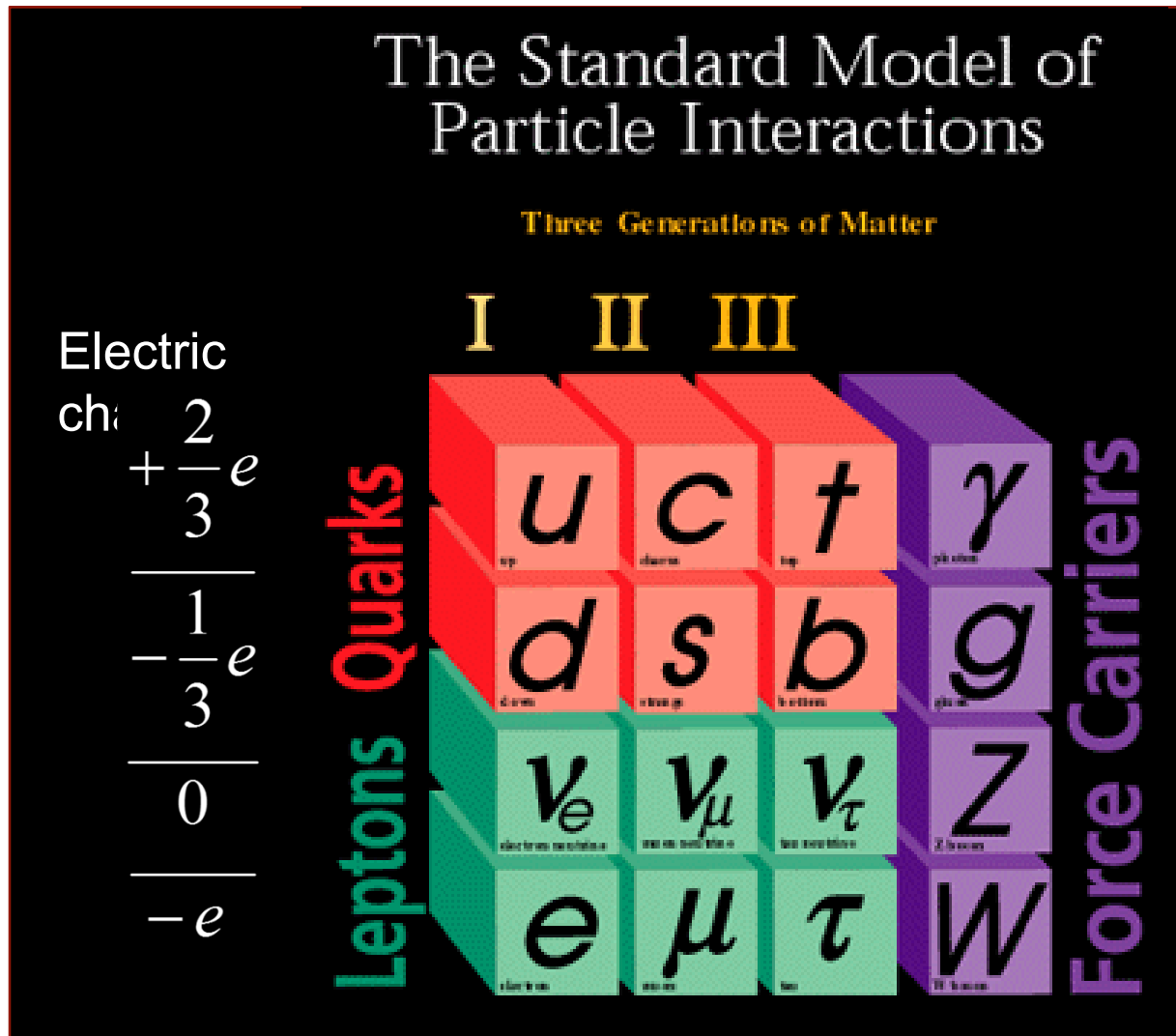
Atsuko K. Ichikawa, Kyoto University

INTRODUCTION OF MYSELF

- Got PhD by detecting doubly-strange nuclei using emulsion
- After that, working on accelerator-based long-baseline neutrino oscillation experiments in Japan, especially on neutrino production, neutrino detector in accelerator-site and analysis.
- Recently, started a high-pressure Xenon gas project for double-beta decay search



CONSTITUENTS OF THIS WORLD



How can we distinguish btw.

u, c and t

d, s and b

ν_e , ν_μ and ν_τ

e, μ and τ

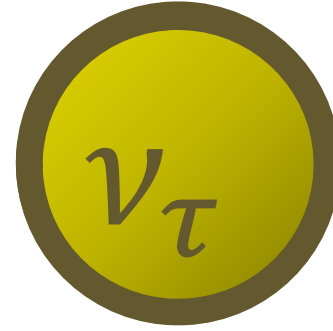
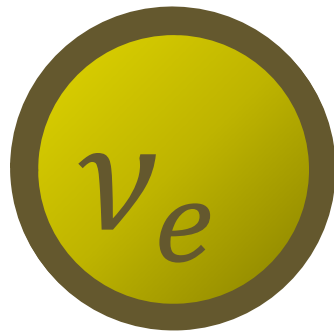
Same spin, same charge...

Only by mass!
Except for ν 's.

HOW CAN WE DISTINGUISH NEUTRINOS?

- IT IS TWO SIDES OF COINS-

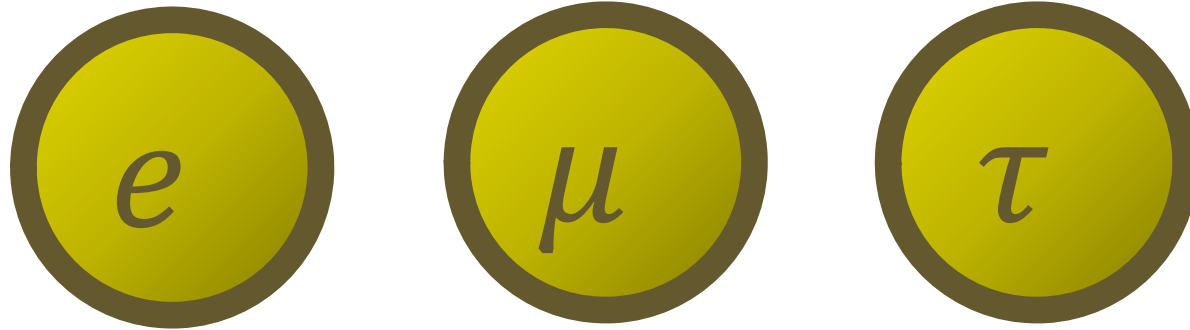
Neutrinos do interact with matter and



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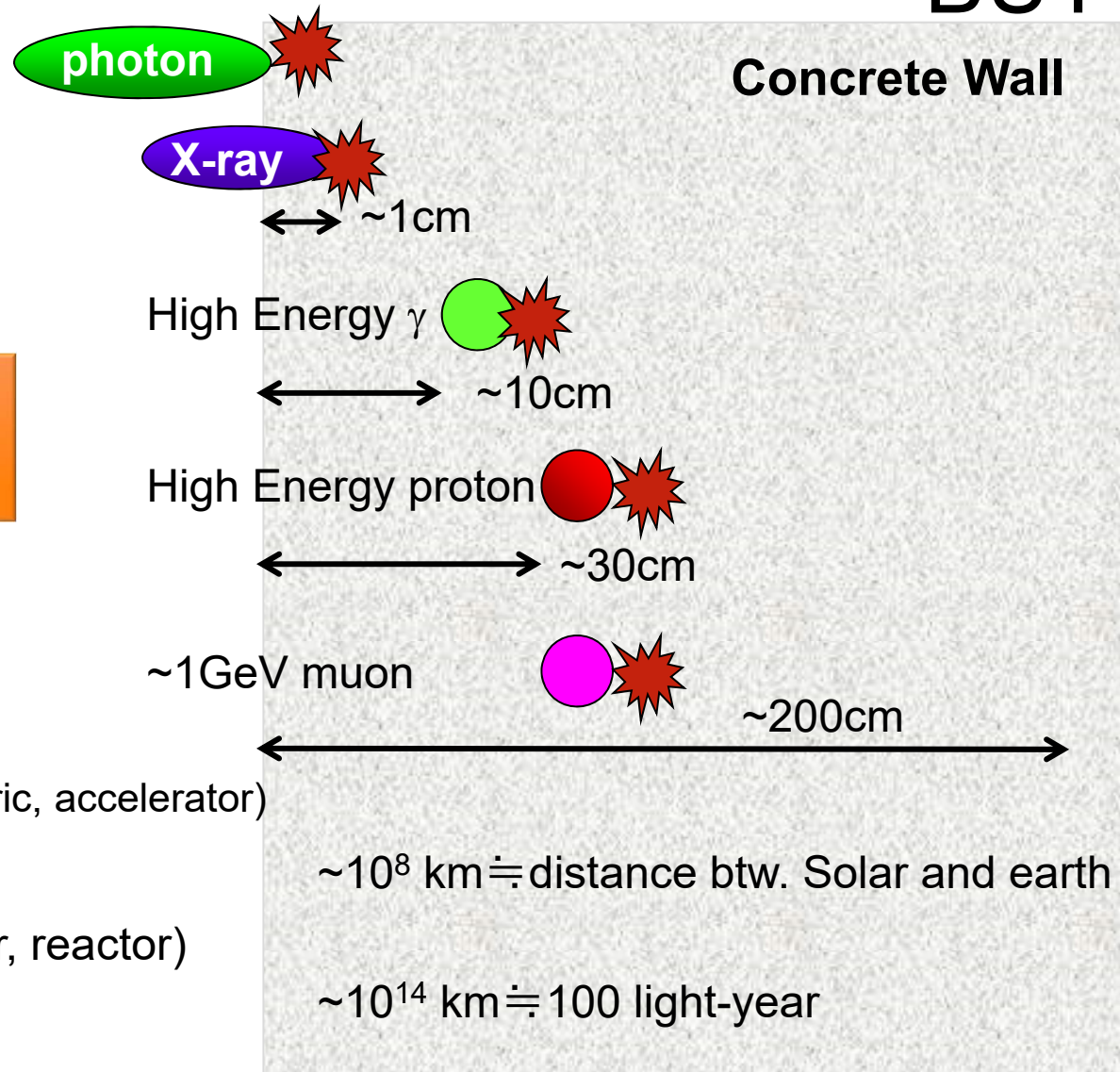


- **An electron neutrino** changes to **an electron**.
- **A muon neutrino** changes to **a muon**.
- **A tau neutrino** changes to **a tau**.

We call this categorization 'flavor'.

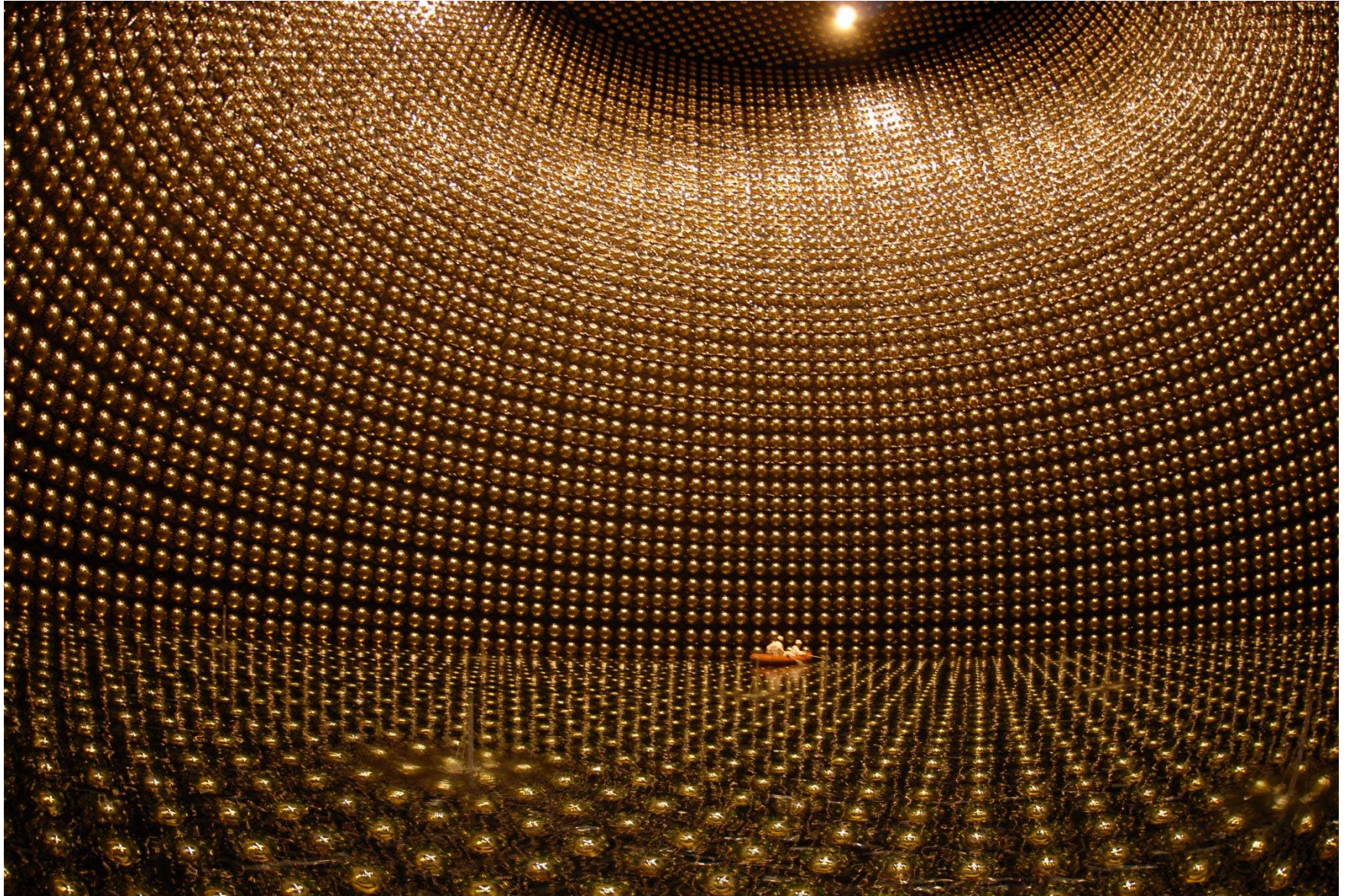
And it was believed that electron neutrino only changes to electron, never into muon nor tau before the neutrino oscillation was found.

NEUTRINOS DO INTERACT, BUT

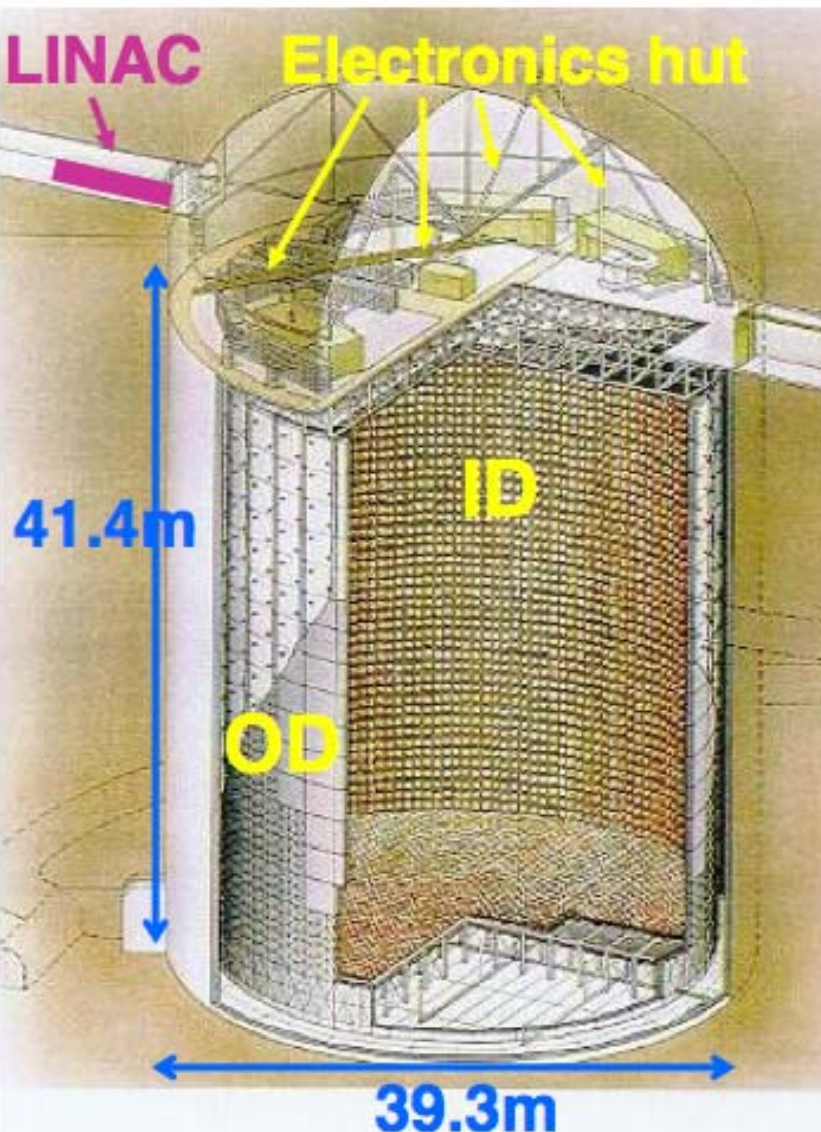


Mean Free path
of particles

SUPER-KAMIOKANDE



SUPER-KAMIOKANDE



~11000 x 20inch PMTs (inner detector, ID)

Since April 1996

Water Cherenkov detector w/ fiducial volume 22.5kton

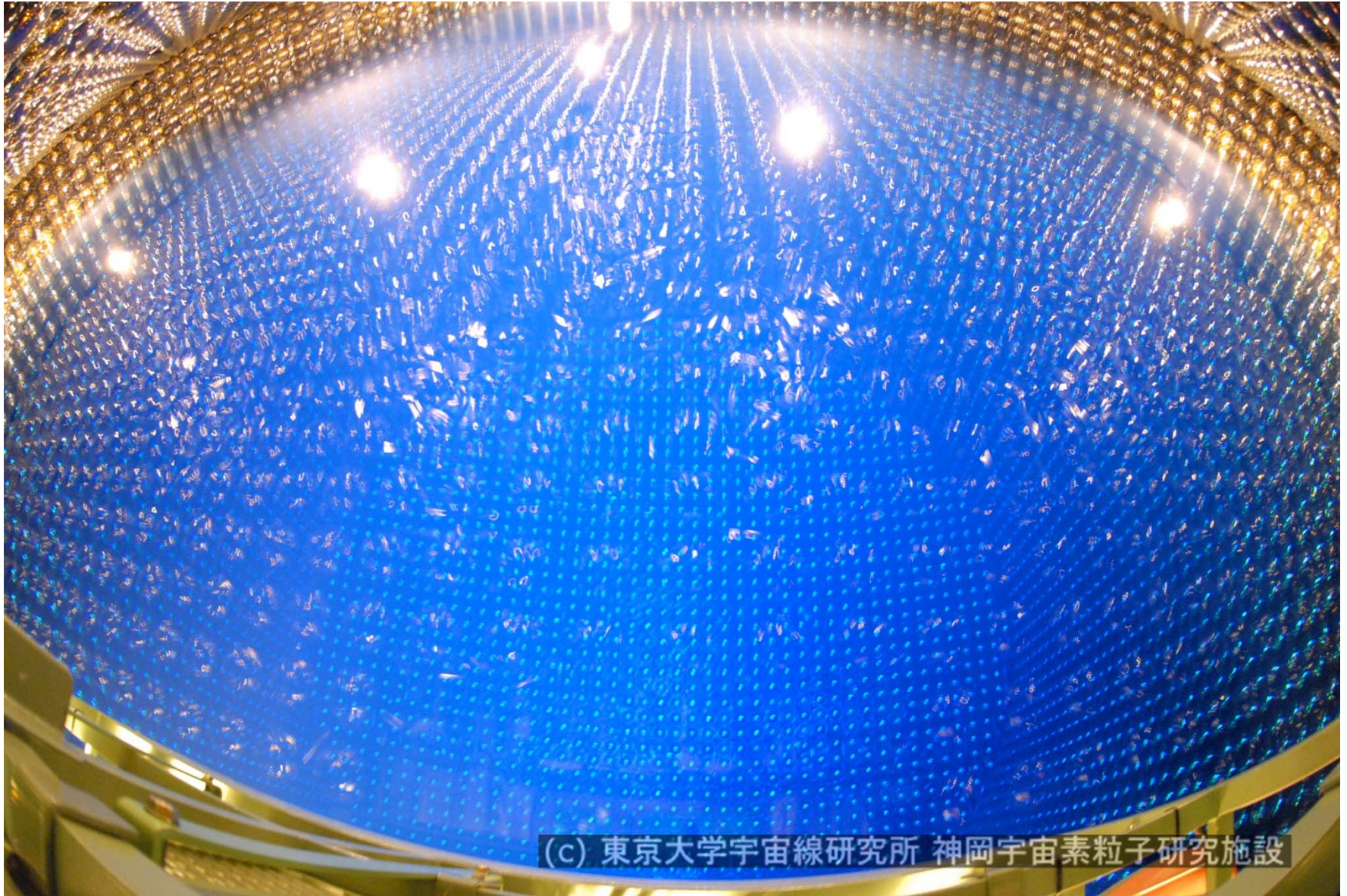
Detector performance is well-matched at sub GeV

Excellent performance for single particle event

Good e-like(shower ring) / μ -like separation
(next page)

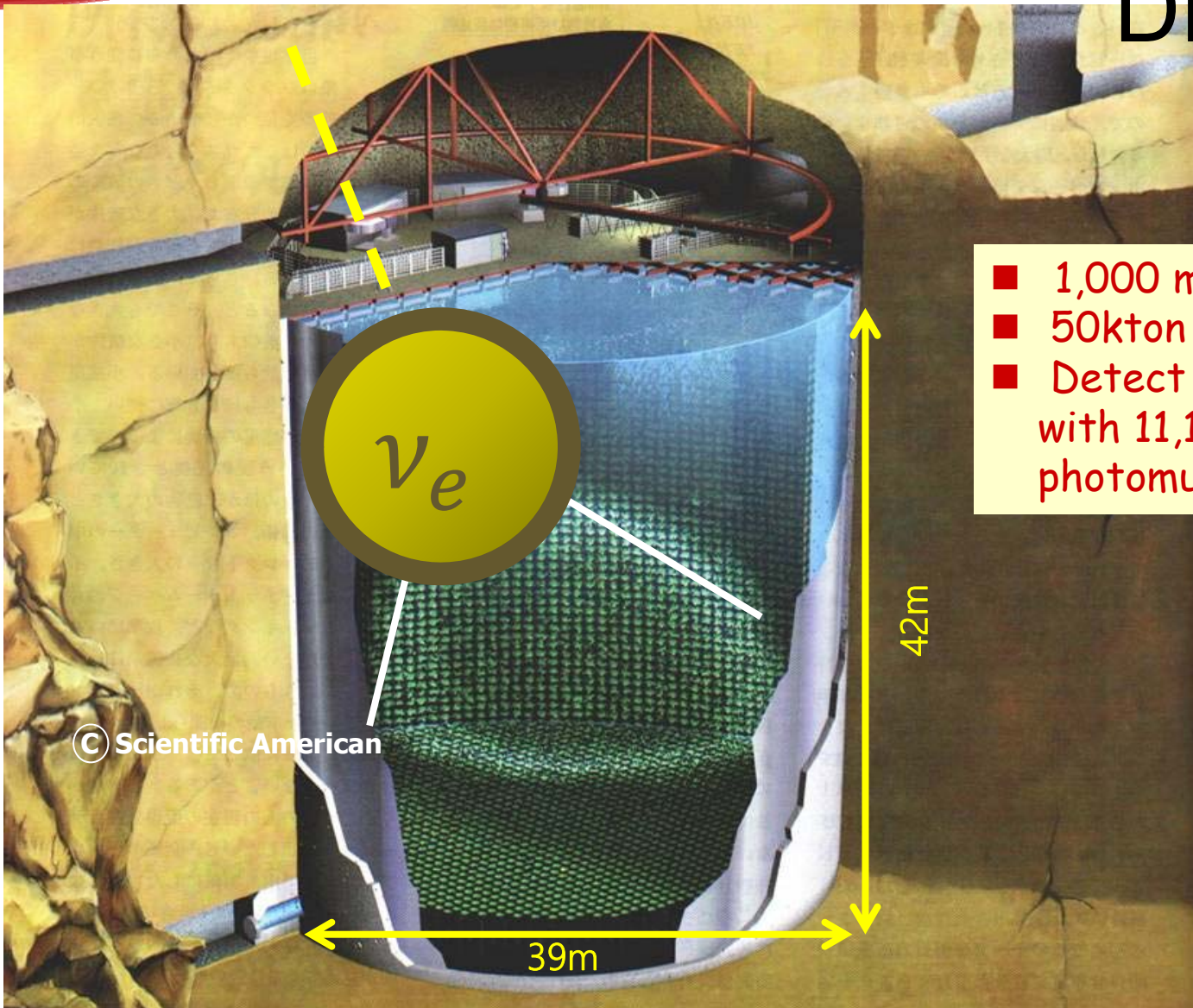


SUPER-KAMIOKANDE



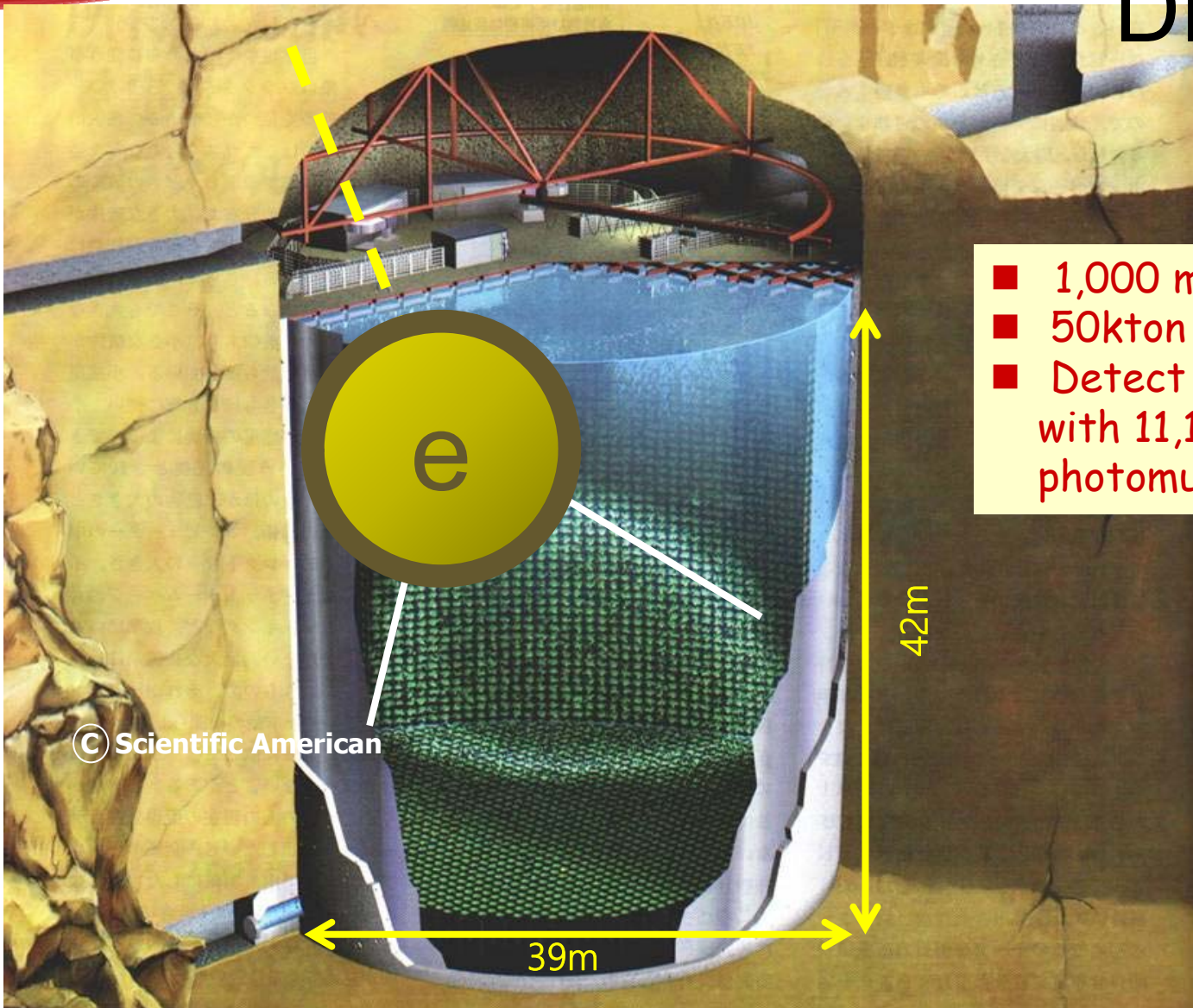
(c) 東京大学宇宙線研究所 神岡宇宙素粒子研究施設

SUPER KAMIOKANDE DETECTOR



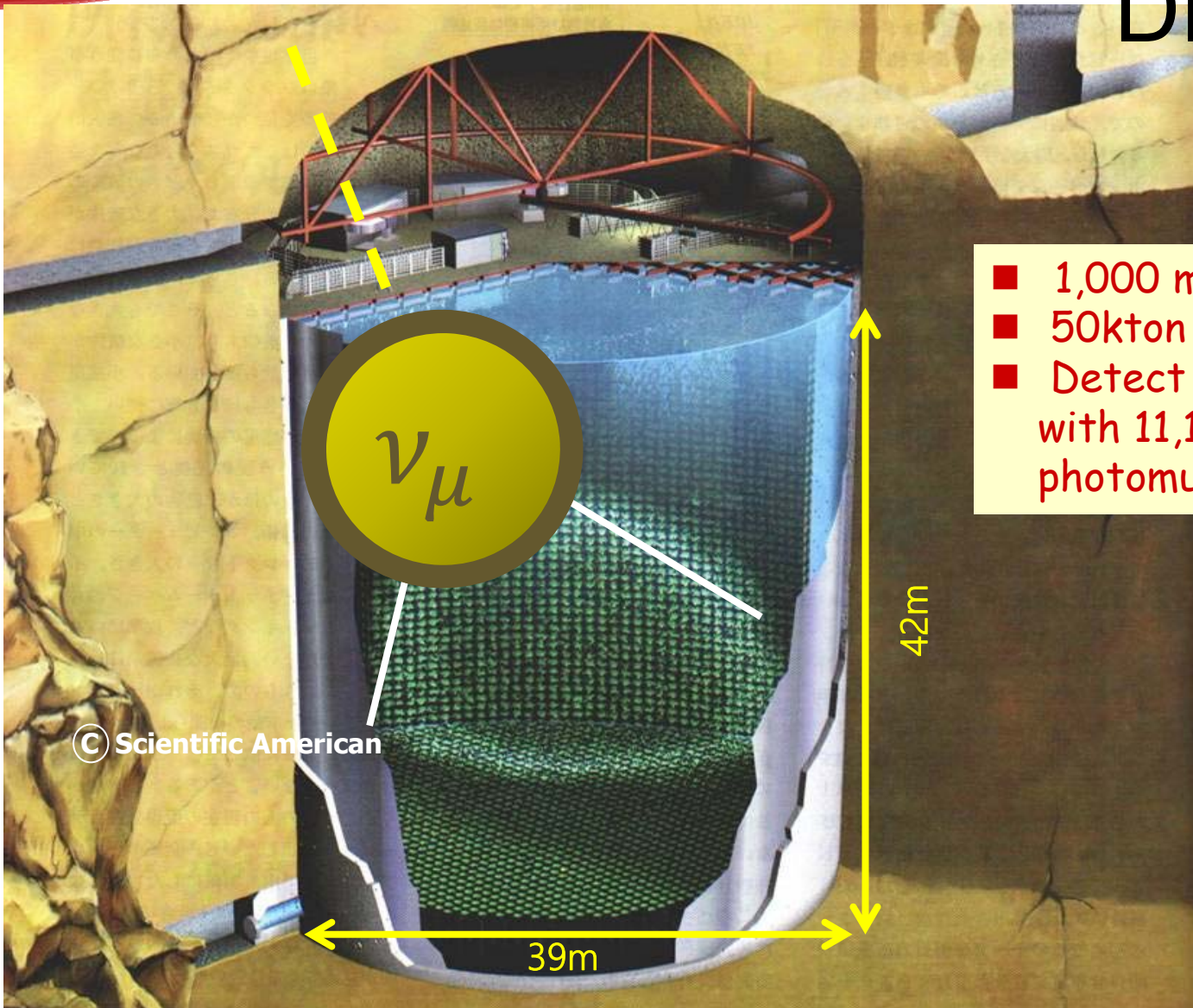
- 1,000 m underground
- 50kton water tank
- Detect Cherenkov photons with 11,146 20 inch photomultiplier tubes

SUPER KAMIOKANDE DETECTOR



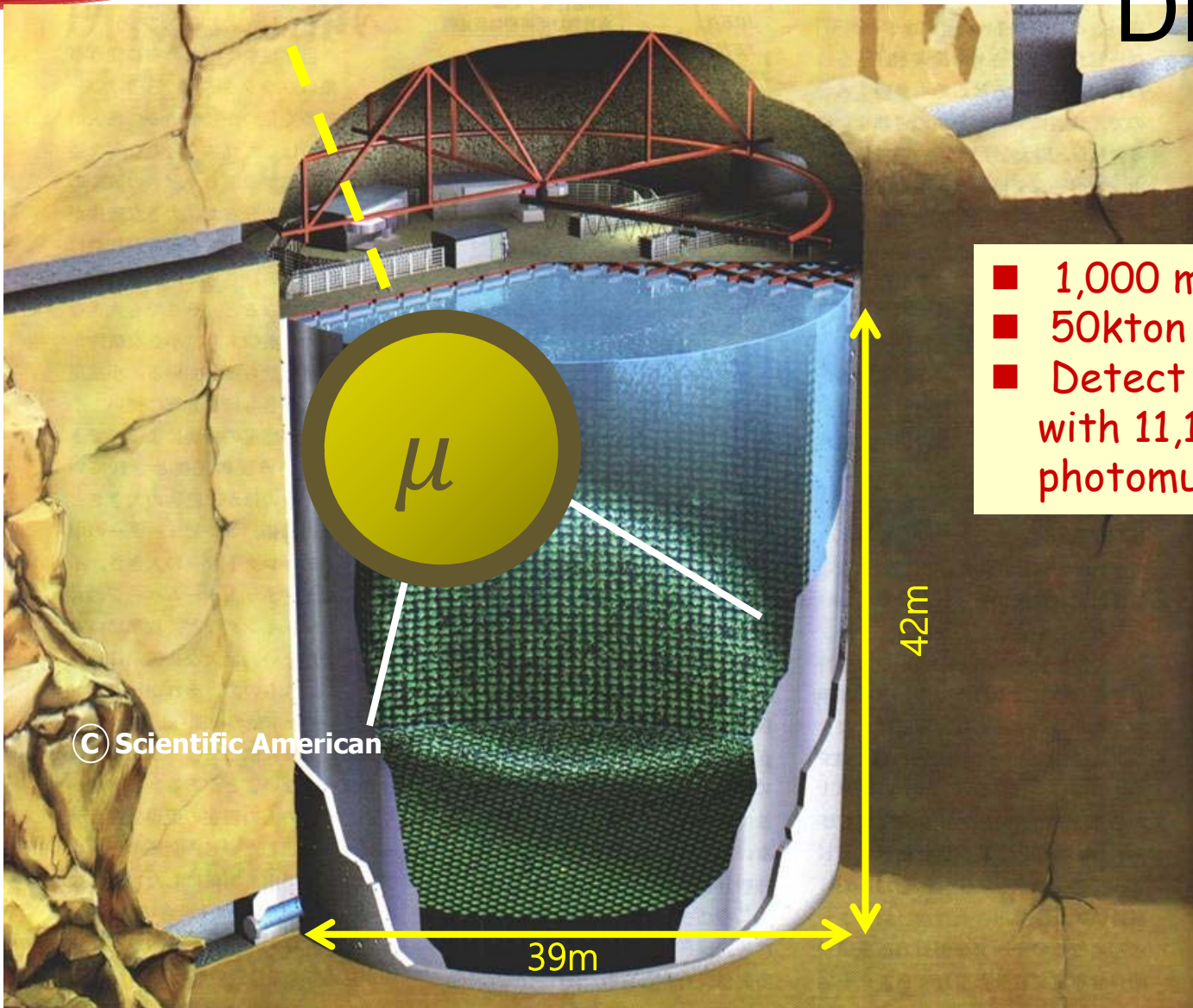
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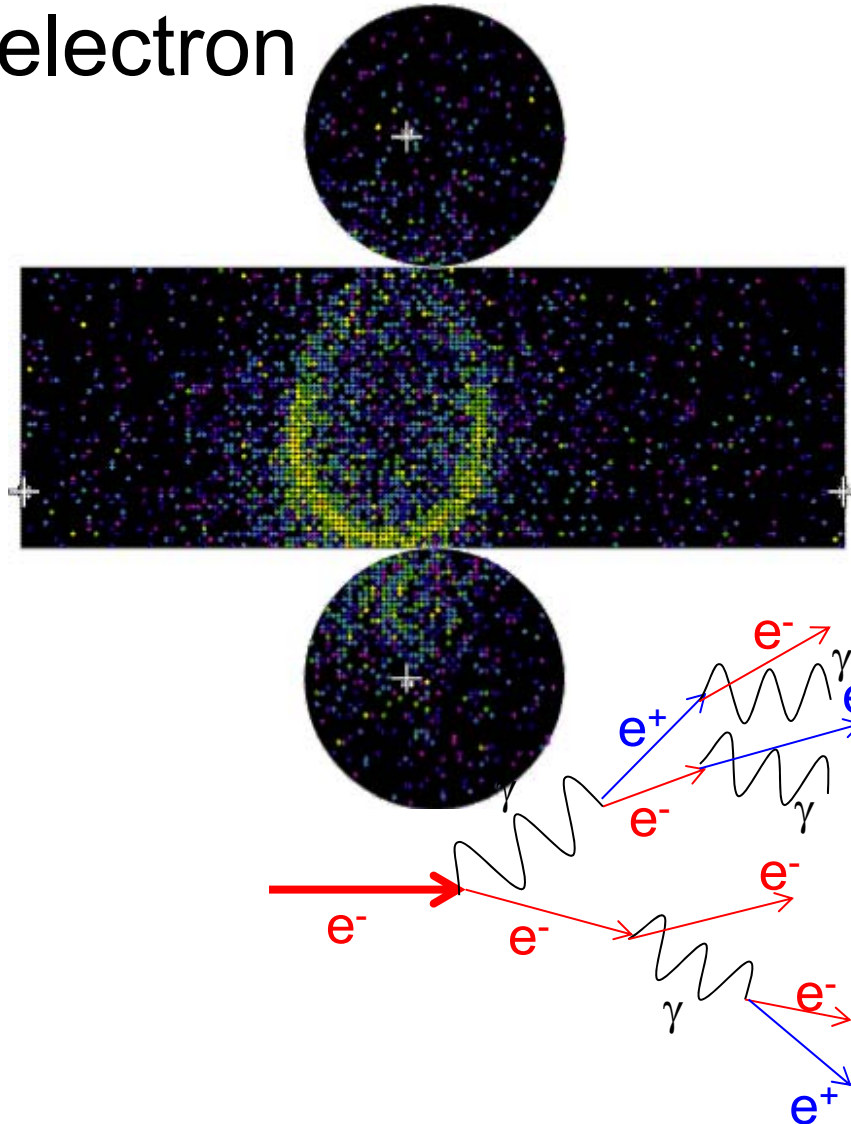
SUPER KAMIOKANDE DETECTOR



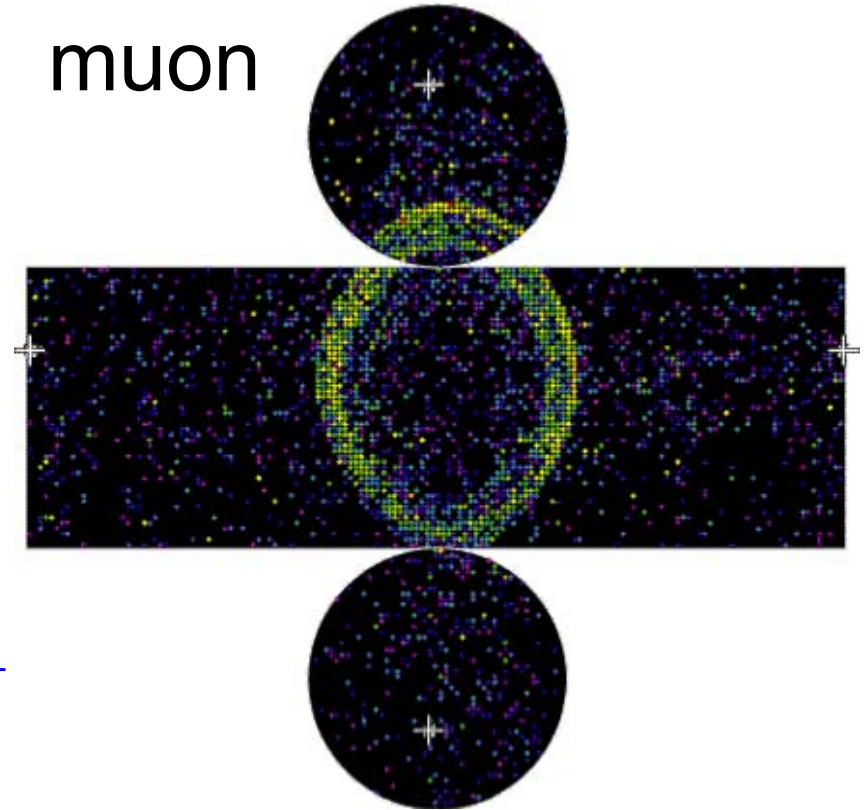
- 1,000 m underground
- 50kton water tank
- Detect Cherenkov photons with 11,146 20 inch photomultiplier tubes

THEY CAN BE DISTINGUISHED.

electron

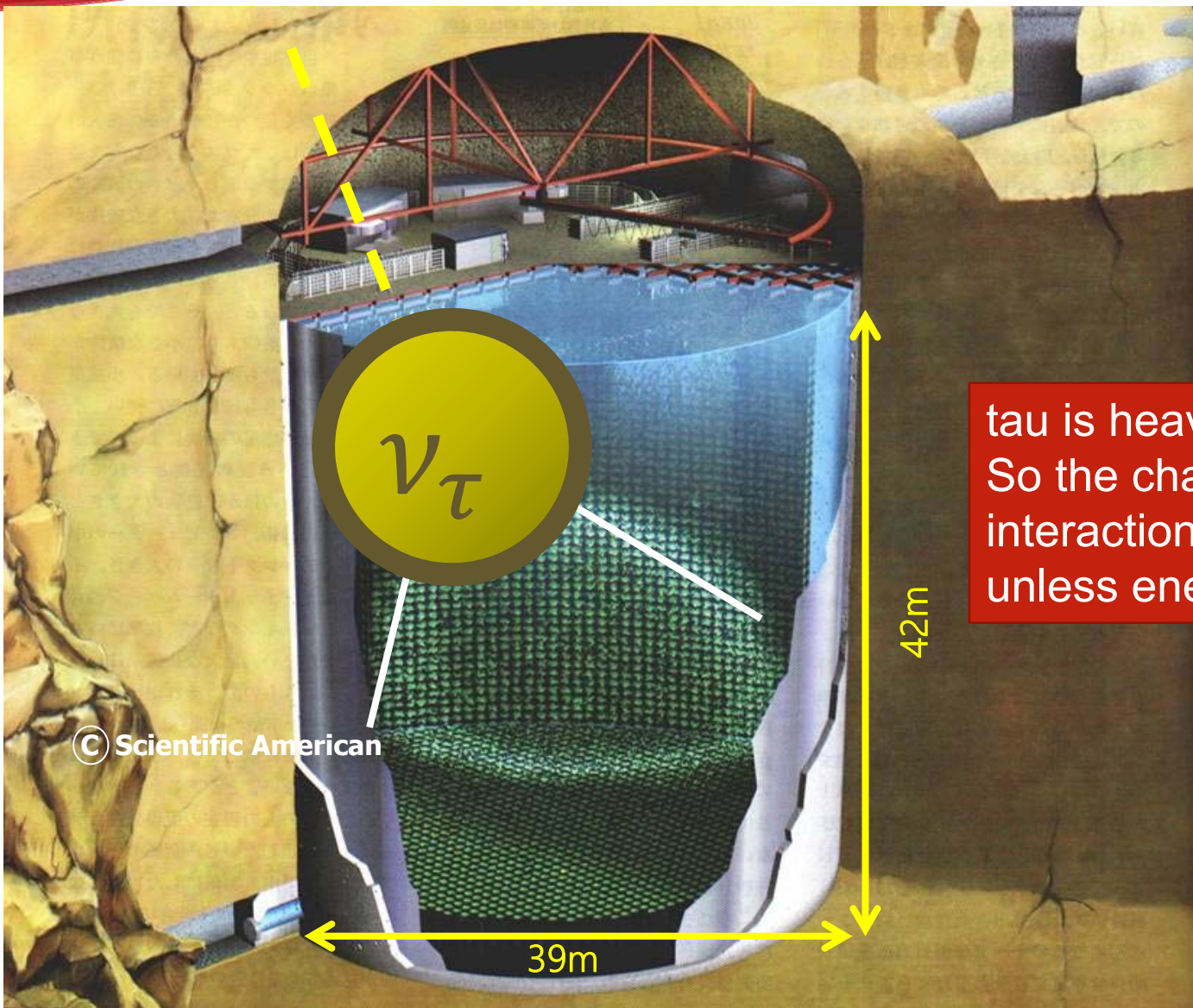


muon



SK MC event display

TAU NEUTRINO....



tau is heavy.
So the charged current
interaction dose not occur
unless energy $> \sim 3.5$ GeV.

MIXING BTW. FLAVOR AND MASS

We know that there are three types of charged lepton (e, μ, τ), distinguishable only via **mass**.

Then, we found that there are three types of neutrinos, distinguishable via interaction w/ matter.

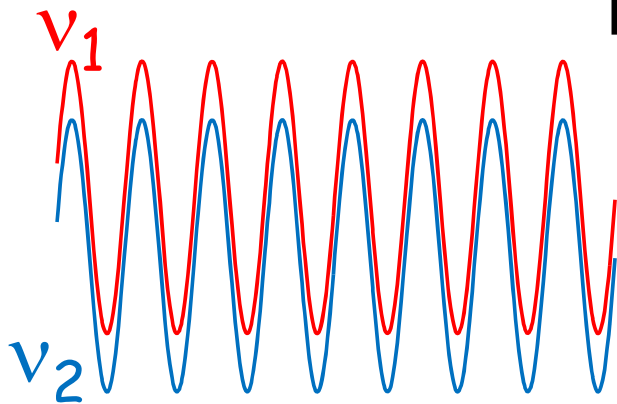
IF NEUTRINOS HAVE MASSES, there is no need for three types to be mass eigenstates.

$$|\nu_e\rangle = a|\nu_1\rangle + b|\nu_2\rangle + c|\nu_3\rangle, \quad \nu_1, \nu_2, \nu_3 : \text{mass eigenstates}$$

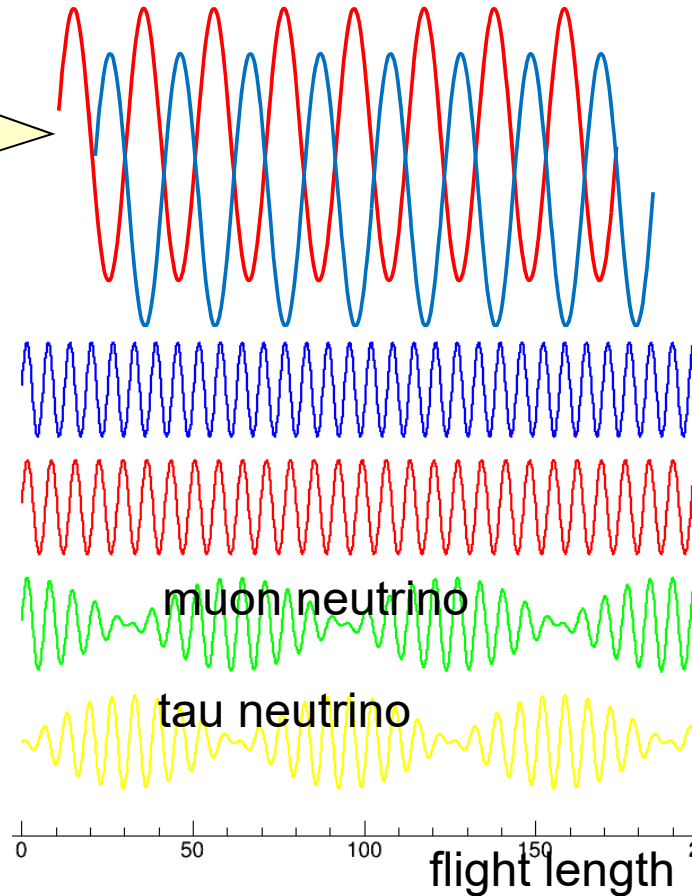
Same thing is happening for quarks. (e.g. Cabibbo angle)

$$\text{Partner of } |up\rangle \text{ quark} = a|down\rangle + b|strange\rangle + c|bottom\rangle$$

Then,
a neutrino is produced as an eigenstate of one flavor,
but propagate with two different speed



Traveling
distance L



$$|v_\alpha\rangle = |v_1\rangle \cos\theta + |v_2\rangle \sin\theta, \quad \alpha=e,\mu,\tau$$

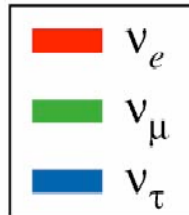
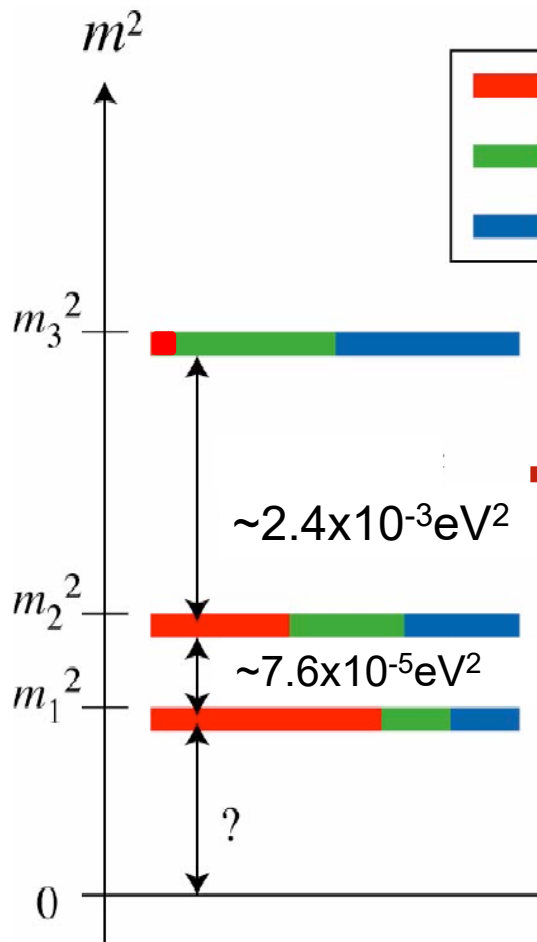
Neutrino Oscillation!

$$|v_1\rangle e^{-i\frac{m_1^2}{2E}L} \cos\theta + |v_2\rangle e^{-i\frac{m_2^2}{2E}L} \sin\theta \Rightarrow |v_\beta\rangle$$

$$P(\nu_\alpha \rightarrow \nu_\beta) = \left| \langle \nu_\beta | \nu_\alpha \rangle \right|^2 = \sin^2 2\theta \sin^2 \left(\Delta m^2 \frac{L}{4E} \right)$$

OSCILLATE ACROSS LOOOOONG DISTANCE

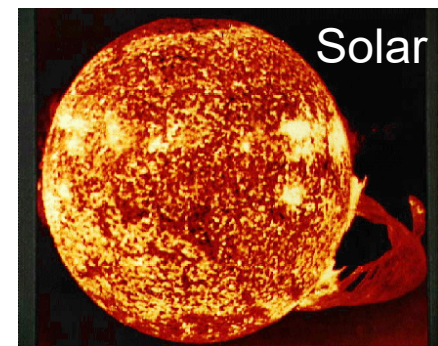
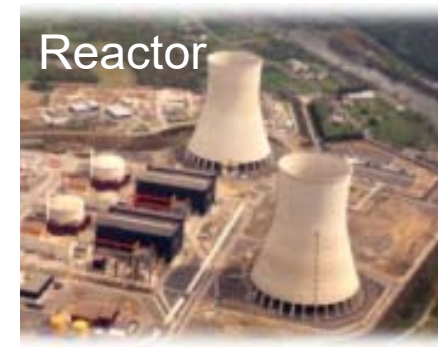
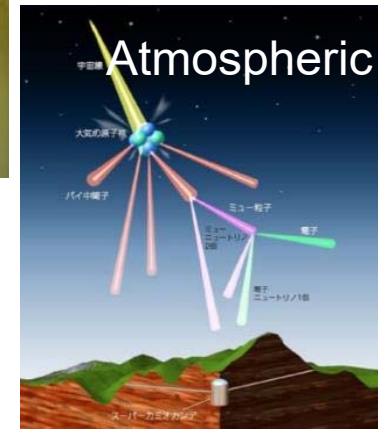
$$P(\nu_{\alpha} \rightarrow \nu_{\beta}) = \sin^2 2\theta \sin^2 \left(1.27 \Delta m^2 (eV^2) \frac{L(km)}{E(GeV)} \right)$$



~300 km @
 $E_{\nu}=700 \text{ MeV}$

~50 km @
 $E_{\nu}=3 \text{ MeV}$

However, in other words, nature allows oscillations corresponding to three Δm^2 's happen in Earth Scale



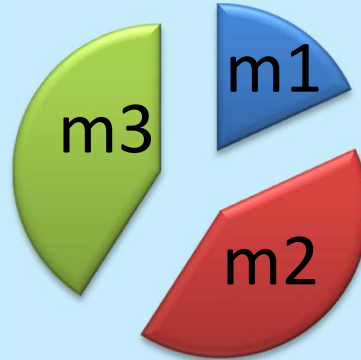
$m_3^2 \gg m_2^2 > m_1^2$ case
 $m_2^2 > m_1^2 > m_3^2$ could be also possible

AND WE FOUND,

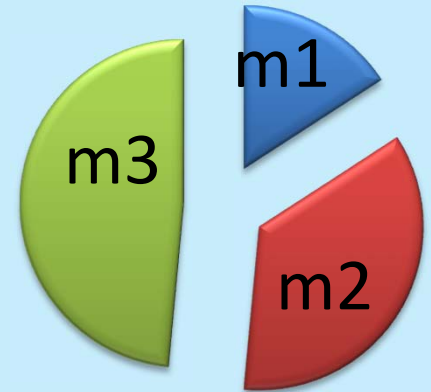
m3 electron neutrino



muon neutrino



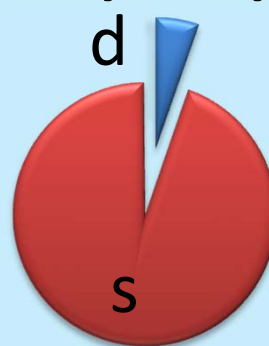
tau neutrino



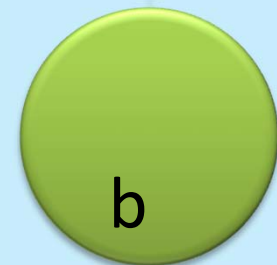
u quark partner



c quark partner

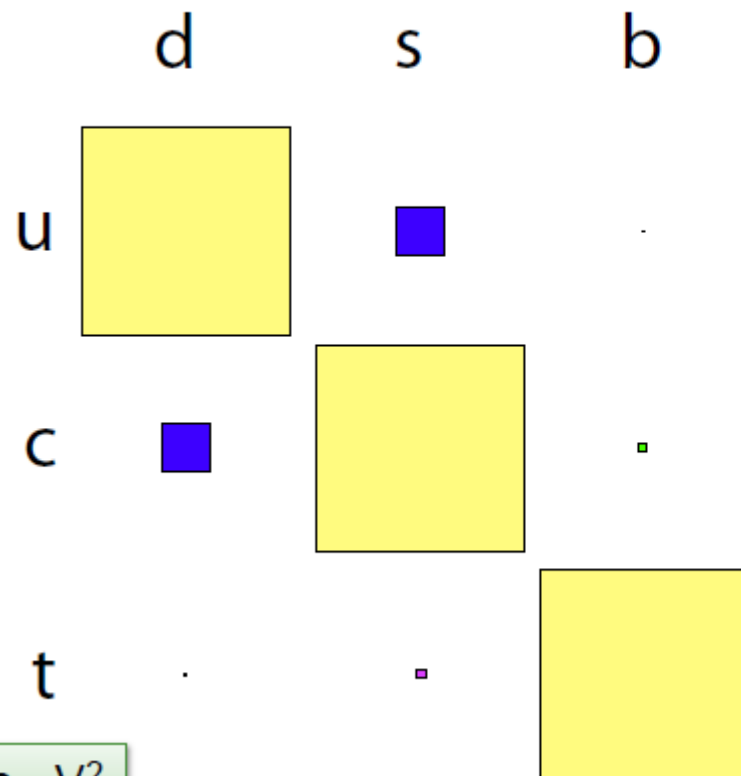


t quark partner



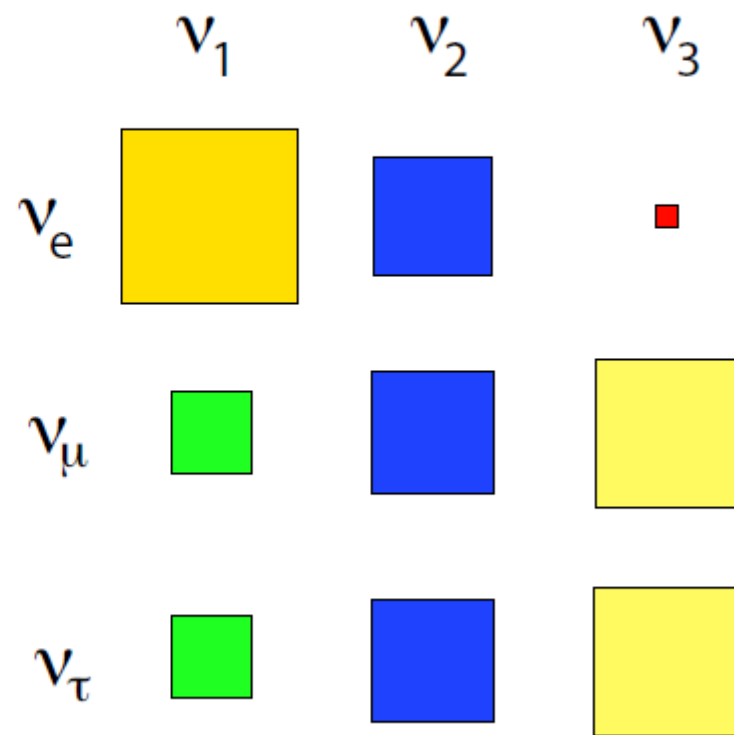
$$\begin{pmatrix} \text{flavor}_1 \\ \text{flavor}_2 \\ \text{flavor}_3 \end{pmatrix} = U_{3 \times 3} \begin{pmatrix} m_1 \\ m_2 \\ m_3 \end{pmatrix}$$

CKM



Area $\sim V^2$

PMNS



DO THEY MEAN SOMETHING?

$$\begin{pmatrix} \text{flavor}_1 \\ \text{flavor}_2 \\ \text{flavor}_3 \end{pmatrix} = U_{3 \times 3} \begin{pmatrix} m_1 \\ m_2 \\ m_3 \end{pmatrix}$$

quark

$$U_{CKM} \approx \begin{pmatrix} 0.97 & 0.23 & 0.004 \\ 0.23 & 1.01 & 0.04 \\ 0.008 & 0.04 & 0.89 \end{pmatrix}$$

lepton

$$U_{PMNS} \approx \begin{pmatrix} 0.82 & 0.55 & 0.16 \\ -0.49 & 0.52 & 0.55 \\ 0.20 & -0.65 & 0.70 \end{pmatrix}$$

Assuming some symmetry among quarks and leptons, some models predict

$$U_{CKM} \approx \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$$U_{MNS} = \begin{pmatrix} \sqrt{2/3} & \sqrt{1/3} & 0 \\ -\sqrt{1/6} & \sqrt{1/3} & \sqrt{1/2} \\ \sqrt{1/6} & -\sqrt{1/3} & \sqrt{1/2} \end{pmatrix} = \begin{pmatrix} 0.816 & 0.577 & 0 \\ -0.408 & 0.577 & 0.707 \\ 0.408 & -0.577 & 0.707 \end{pmatrix}$$

MIXING BTW. FLAVOR AND MASS

$$\begin{pmatrix} \text{flavor}_1 \\ \text{flavor}_2 \\ \text{flavor}_3 \end{pmatrix} = U_{3 \times 3} \begin{pmatrix} m_1 \\ m_2 \\ m_3 \end{pmatrix}$$

Kobayashi-Maskawa theory

Mixing with ≥ 3 types can have imaginary components.

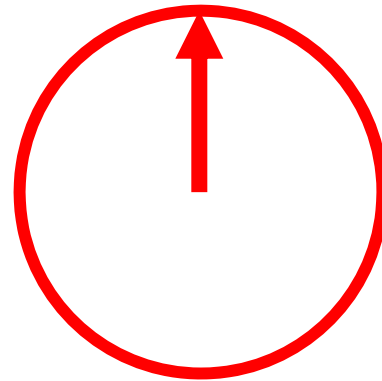
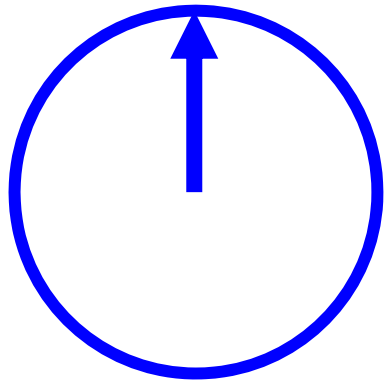
In case of 3 generations, **one CP phase**.

$$U_{\text{PMNS}} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & +c_{23} & +s_{23} \\ 0 & -s_{23} & +c_{23} \end{pmatrix} \begin{pmatrix} +c_{13} & 0 & +s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & +c_{13} \end{pmatrix} \begin{pmatrix} +c_{12} & +s_{12} & 0 \\ -s_{12} & +c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$(c_{ij} = \cos \theta_{ij}, s_{ij} = \sin \theta_{ij})$

ANTI-PARTICLE

- Reverse the rotation of the clock i.e. phase of wavefunction
→ particle behaves like oppositely-charged particle.



$$e^{-i\{(E+eV)t-(\vec{p}+e\vec{A})\cdot\vec{x}\}} \text{ vs. } e^{i\{(E-eV)t-(-\vec{p}-e\vec{A})\cdot\vec{x}\}}$$

MIXING BTW. FLAVOR AND MASS

$$\begin{pmatrix} \text{flavor}_1 \\ \text{flavor}_2 \\ \text{flavor}_3 \end{pmatrix} = U_{3 \times 3} \begin{pmatrix} m_1 \\ m_2 \\ m_3 \end{pmatrix}$$

Kobayashi-Maskawa theory

Mixing with > 3 generation can have imaginary components, **which introduce different behavior between particle and anti-particle=CP violation.**

In case of 3 generations, **one CP phase.**

CKM (quark sector) $\delta \sim 60^\circ$
PMNS (lepton sector) $\delta \sim ?$

$$U_{\text{PMNS}} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & +c_{23} & +s_{23} \\ 0 & -s_{23} & +c_{23} \end{pmatrix} \begin{pmatrix} +c_{13} & 0 & +s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & +c_{13} \end{pmatrix} \begin{pmatrix} +c_{12} & +s_{12} & 0 \\ -s_{12} & +c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$(c_{ij} = \cos \theta_{ij}, s_{ij} = \sin \theta_{ij})$

LEPTONIC CPV CAN BE MUCH LARGER THAN QUARK'S

$\delta_{CP}^{CKM} \sim 60^\circ \sim 70^\circ$ looks large, but cannot explain matter-dominant universe.

δ_{CP} is dependent on definition.

Jarlskog Invariant : independent of definition. show the size of CP violation effect.

$$J_{CP} \equiv \text{Im}(U_{\mu 3} U_{e 3}^* U_{e 2} U_{\mu 2}^*) = \frac{1}{8} \sin 2\theta_{12} \sin 2\theta_{23} \sin 2\theta_{13} \cos \theta_{13} \sin \delta_{CP}$$
$$J_{CP}^{CKM} \approx 3 \times 10^{-5}$$

$$J_{CP}^{PMNS} \approx 0.03 \sin \delta_{CP}$$

PDG2015 “NEUTRINOMASS, MIXING, AND OSCILLATIONS”

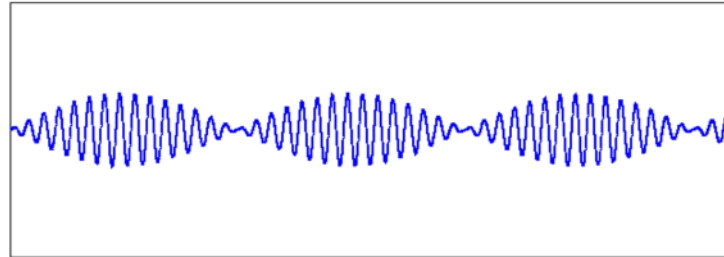
A value of $|\sin \theta_{13} \sin \delta| \gtrsim 0.09$, and thus $\sin \theta_{13} \gtrsim 0.09$, is a necessary condition for a successful “flavoured” leptogenesis with hierarchical heavy Majorana neutrinos when the CP violation required for the generation of the matter-antimatter asymmetry of the Universe is provided entirely by the Dirac CP violating phase in the neutrino mixing matrix [191]. This condition is comfortably compatible both with the measured value of $\sin^2 \theta_{13}$ and with the best fit value of $\delta \cong 3\pi/2$.

$$|\sin \theta_{13} \sin \delta| \geq 0.09 \rightarrow |\sin \delta| \geq 0.58$$

If we find that $|\sin \delta| \geq 0.58$, we can, at least, say that there exists CPV which is as large as to produce matter-dominant universe.

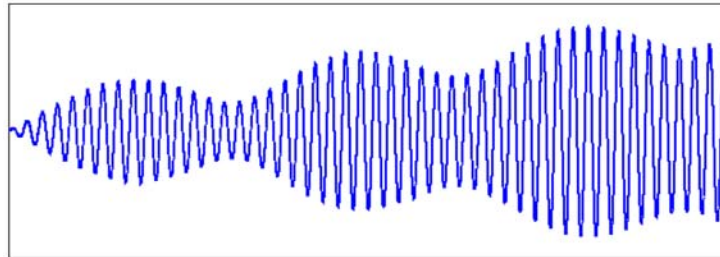
NEUTRINO OSCILLATION OF THREE STATES

Two neutrinos case

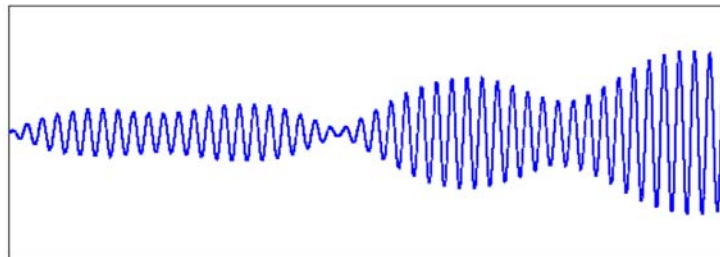


Three neutrinos case

neutrino



antineutrino



Depending on the value of δ_{CP} , neutrino and antineutrino oscillate differently!

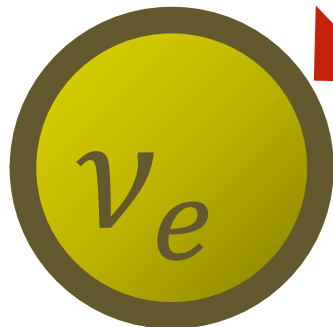
OSCILLATIONS PECULIAR TO THE ACC.-BASED LONG BASELINE EXPERIMENT

CP violation is
accessible only via
appearance, and can be
large for ν_e appearance.

ν_μ disappearance $\sim \sin^2 2\theta_{23} \sim 100\%$ at right energy

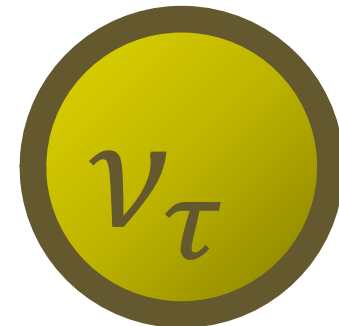
ν_e
 $\sim \sin^2 \theta_{23} \sin^2 2\theta_{13} \sim 5\%$

ν_τ
 $\sim \cos^4 \theta_{13} \sin^2 2\theta_{23} \sim 95\%$



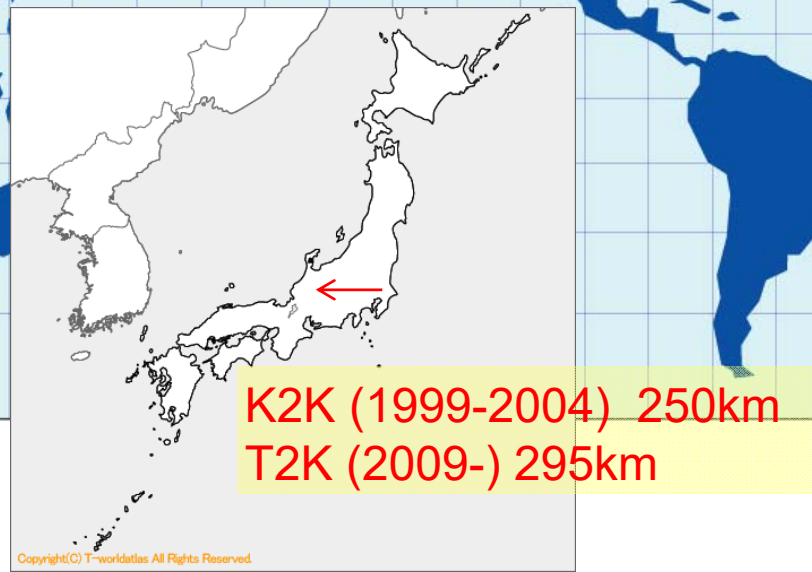
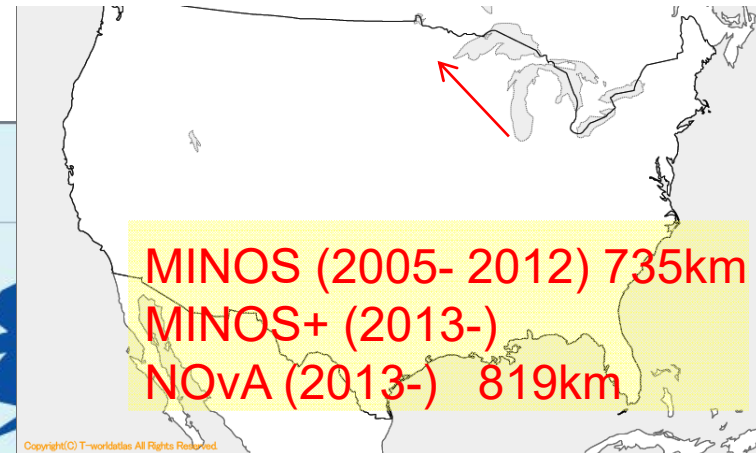
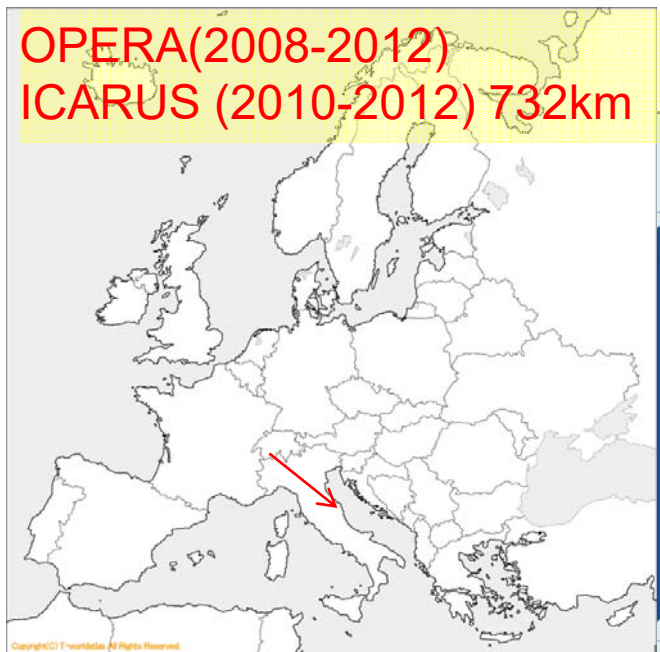
ν

$\bar{\nu}$

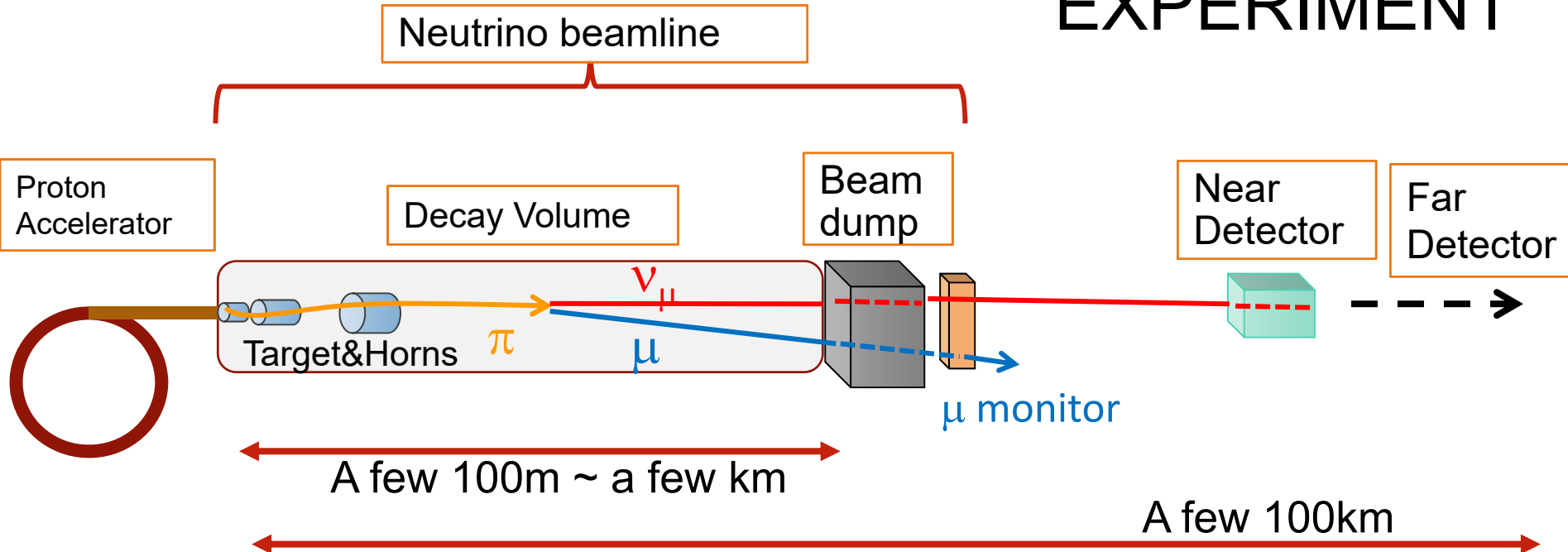


Interference term
 $\sim \sin \delta_{CP}$ for neutrino
 $\sim -\sin \delta_{CP}$ for antineutrino
 $\pm 27\%$ effect on ν_e appearance

World Acc.-based Long baseline ν oscillation experiments



COMPONENTS OF THE LONG BASELINE NEUTRINO EXPERIMENT

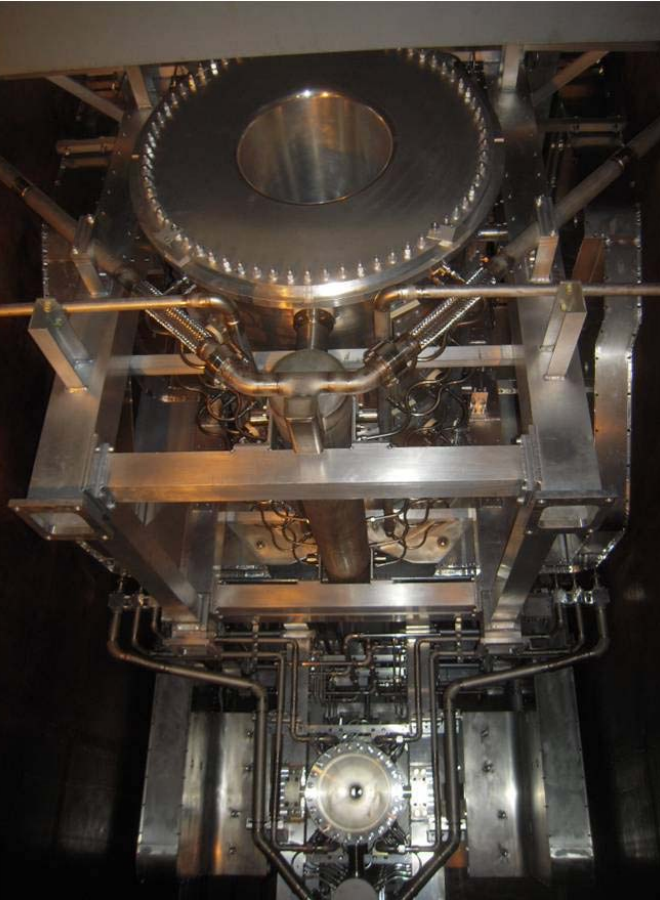


Toroidal magnetic field by 'horn' focuses
 $\pi^+ \rightarrow \nu_\mu$ beam

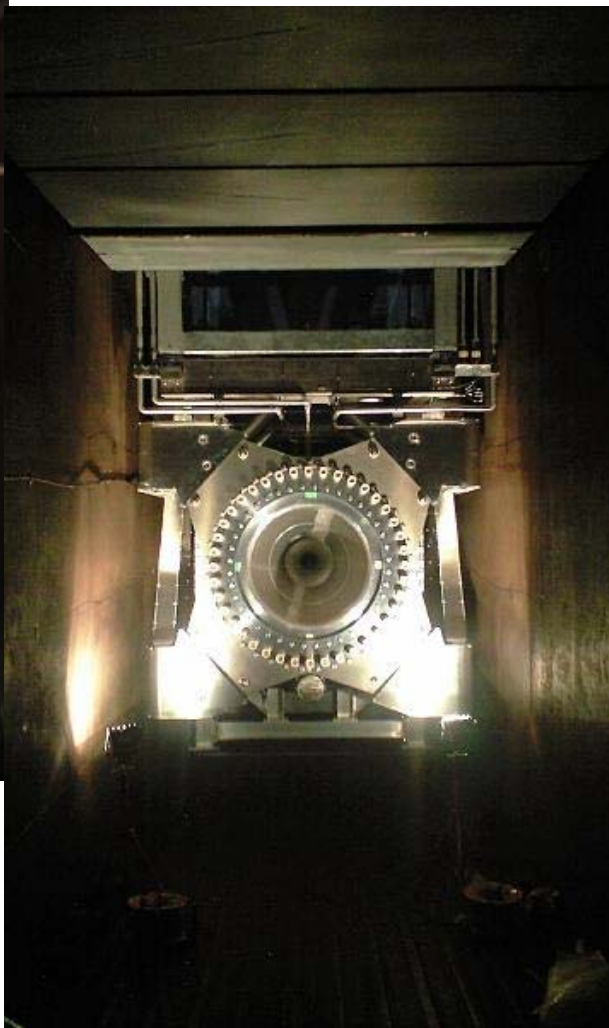
or

$\pi^- \rightarrow \bar{\nu}_\mu$ beam

timing synchronized
using GPS



1st Horn & 2nd Horn



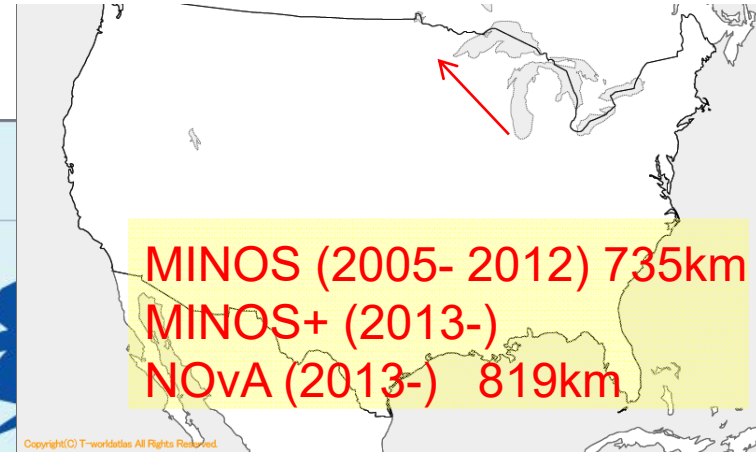
3rd Horn

Decay Volume

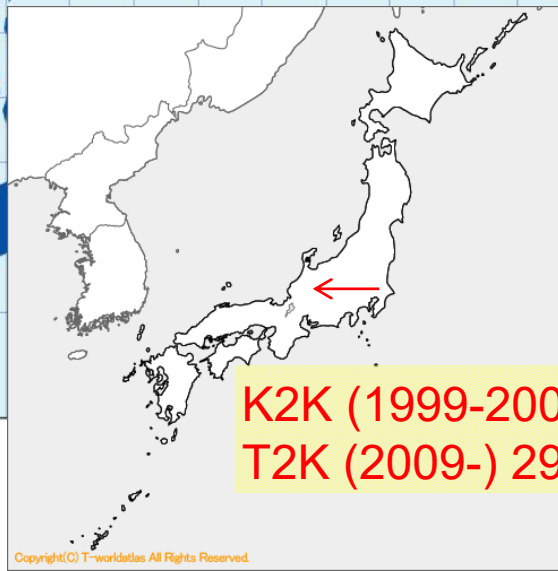


World Acc.-based Long baseline ν oscillation experiments

OPERA(2008-2012)
ICARUS (2010-2012) 732km



K2K (1999-2004) 250km
T2K (2009-) 295km

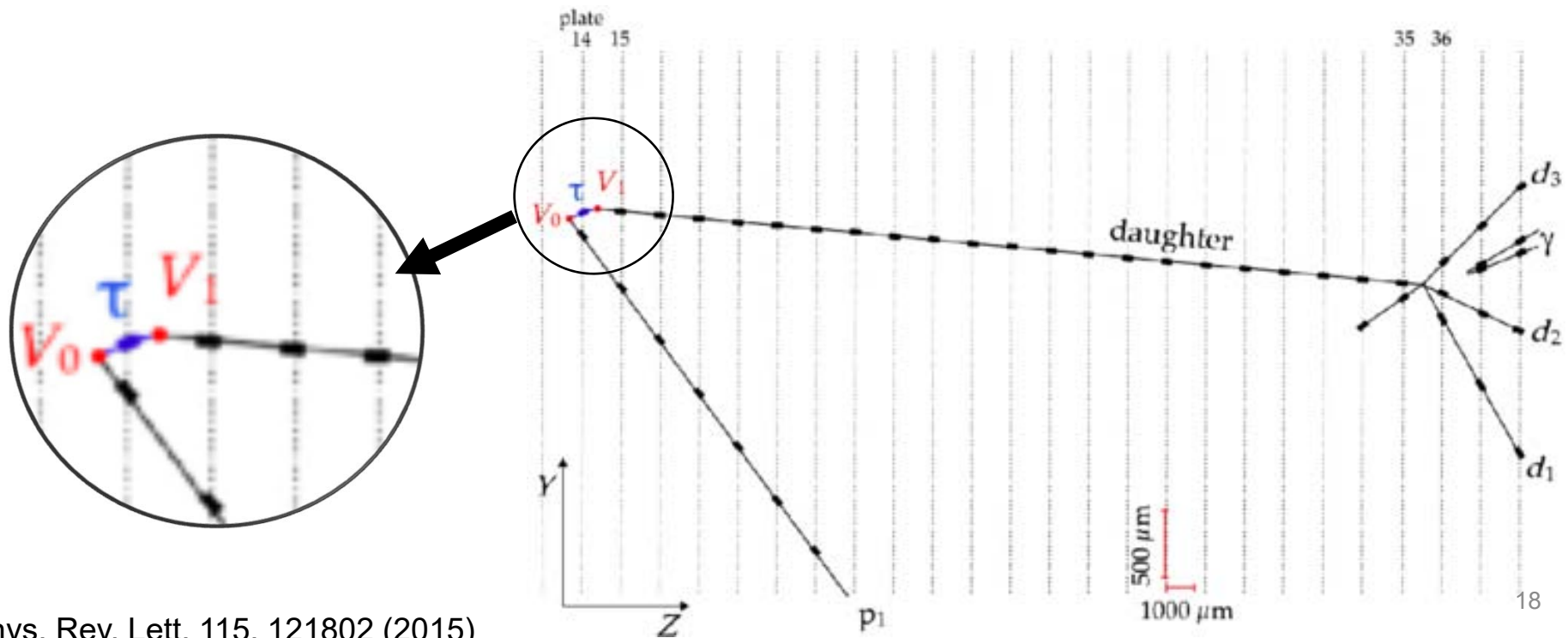


ν_τ APPEARANCE

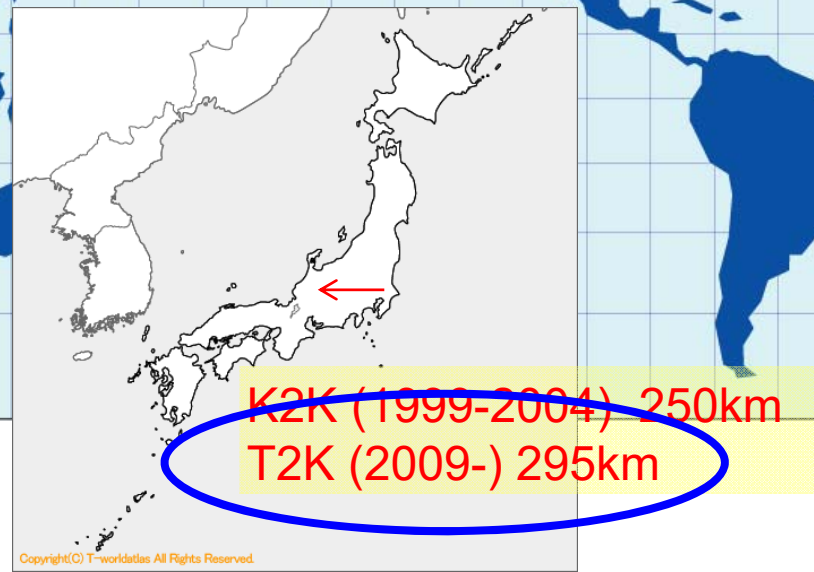
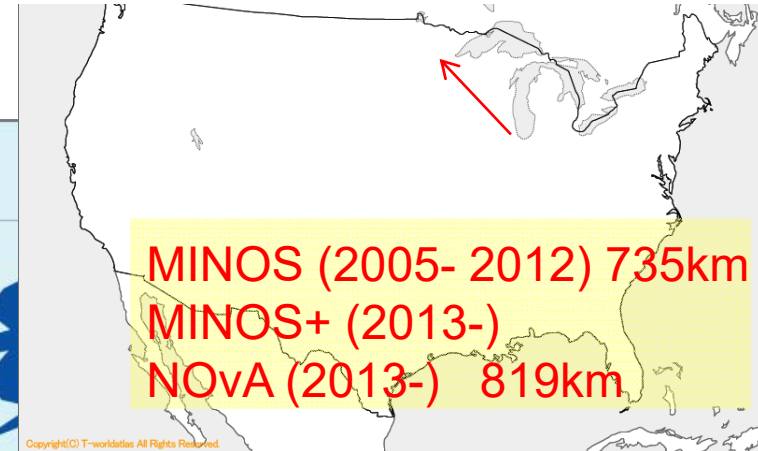
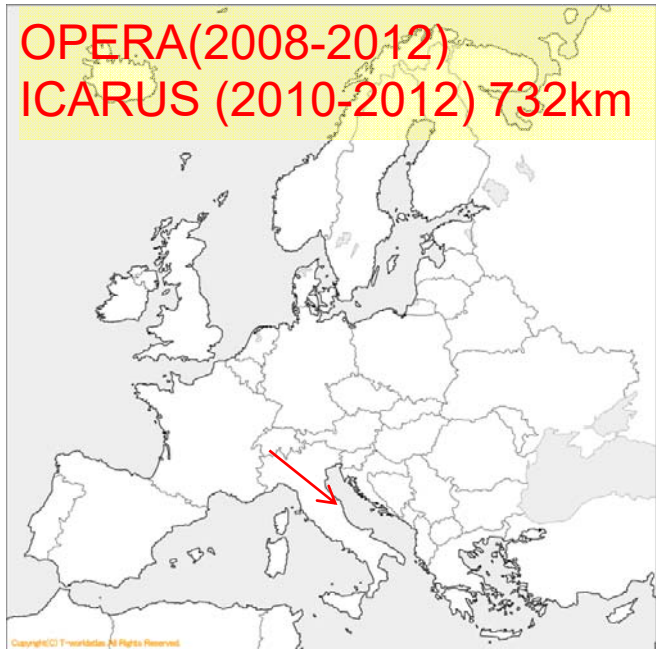
3.8σ ν_τ appearance by Super-K atmospheric data (*Abe et al., PRL 110, 181802 (2013)*) from a sample of enhanced τ -like events.

OPERA experiment, long baseline accelerator neutrino experiments. identifies τ production in event-by-event basis using nuclear emulsion.

5 events observed \rightarrow confirmation by 5.1σ significance.



World Acc.-based Long baseline ν oscillation experiments



ν_e appearance

2011 T2K observed 6 events(1.5 bkgs). 2.5σ significance.

2013 T2K observed 28 events over 4.9bkgs 7.3σ significance.

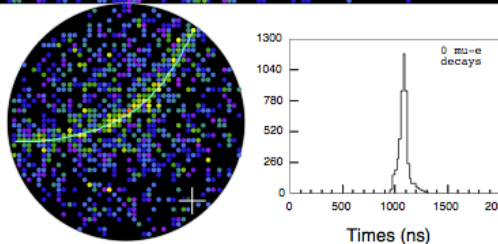
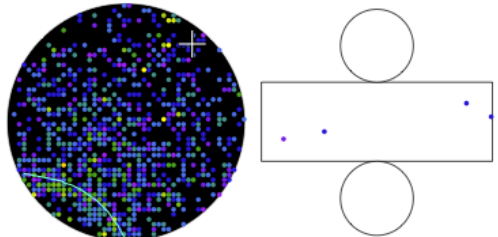
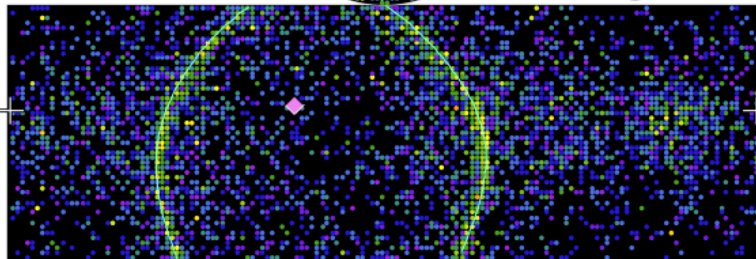
Also first confirmation of ‘appearance’ w/ $>5\sigma$ significance.

Super-Kamiokande IV

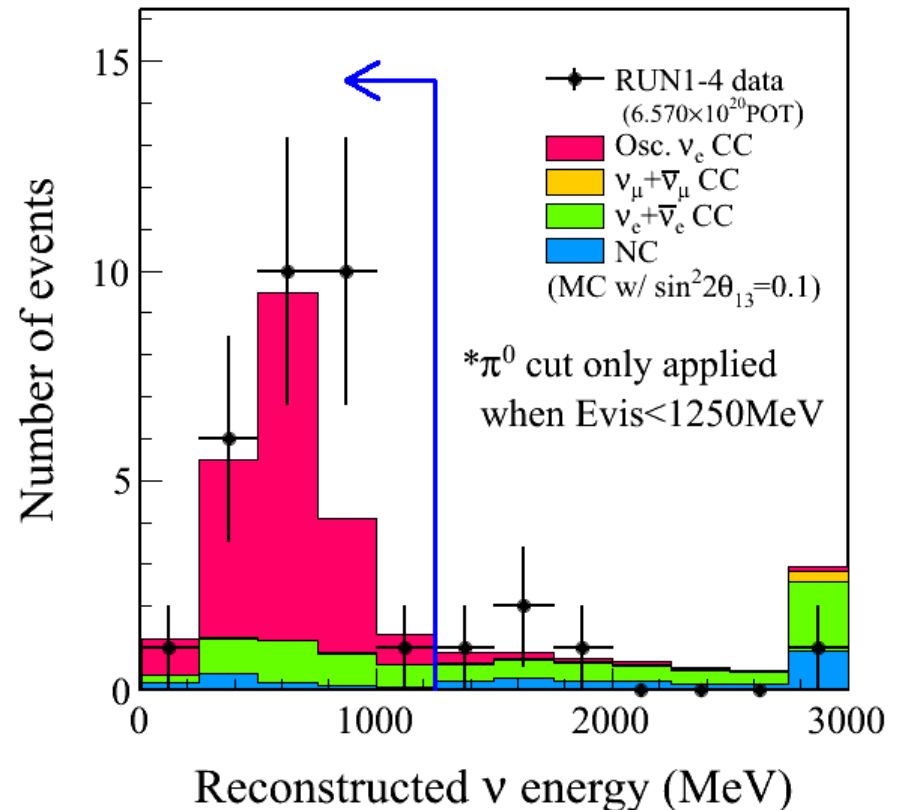
T2K Beam Run 0 Spill 1039222
Run 67969 Sub 921 Event 218931934
10-12-22:14:15:18
T2K beam dt = 1782.6 ns
Inner: 4804 hits, 9970 pe
Outer: 4 hits, 3 pe
Trigger: 0x80000007
D_wall: 244.2 cm
e-like, p = 1049.0 MeV/c

Charge (pe)

- * >26.7
- * 23.3-26.7
- * 20.2-23.3
- * 17.3-20.2
- * 14.7-17.3
- * 12.2-14.7
- * 10.0-12.2
- * 8.0-10.0
- * 6.2- 8.0
- * 4.7- 6.2
- * 3.3- 4.7
- * 2.2- 3.3
- * 1.3- 2.2
- * 0.7- 1.3
- * 0.2- 0.7
- * < 0.2



Example of ν_e candidate event



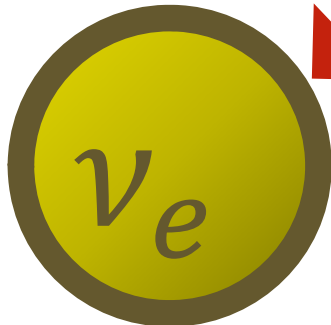
OSCILLATIONS PECULIAR TO THE LONG BASELINE EXPERIMENT



ν_μ disappearance $\sim \propto \sin^2 2\theta_{23} \sim 100\%$ at right energy

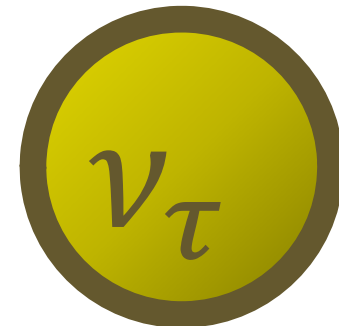
Since $\theta_{12}, \theta_{13}, \theta_{23}$ are known from other measurements,

$\nu_e \sim \propto \sin^2 \theta_{23}$ δ_{CP} can be determined $\nu_\mu \rightarrow \nu_e$ measurement, $\sim 95\%$
even w/o $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ meas.



ν

$\bar{\nu}$



Interference term

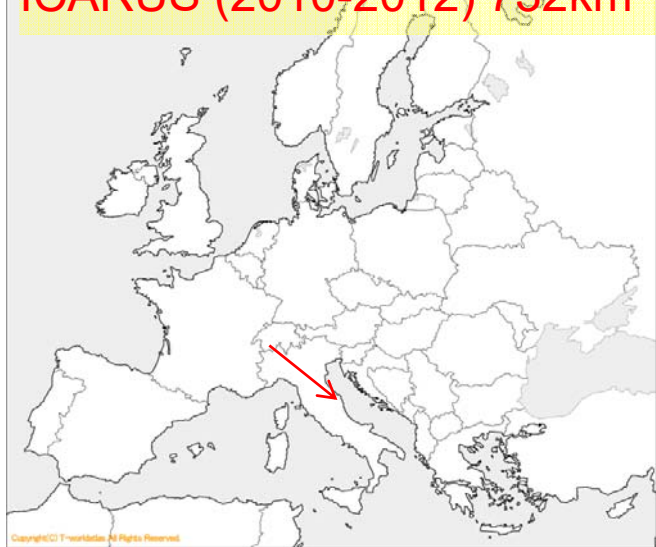
$\sim \propto \sin \delta_{CP}$ for neutrino

$\sim \propto -\sin \delta_{CP}$ for antineutrino

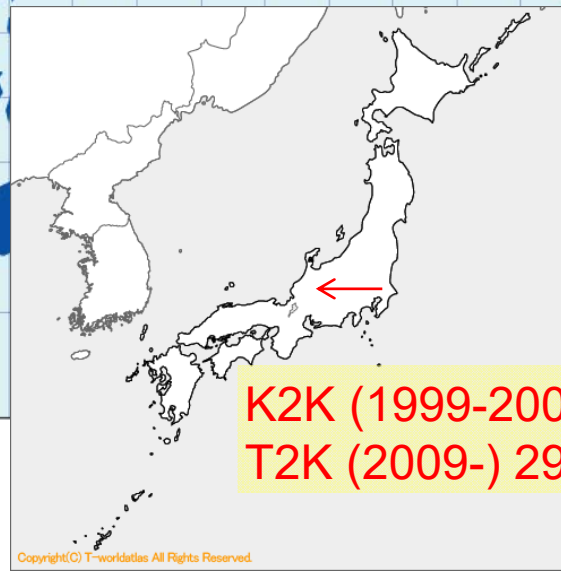
$\pm 27\%$ effect on ν_e appearance

World Acc.-based Long baseline ν oscillation experiments

OPERA(2008-2012)
ICARUS (2010-2012) 732km



MINOS (2005- 2012) 735km
MINOS+ (2013-)
NOvA (2013-) 819km



K2K (1999-2004) 250km
T2K (2009-) 295km

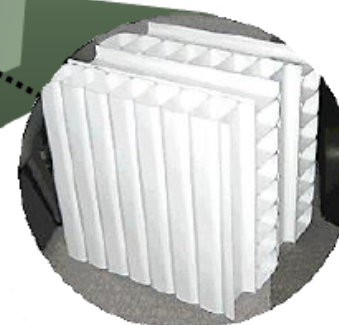
NO ν A detectors

Extruded PVC cells filled with
11M liters of scintillator
instrumented with
 λ -shifting fiber and APDs

15.6 m

Far Detector
14 kton
896 layers

Detector

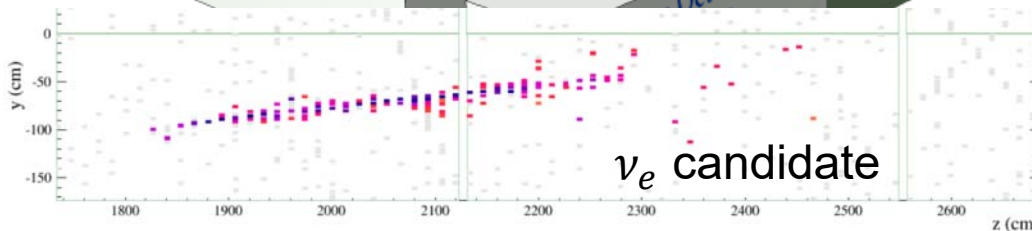


A NO ν A cell

To APD

1560 cm

4 cm \times 6 cm



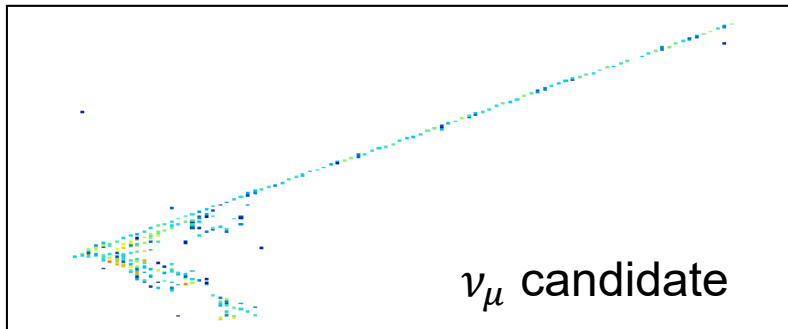
ν_e candidate

Far detector:

14-kton, fine-grained,
low-Z, highly-active
tracking calorimeter
 \rightarrow 344,000 channels

Near detector:

0.3-kton version of
the same
 \rightarrow 18,000 channels



ν_μ candidate

DEEP LEARNING!

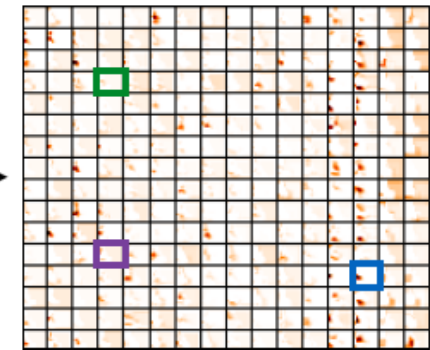
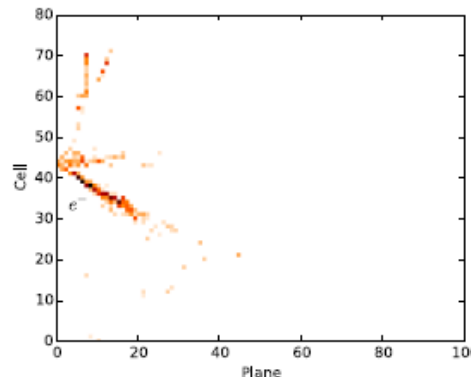
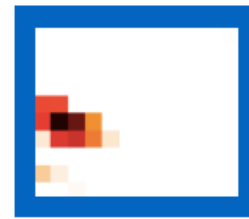
Improved Event Selection

9



P. Vahle, Neutrino 2016

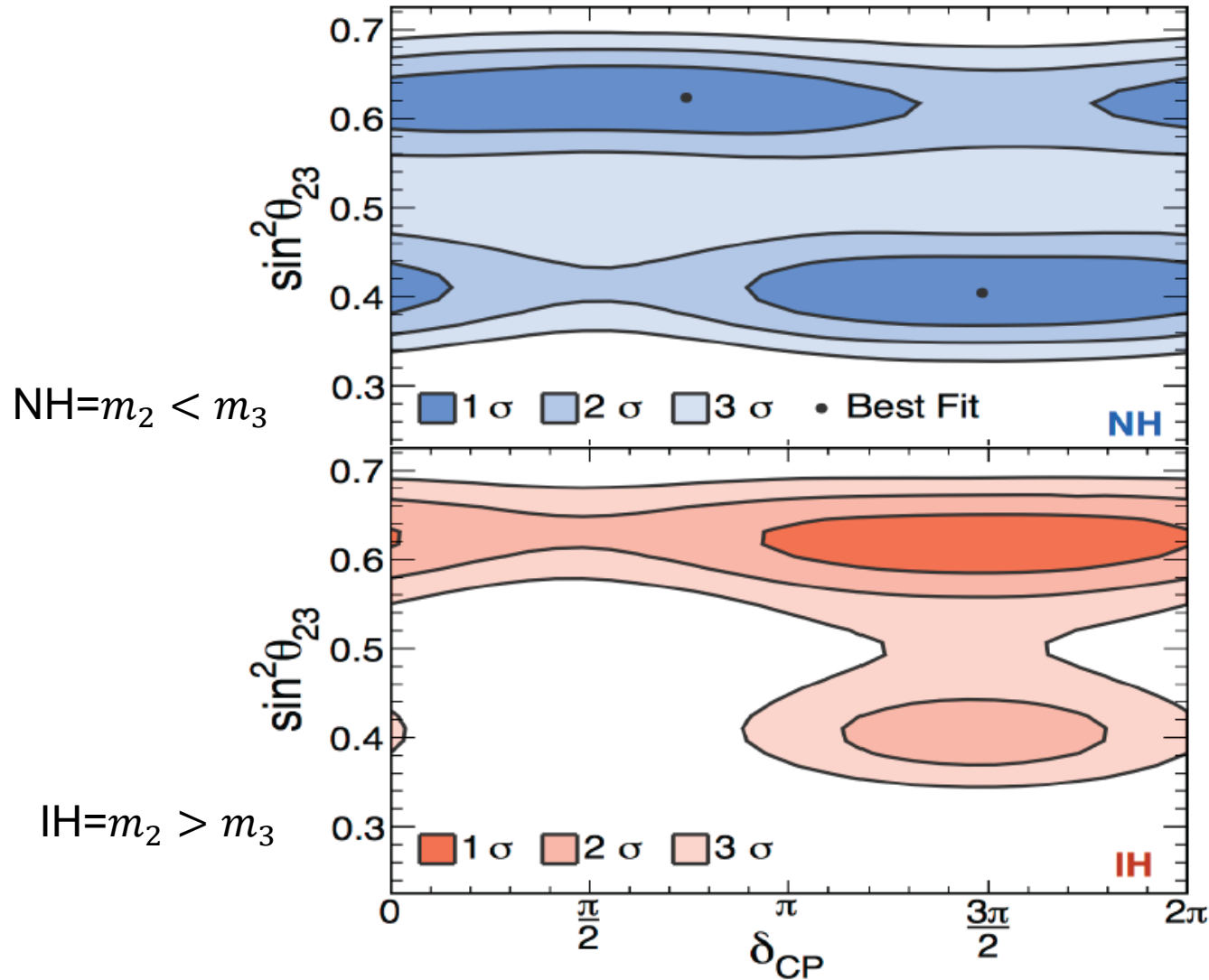
- This analysis features a new event selection technique based on ideas from computer vision and deep learning
- Calibrated hit maps are inputs to Convolutional Visual Network (CVN)
- Series of image processing transformations applied to extract abstract features
- Extracted features used as inputs to a conventional neural network to classify the event



A. Aurisano et al., arXiv:1604.01444
Posters P1.028 by A. Radovic, P1.032 by
F. Psihas and A. Himmel for more detail

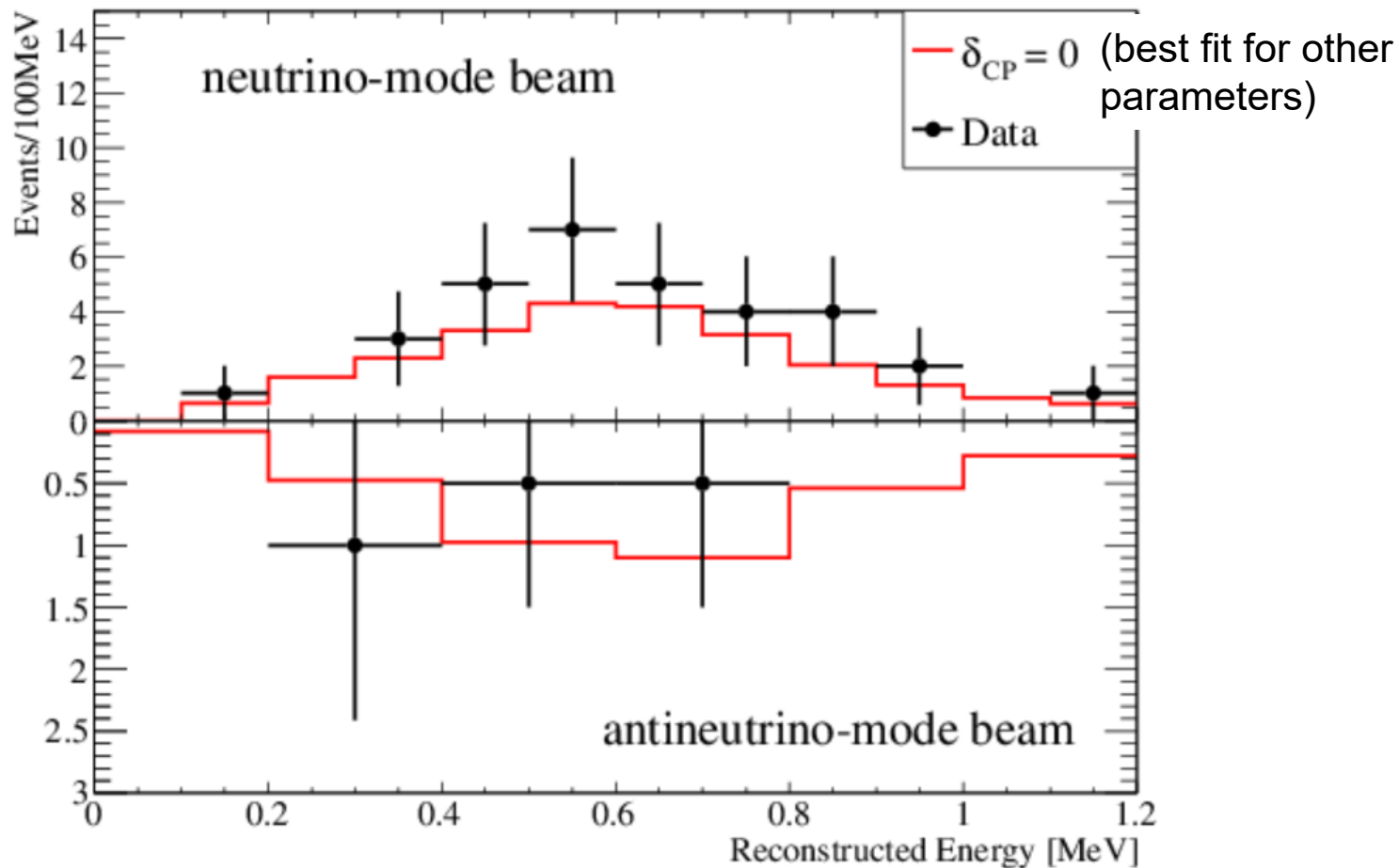
NOvA

Based on neutrino data by 2016

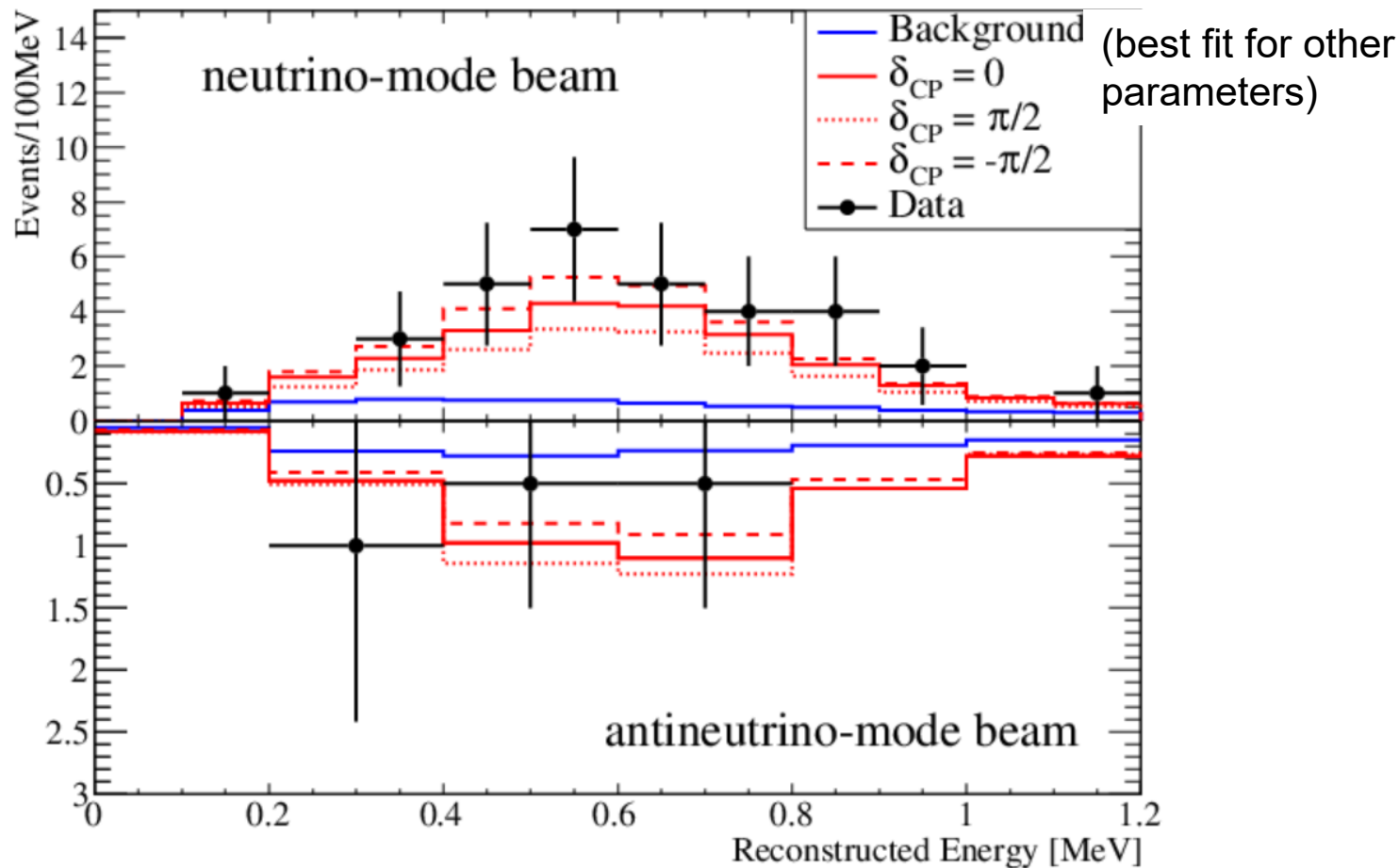


Now accumulating
antineutrino data!

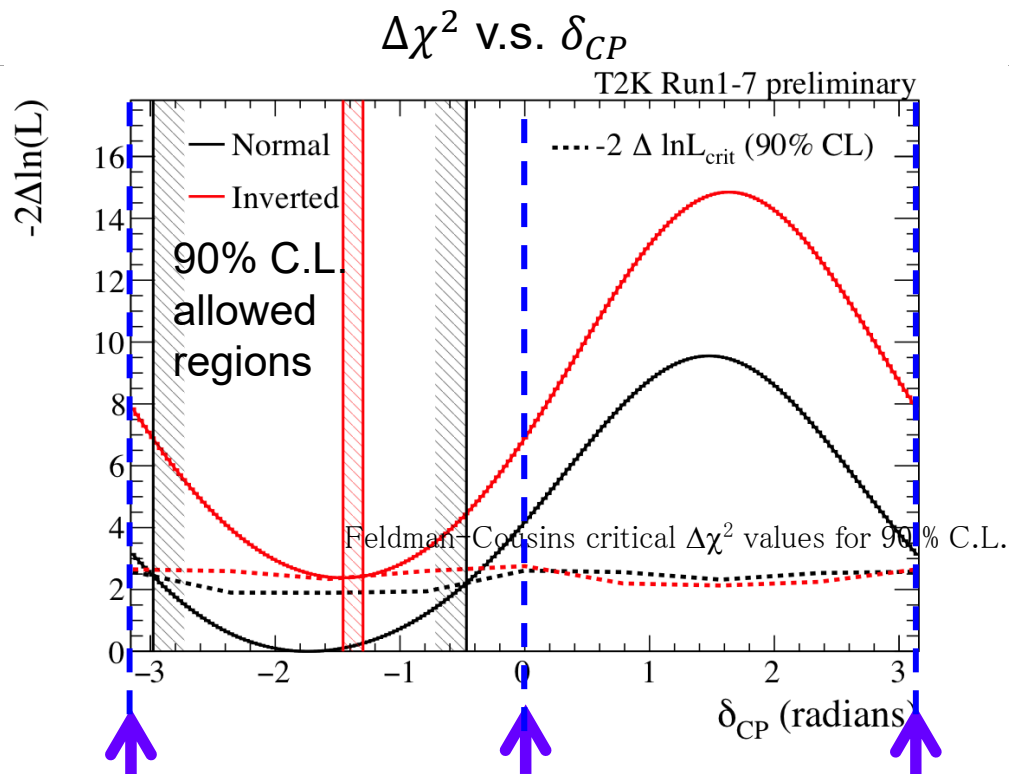
T2K HAS STARTED ANTINEUTRINO DATA TAKING IN 2013



T2K ν_e AND $\bar{\nu}_e$ SELECTED EVENT DISTRIBUTIONS



INDICATION OF CP VIOLATION???



no CPV

90% Confidence Interval:

Normal mass ordering	$(-171^\circ, -27^\circ)$
Inverted mass ordering	$(-84^\circ, -73^\circ)$

CP conserving case is disfavored by $>90\%$ C.L.

Need more statistics!
This year, we doubled neutrino-beam data. Results coming soon.

PROSPECT IN 10 YEARS

- **T2K**

- beam power 350 kW → 900 kW by ~2021
- CP violation may be indicated with >90% C.L. if it is maximally violated.

- **NOvA**

- Mass ordering determination ($\sim 3\sigma$)
- CP violation may be indicated with >90% C.L. if it is maximally violated.

- **T2K-II**

- Target Beam power 1.3 MW, near detector upgrade
- x 3 times T2K POT in total by ~2026
- CP violation can be observed with 3σ C.L. if it is maximally violated
- (new collaborators are welcomed)

T2K-II Expected number of events (1:1 ν : $\bar{\nu}$ running case)

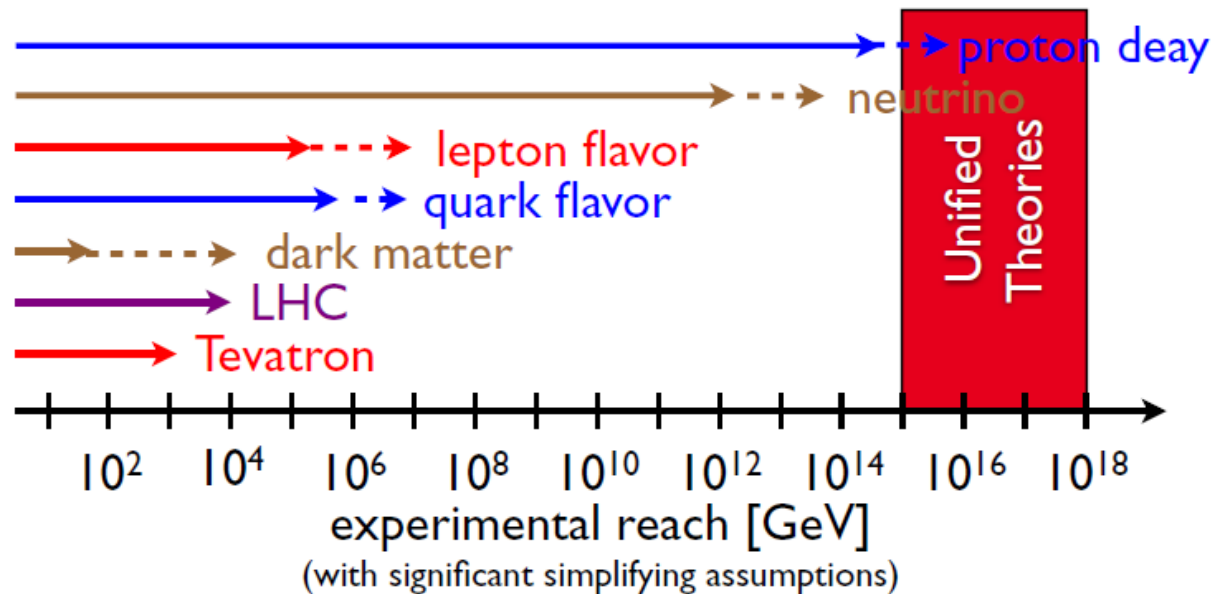
ν_e sample : 455 evts $\pm 20\%$ change depending on δ_{CP}

$\bar{\nu}_e$ sample : 129 evts $\pm 13\%$ change depending on δ_{CP}

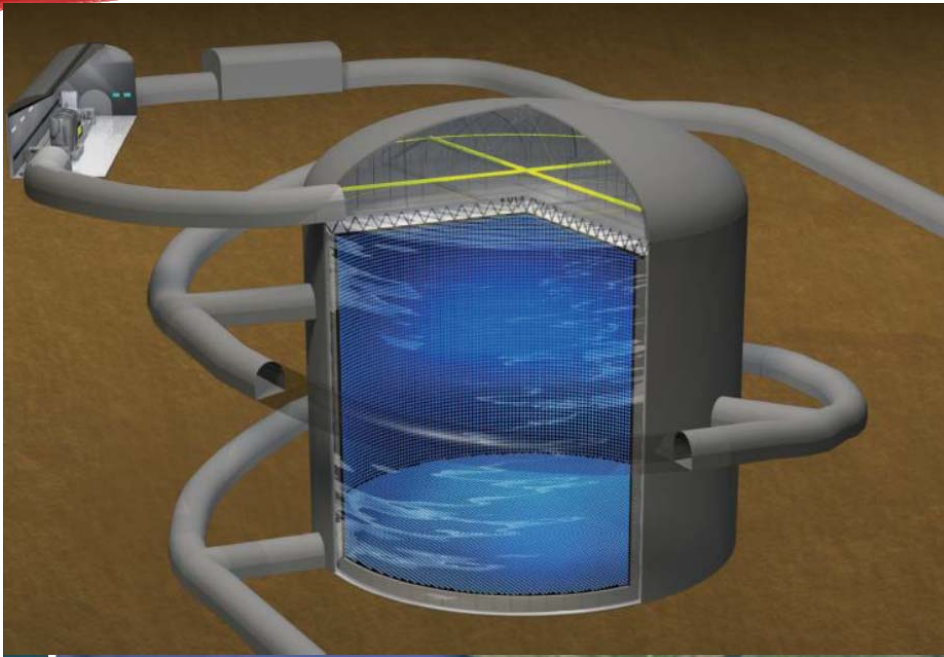
SURPRISE?

$K/B/\mu/\tau/n$ and LHC is searching TeV scale new physics. difficult with neutrinos because of low statistics.

However, extremely-small neutrino mass is possibly generated by Grand-unification Theory(GUT) scale physics ($m_\nu \sim \frac{m_D^2}{M_N}$ (m_D Dirac mass like quarks/charged leptons, $M_N \lesssim \text{GUT scale}$)). So m_ν scale measurement(=neutrino oscillation) may reveal surprising phenomenon from GUT scale physics



BEYOND 10 YEARS,



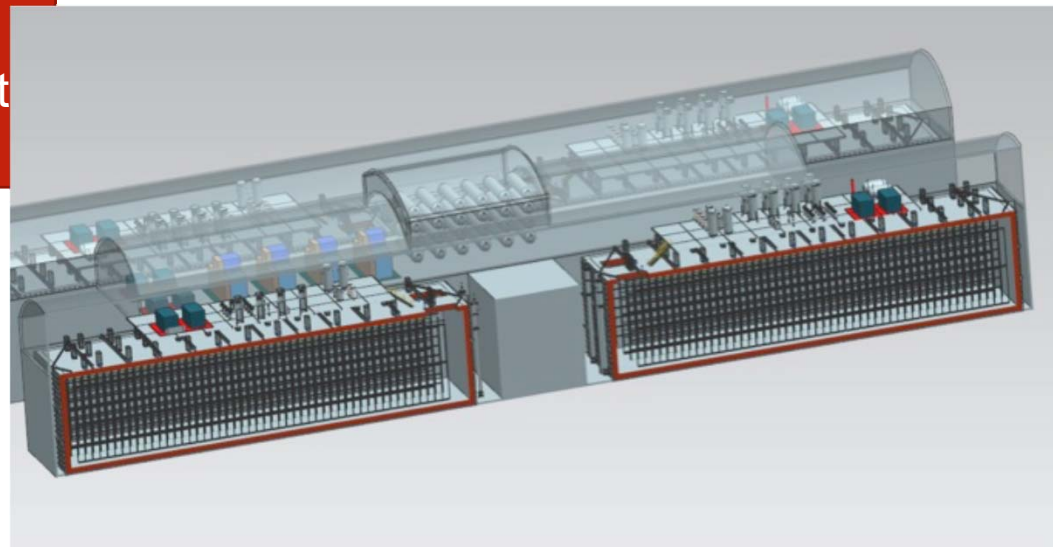
Hyper-Kamiokande
~1 Mton Water Cherenkov Detector at
Kamioka,

CPV for wide range of δ_{CP}
 ν 's from Solar, Supernova,
atmospheric
Proton decay, dark matter...

Liquid Ar technology
established by ICARUS



In US, DUNE experiment,
with 40kt liquid Ar TPC



SUMMARY

- Since the discovery of neutrino oscillation, understanding of 3 generation mixing has been steadily progressed.
 - Still we don't know why mass-flavor mixing is small in quark and large in lepton.
- Remaining known unknowns are: CP violation, mass ordering and maximal or non-maximal θ_{23}
 - They will be solved in 1~20 years by the accelerator long baseline experiments, reactor long baseline experiments and atmospheric measurements
- Our knowledge on the origin of the matter-dominant universe will be progressed by neutrino CP violation.
- It is likely that the tiny mass of neutrino originates from the physics in high energy ($>10^{14}$ GeV) scales. Then, unknown unknown may be also found in the neutrino oscillations
- We are entering the era where precise comparison among measurements are possible like the quark sector.