

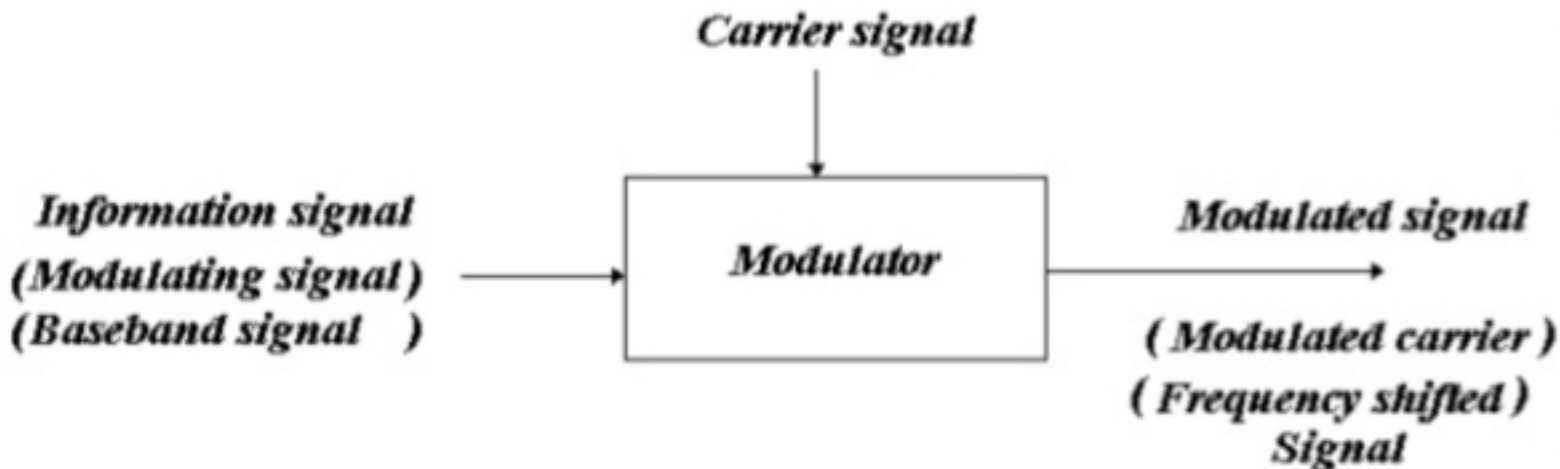
*A primer
on
a **Real** Gravitational Wave Detector*

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Basic concepts

- *Laudatio (encomium)* of the Modulation Method
- Formalism to treat optical elements
- Michelson Interferometer
- F-P cavity
- Light detection method
- Interferometer Control
- Sensitivity
- Conclusion → the Virgo detector

- Modulation is the process by which some characteristic of the carrier is varied in accordance with a modulation wave
- It permits to put information onto a high frequency carrier for transmission (*frequency translation*)
- In the modulation process, the modulation wave is at lower frequency, “the baseband signal”, while the signal transporting the information , “*the carrier*”, is at higher frequency

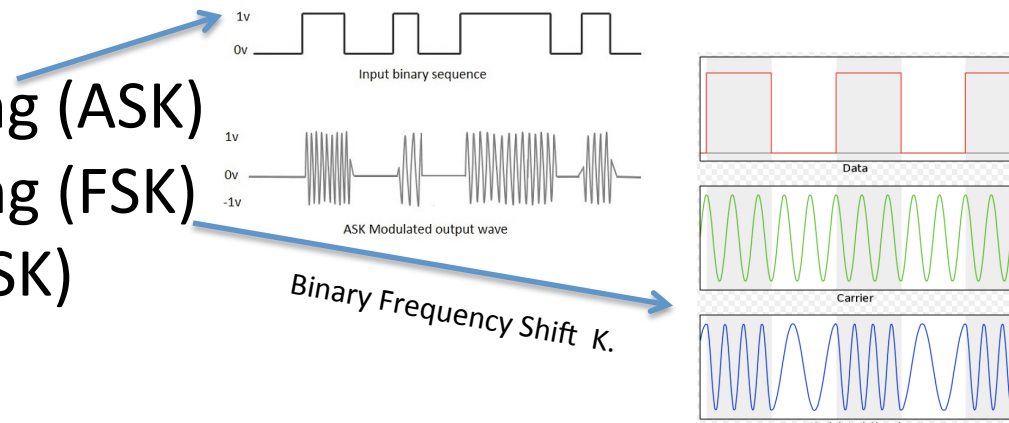


ANALOG Modulation

- Modulating signal and Carrier are both analog
- TYPES
 - Amplitude Modulation (AM)
 - Frequency Modulation (FM)
 - Phase modulation (PM)

=====

- DIGITAL demodulation
 - Modulating signal is digital, while the carrier is analog
- TYPES
 - Amplitude Shift Keying (ASK)
 - Frequency Shift Keying (FSK)
 - Phase Shift Keying (PSK)

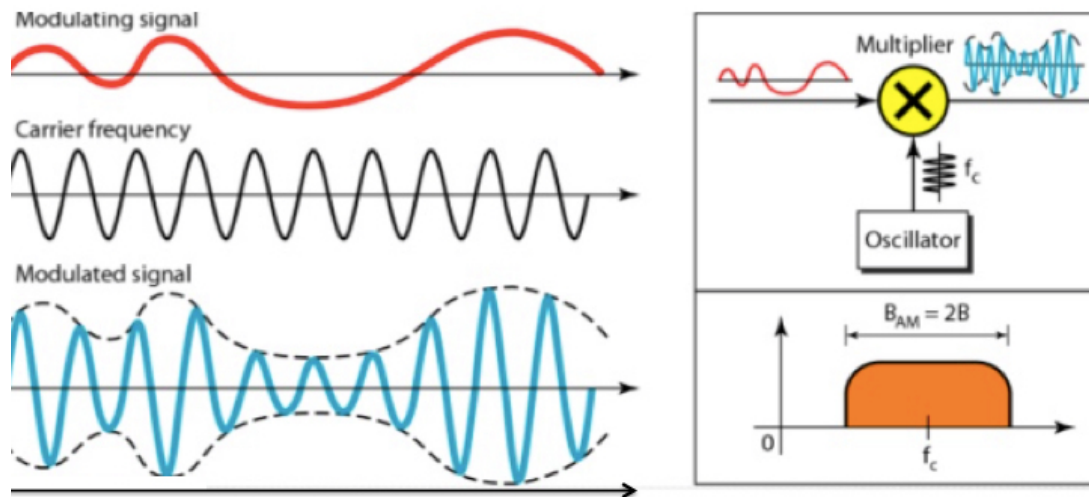


The simplest and earliest form of modulation

$$E(t) = A_c [1 + K m(t)] \cos(\omega_o t + \vartheta)$$

Non-linear process

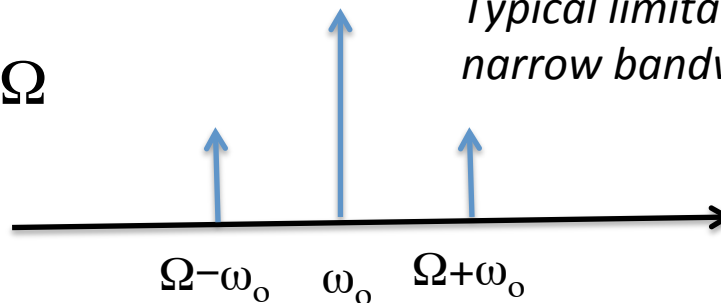
$$m(t) = A_m \cos(\Omega t + \phi)$$



Spectrum

Bandwidth $\Rightarrow \Delta\omega = 2\Omega$

*Typical limitation for sound transmission
narrow bandwidth*



Frequency (FM) and Phase (PM) Modulation

- FM the carrier's frequency vary with the signal's input voltage
- PM the carrier's phase vary with the signal's input voltage

Two types of angle modulation schemes they differ from each other in terms of their respective transmission bandwidth

The simplest single frequency phase modulation:

$$E(t) = \mathcal{E}_0 \cos[\omega_0 t + \delta_m \cos(\Omega t + \phi_m)] \quad \text{with } \Omega \ll \omega_0 \text{ and } \delta_m \ll 1$$

Complex representation of the modulated amplitude of the carrier

$$\mathcal{E}_0 (1/\sqrt{2}) \exp [i \delta_m \cos(\Omega t + \phi_m)]$$

Quadrature,

$$\mathcal{E}_c(t) \approx \mathcal{E}_0$$

$$\mathcal{E}_s(t) \approx \delta_m \mathcal{E}_0 \cos(\Omega t + \phi_m)$$

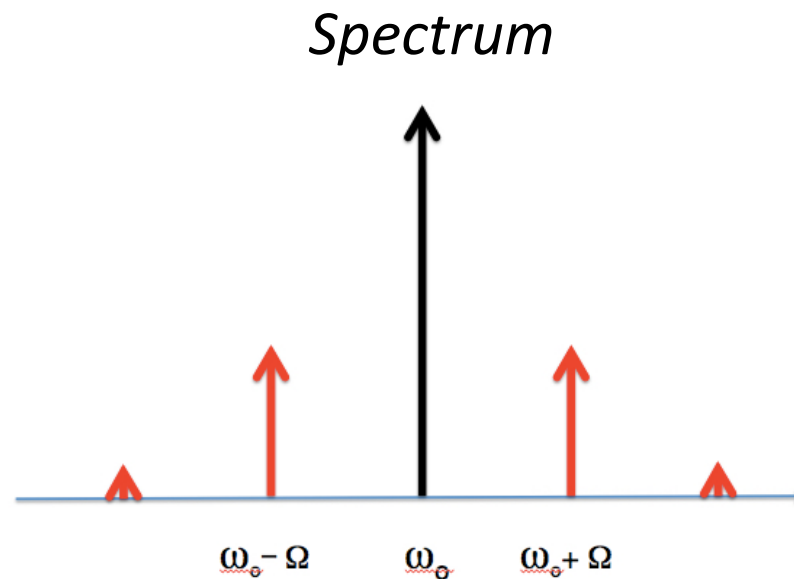
Note: only the sine quadrature to the respect of the carrier contains the modulation signal

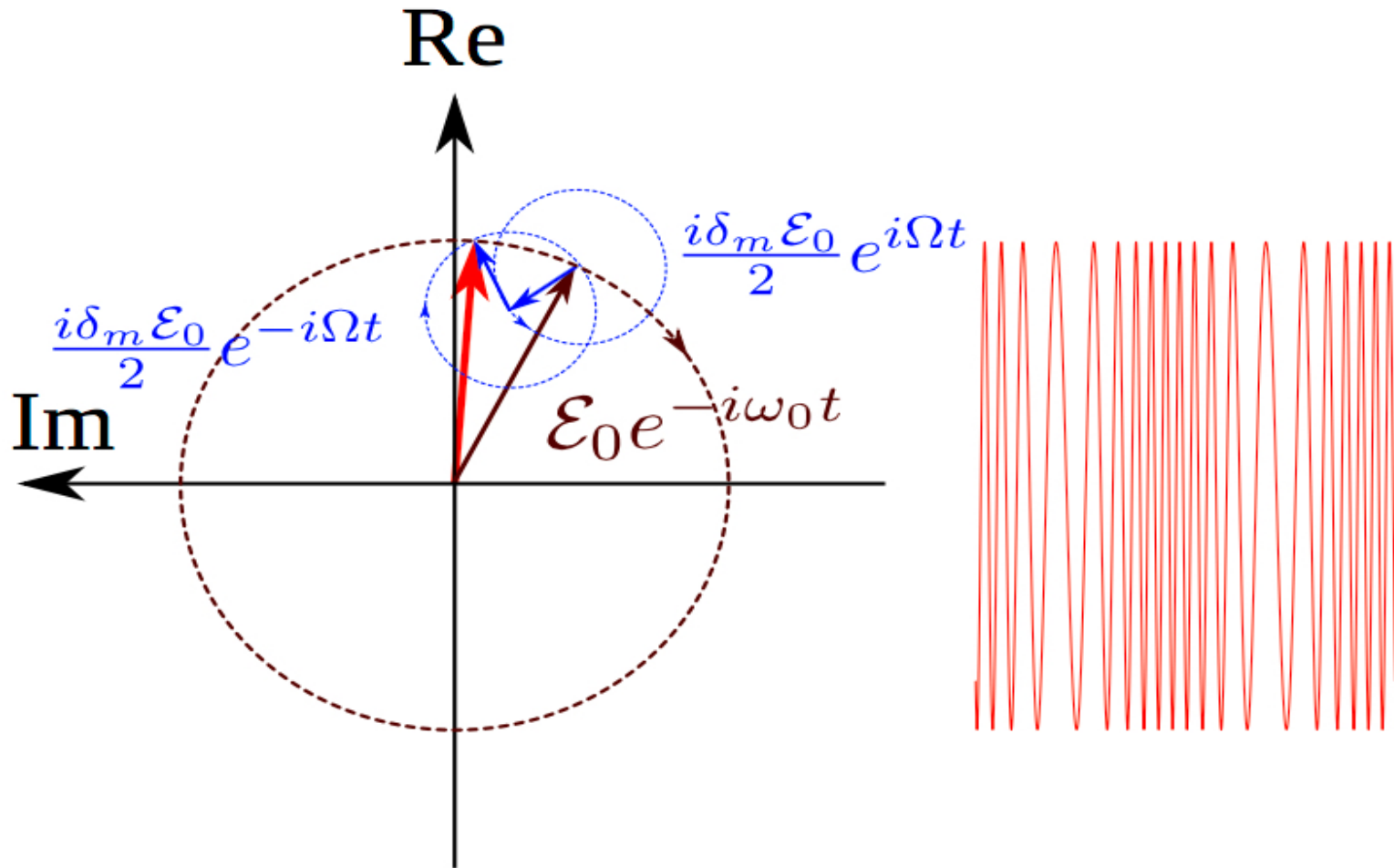
$$\mathcal{E}_0 (1/\sqrt{2}) \exp [i \delta_m \cos(\Omega t + \phi_m)] \simeq$$

$$\simeq J_0(\delta_m) + i(J_1(\delta_m) e^{i(\Omega t + \phi_m)} + J_{-1}(\delta_m) e^{-i(\Omega t + \phi_m)}) + \dots$$

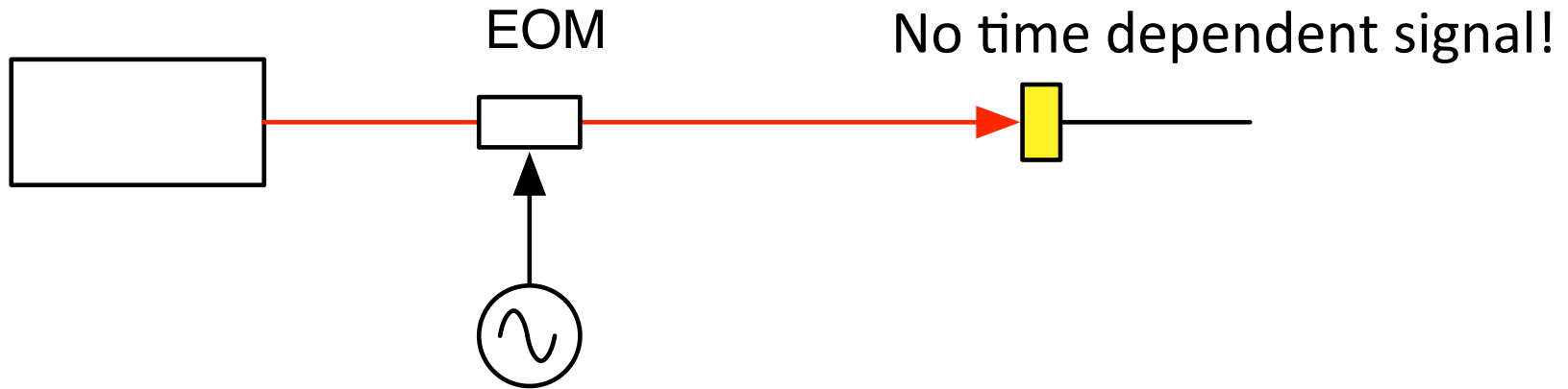
with

$$J_0(\delta_m) \simeq 1 \quad J_1(\delta_m) = J_{-1}(\delta_m) \simeq \delta_m/2$$





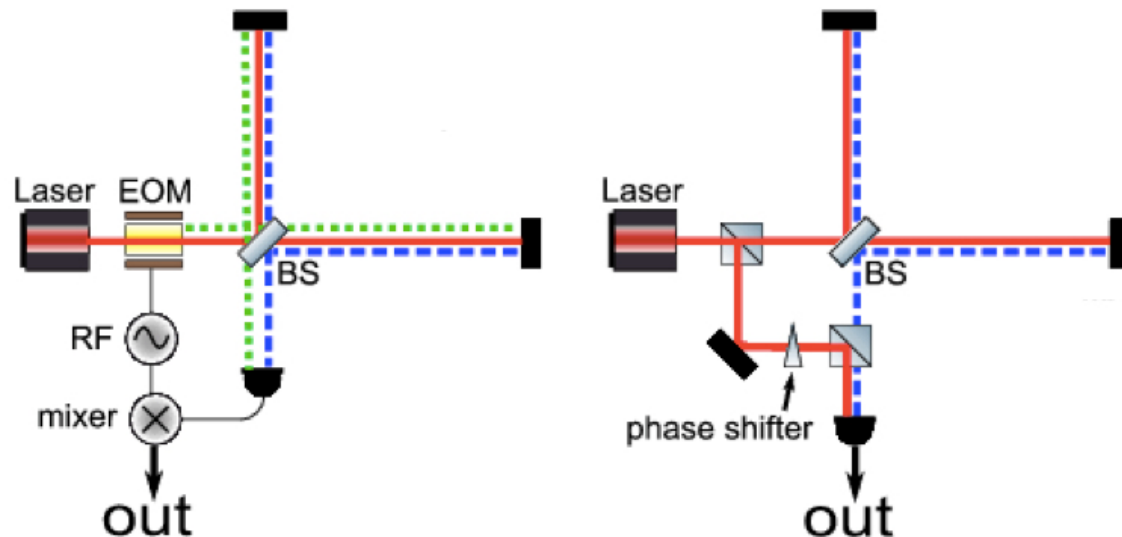
The resulting modulated oscillation vector (red arrow) has approximately the same length as the carrier field vector but outruns or lags behind the latter periodically with the modulation frequency Ω .



$$\begin{aligned}
 EE^* &= E_{\text{in}}^2 e^{i(-\Omega t + m \cos \omega_m t)} e^{-i(-\Omega t + m \cos \omega_m t)} \\
 &= E_{\text{in}}^2 \quad (\text{constant})
 \end{aligned}$$

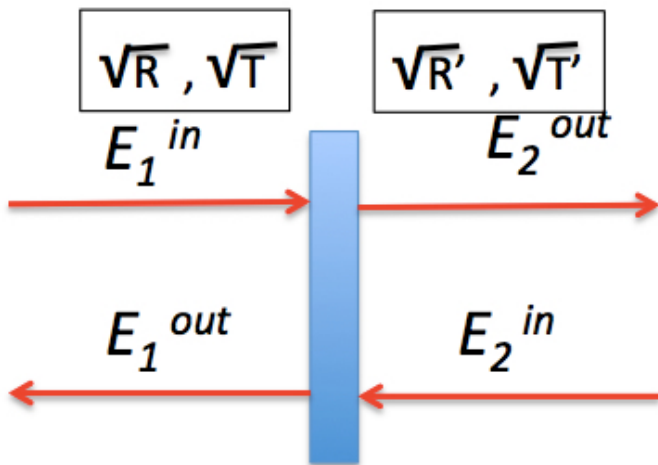
Note: Photodetectors don't feel phase modulation!

- PM detection must start before the photodetector: interfere the phase-modulated signal (*target beam*) with another signal (the *local oscillator* acting as a phase reference), prior to photodetection.
- *Local oscillator and Target beam* combine like an amplitude modulation to produce a detectable oscillation
 - **Homodyne detection:** local oscillator and target beam have the same frequencies
 - **Heterodyne detection:** local oscillator and target beam have different frequencies (frequency-shifted local oscillator beam, which beats with each individual sideband in the target beam with a different beat frequency)



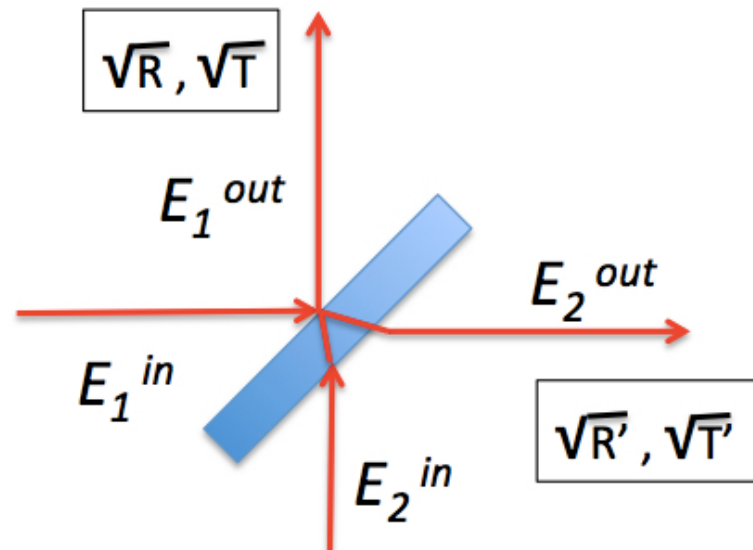
- Mirror and Beam Splitter \rightarrow linear systems with two input and two output ports

$r^2 = R \rightarrow$ Reflectivity $t^2 = T \rightarrow$ Transmittivity *in absence of loss*
 $R + T = 1$



$$E_1^{out} = -\sqrt{R} E_1^{in} + \sqrt{T} E_2^{in}$$

$$E_2^{out} = \sqrt{R'} E_2^{in} + \sqrt{T'} E_1^{in}$$



$$E_1^{out} = \sqrt{T'} E_2^{in} - \sqrt{R} E_2^{in}$$

$$E_2^{out} = \sqrt{R'} E_2^{in} + \sqrt{T} E_1^{in}$$

Mirror: Input – output electric fields connected via the Mirror Matrix M

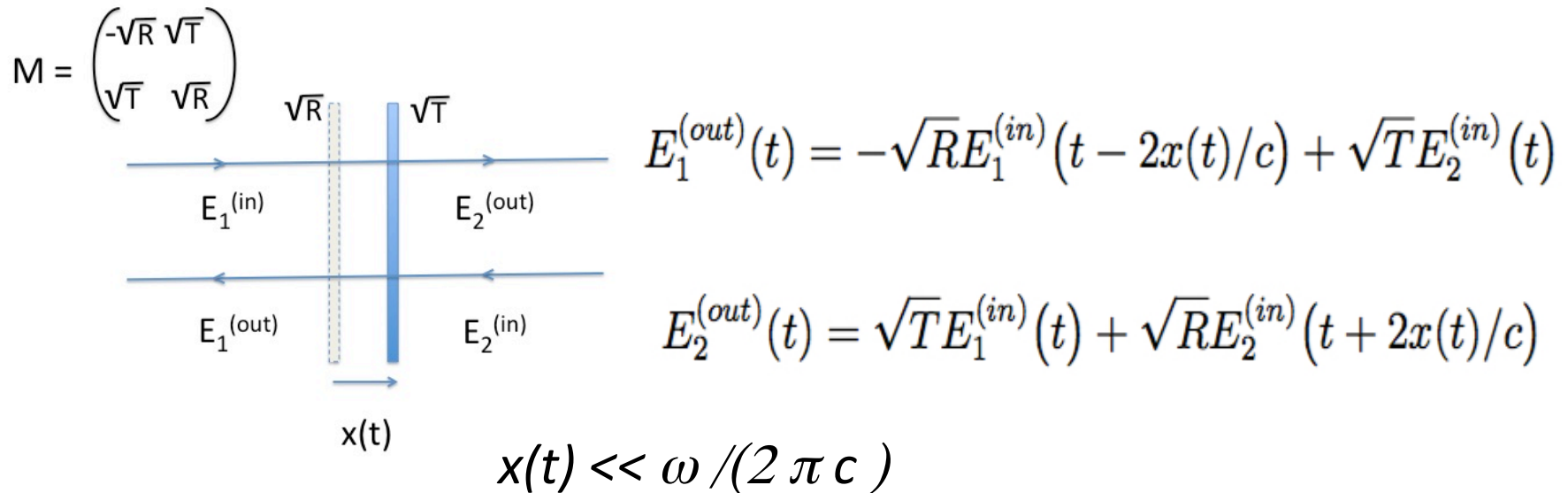
$$M = \begin{pmatrix} -\sqrt{R} & \sqrt{T} \\ \sqrt{T} & \sqrt{R} \end{pmatrix}$$

Mirror : electric field
quadrature
components

$$\begin{pmatrix} \mathcal{E}_{1c}^{out} \\ \mathcal{E}_{1s}^{out} \\ \mathcal{E}_{2c}^{out} \\ \mathcal{E}_{2s}^{out} \end{pmatrix} = \begin{pmatrix} -\sqrt{R} & 0 & \sqrt{T} & 0 \\ 0 & -\sqrt{R} & 0 & \sqrt{T} \\ \sqrt{T} & 0 & \sqrt{R} & 0 \\ 0 & \sqrt{T} & 0 & \sqrt{R} \end{pmatrix} \begin{pmatrix} \mathcal{E}_{1c}^{in} \\ \mathcal{E}_{1s}^{in} \\ \mathcal{E}_{2c}^{in} \\ \mathcal{E}_{2s}^{in} \end{pmatrix}$$

50% Beam Splitter :
matrix for transforming
electric field quadrature
components

$$M_{50\%,50\%} = \begin{pmatrix} -\sqrt{\frac{1}{2}} & 0 & \sqrt{\frac{1}{2}} & 0 \\ 0 & -\sqrt{\frac{1}{2}} & 0 & \sqrt{\frac{1}{2}} \\ \sqrt{\frac{1}{2}} & 0 & \sqrt{\frac{1}{2}} & 0 \\ 0 & \sqrt{\frac{1}{2}} & 0 & \sqrt{\frac{1}{2}} \end{pmatrix}$$



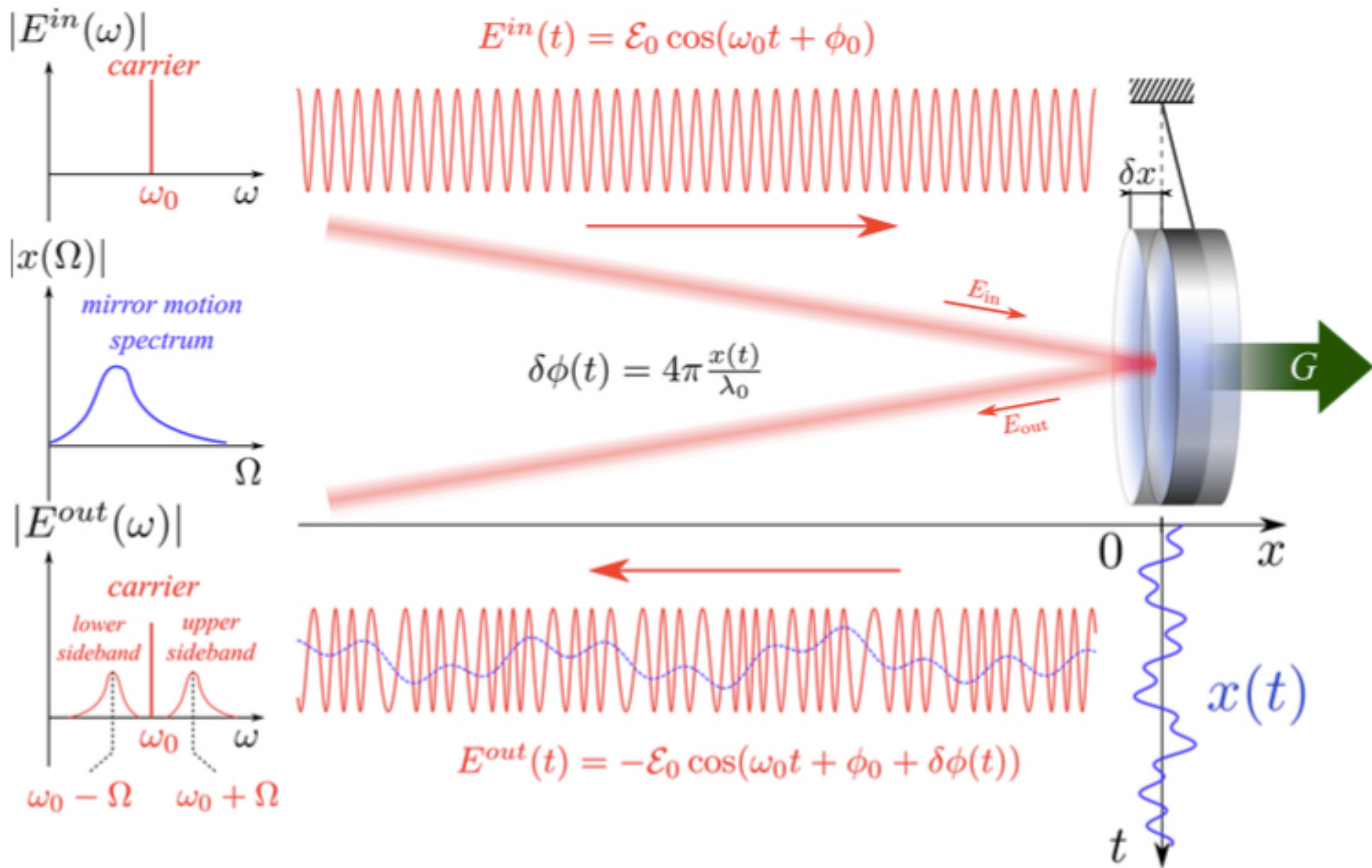
The output electric fields are phase modulated by the mirror motion



The modulation effect is encoded just in the s-quadrature of the electric field.

The light spectral component $L_s(\Omega)$ containing information of the motion x (with spectral component $X(\Omega)$) is proportional to the incoming light \mathcal{E}_0

$$L_s(\Omega) = (\omega_0 / 2 \pi c) X(\Omega) \mathcal{E}_0$$



The MICHELSON Interferometer

Two different mirrors : $r_1 = \sqrt{R_1} \neq r_2 = \sqrt{R_2}$

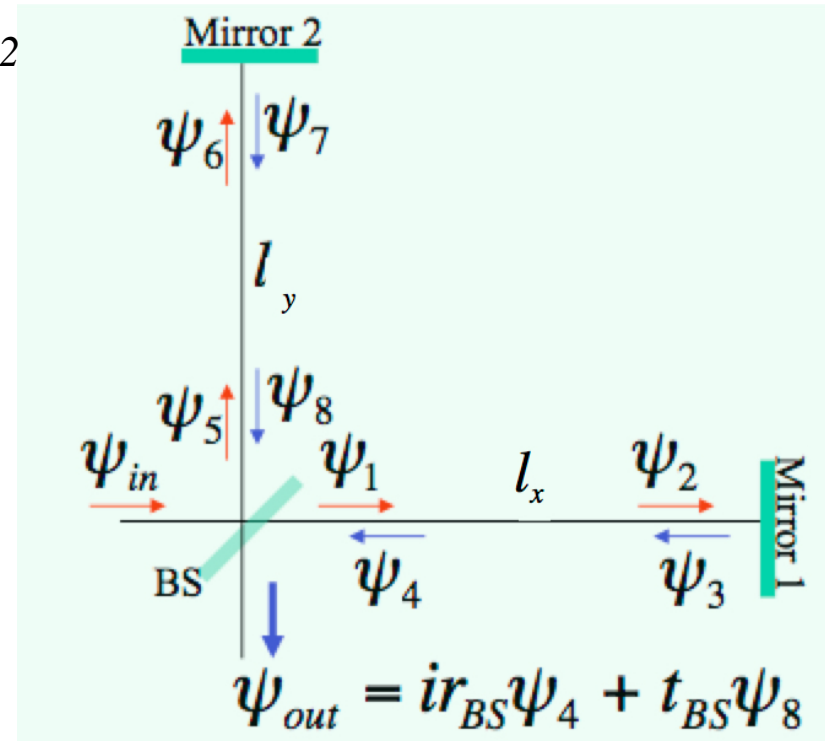
$$\psi_{in} = K e^{i\chi}$$

$$\psi_1 = t_{BS} \psi_{in} \quad \psi_5 = i r_{BS} \psi_{in}$$

$$\psi_2 = e^{-ikl_1} \psi_1 \quad \psi_6 = e^{-ikl_2} \psi_5$$

$$\psi_3 = i r_1 \psi_2 \quad \psi_7 = i r_2 \psi_6$$

$$\psi_4 = e^{-ikl_1} \psi_3 \quad \psi_8 = e^{-ikl_2} \psi_7$$



$$|\psi_{out}|^2 = P_{in} r_{BS}^2 t_{BS}^2 (r_1^2 + r_2^2) \left[1 + \frac{2r_1 r_2}{r_1^2 + r_2^2} \cos(2k\delta l) \right]$$

$$\delta l = l_x - l_y$$

$$C = \text{contrast} = \frac{2r_1 r_2}{r_1^2 + r_2^2} = \frac{P_{out}^{\max} - P_{in}^{\min}}{P_{out}^{\max} + P_{in}^{\min}}$$

At the dark fringe with $l_x = l_y$ for all light frequencies the output is zero (zero frequency noise)

In absence of a modulation signal the electric field at the output of an interferometer operated on a dark fringe

GW perturbs the arm length $h \ll l$

$$r_1 = r_2 = \sqrt{R}$$

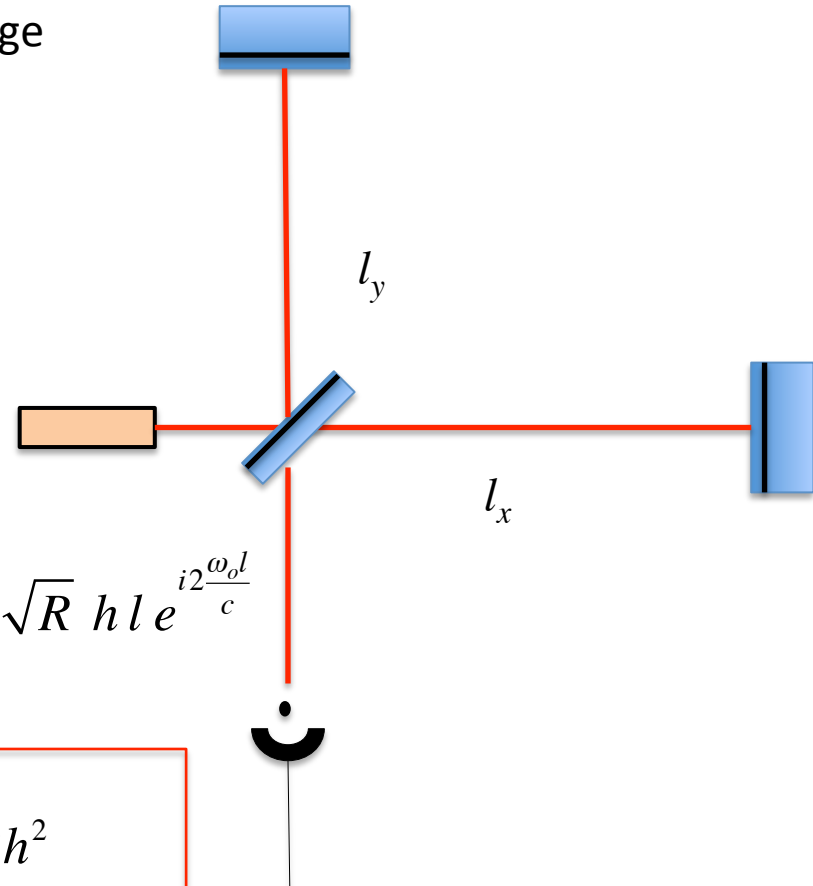
$$\Psi_{out} = -\Psi_{in} \frac{1}{2} \sqrt{R} [e^{i2\frac{\omega_o l_x}{c}} - e^{i2\frac{\omega_o l_y}{c}}]$$

$$l = \frac{l_x + l_y}{2}$$

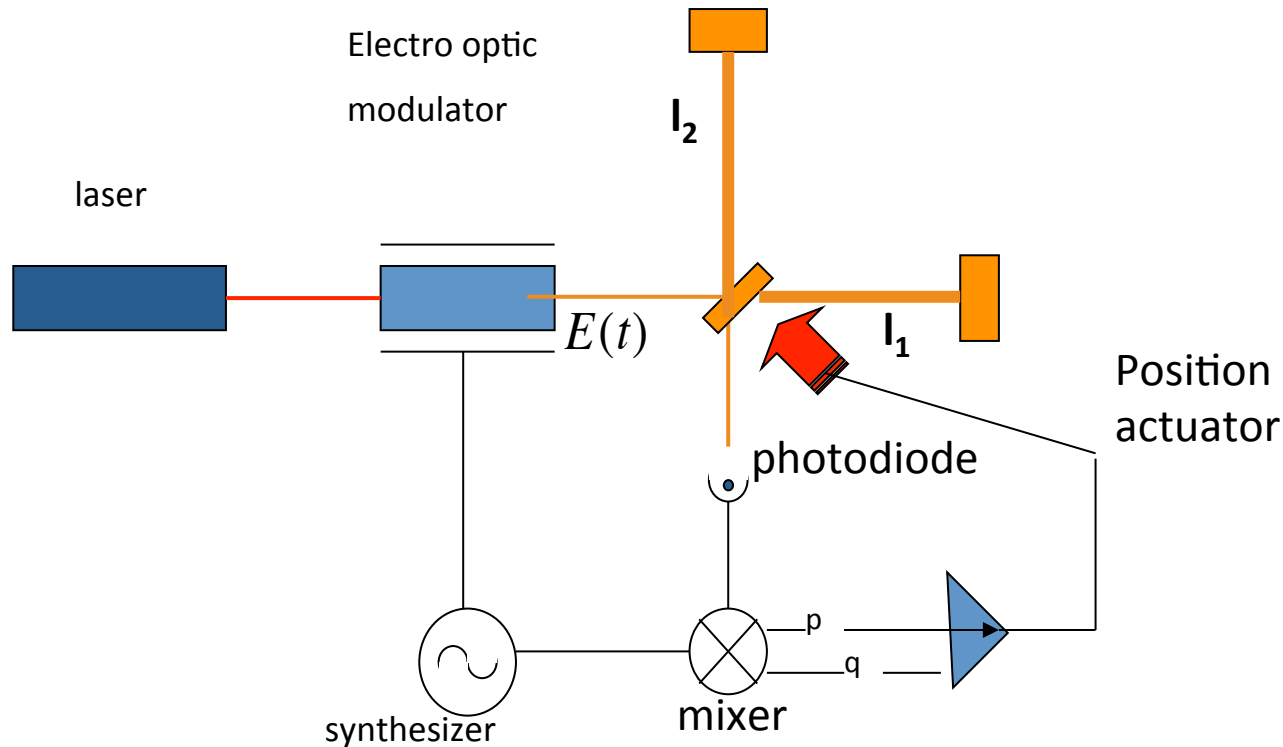
$$h = \frac{\delta l_x - \delta l_y}{l} \ll 1$$

$$\Psi_{out} \approx -i \Psi_{in} \frac{\omega_o}{c} \sqrt{R} h l e^{i2\frac{\omega_o l}{c}}$$

$$P_{out} \approx P_{in} \left(\frac{\omega_o}{c} \right)^2 l^2 h^2$$



The photodiode output depends on $h^2 \rightarrow$ detection impossible



Field at the input of the modulator

$$\Psi_{in}(t) = A_0 e^{i\omega_0 t}$$

Field at the modulator output

$$\Psi_{in}^m(t) = E_0 \cdot e^{i(\omega_0 t + \delta_m \sin \Omega t)} \approx E_0 (J_0(\delta_m) e^{i\omega_0 t} + J_1(\delta_m) e^{i(\omega_0 + \Omega)t} - J_1(\delta_m) e^{-i(\omega_0 - \Omega)t})$$

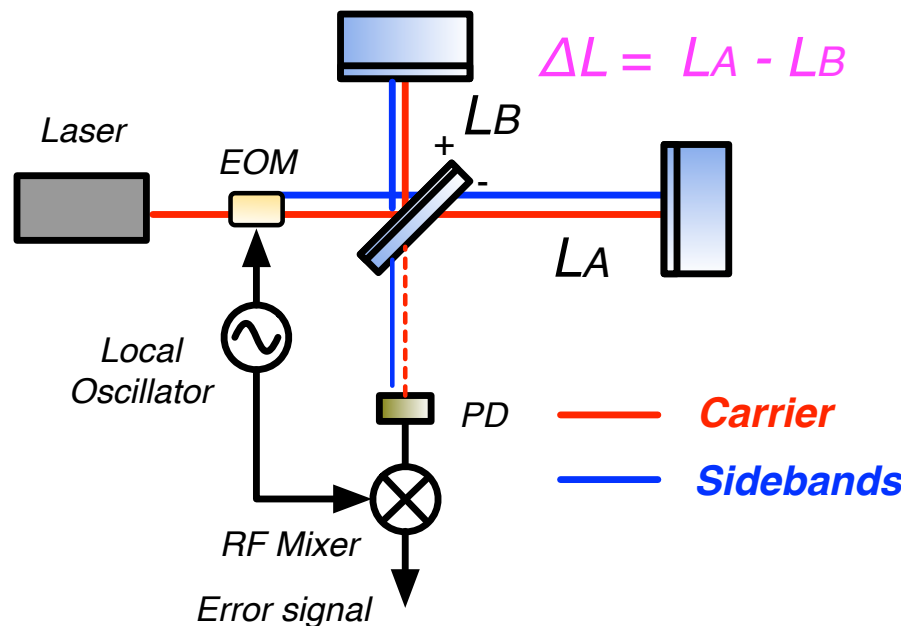
- Michelson is operated at the dark fringe

At the dark fringe, DC signals can't be a good error signal

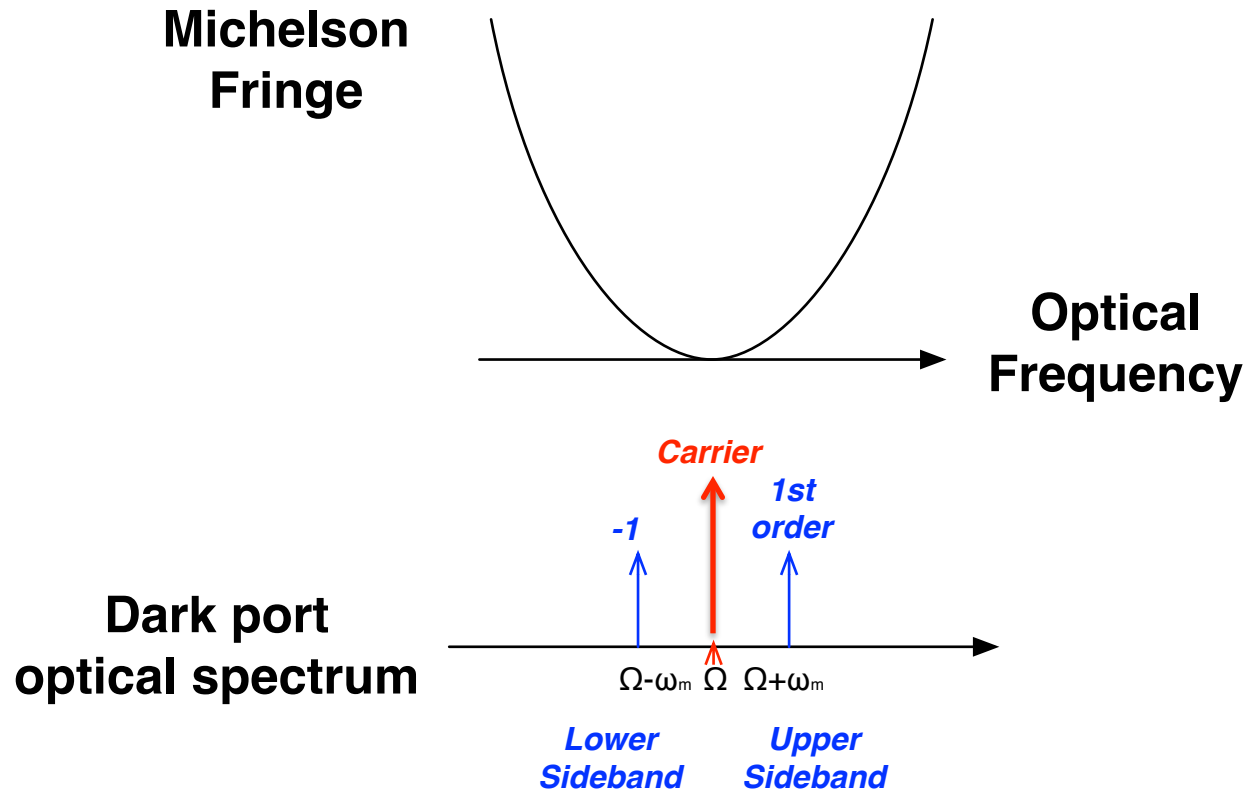
- Schnupp asymmetry:

Introduce small arm length asymmetry

=> RF sidebands leaks to the dark port

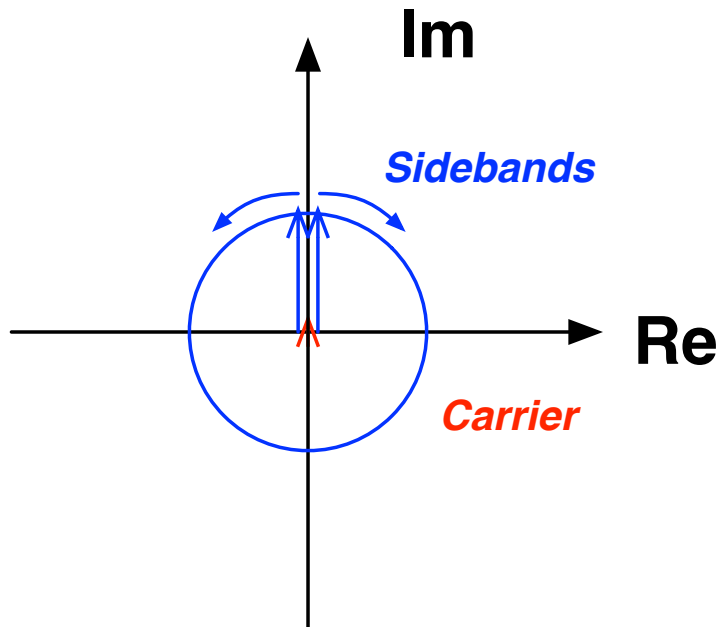


- Because of the asymmetry, the output port is no longer dark for the sidebands even if the carrier is at dark.

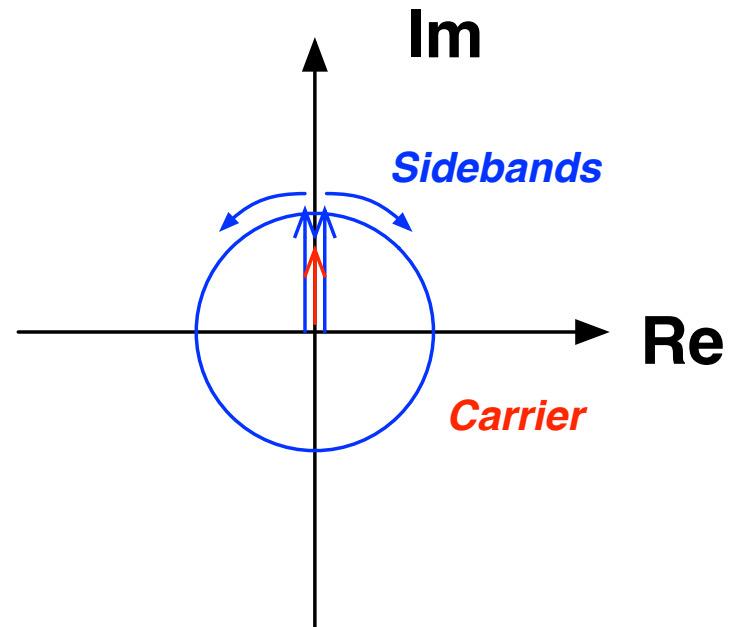


- In presence of the Schnupp asymmetry

Carrier at dark



Carrier deviates from dark finger



The interferometer output is set equal zero for the carrier component –dark fringe –insensitive to the noise (but also to the GW signal). Let us add modulation and a difference in the arms

$$\Psi_{out}^m(t) = [A_c \Psi_0 + A_+ \Psi_+ e^{i\Omega t} + A_- \Psi_- e^{-i\Omega t}] e^{i\omega_0 t}$$

$$\Delta l_{Sc} = l_x - l_y \quad \delta l_{gw} = l h$$

$$\delta l = \delta l_{gw} + \Delta l_{Sc}$$

$$\Psi_{out}^m(t) = i\Psi_{in} e^{i\frac{\omega_0 l}{c}} [J_0 l h \frac{\omega_0}{c} + 2J_1 \sin(\frac{\Omega}{c} \Delta l_{Sc}) \cos(\Omega t + 2\frac{\Omega}{c} l)]$$

PHD output:

$$P_{out} \approx P_{in} \left\{ J_0^2 \left(\frac{\omega_0}{c} \right)^2 l^2 h^2 + 2J_1^2 \frac{\omega_0}{c} [1 + \cos(2\Omega t + 4\frac{\Omega}{c} l)] \sin^2 \left(\frac{\Omega}{c} \Delta l_{Sc} \right) + 2J_0 J_1 \frac{\omega_0}{c} l h \cos(\Omega t + 2\frac{\Omega}{c} l) \sin \left(\frac{\Omega}{c} \Delta l_{Sc} \right) \right\}$$

A Schnupp asymmetry of the arms is needed

h the GW signal can appear in the phase component at Ω

The Fabry-Perot Interferometer

$$\psi_{in} = Ke^{ix}$$

$$\psi_1 = \text{unknown}$$

$$\psi_2 = e^{-ikl} \psi_1$$

$$\psi_3 = ir_2 \psi_2$$

$$\psi_4 = e^{-ikl} \psi_3$$

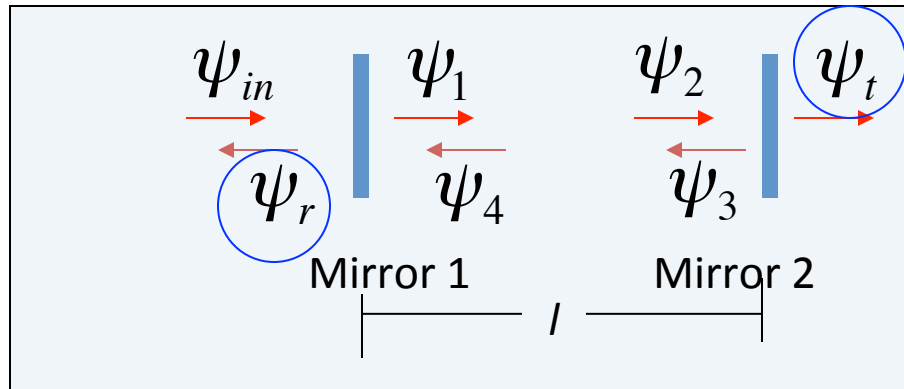
steady state: $\psi_1 = t_1 \psi_{in} + ir_1 \psi_4 \implies$

$$\psi_1 = \frac{t_1}{1 + r_1 r_2 e^{-2ikl}} \psi_{in}$$

$$\psi_4 = \frac{ir_2 t_1}{1 + r_1 r_2 e^{-2ikl}} \psi_{in}$$

Reflected wave: $\psi_r = ir_1 \psi_{in} + t_1 \psi_4 = i \frac{r_1 + r_2 e^{-2ikl}}{1 + r_1 r_2 e^{-2ikl}} \psi_{in}$

Transmitted wave: $\psi_t = t_2 \psi_2 = \frac{t_1 t_2 e^{-ikl}}{1 + r_1 r_2 e^{-2ikl}} \psi_{in}$



Wave amplitudes:

$$|\psi_r|^2 = |A_r|^2$$

$$|\psi_t|^2 = |A_t|^2$$

$$A_r = \sqrt{\frac{r_1^2 r_2^2 + 2r_1 r_2 \cos 2kl}{1 + r_1^2 r_2^2 + 2r_1 r_2 \cos 2kl}}$$

$$A_t = \sqrt{\frac{1}{1 + r_1^2 r_2^2 + 2r_1 r_2 \cos 2kl}} t_1 t_2$$

$$L = (2m + 1) \frac{\lambda}{4} \implies A_r = \min = \frac{r_2 - r_1}{1 - r_1 r_2}; A_t = \max = \frac{t_1 t_2}{1 - r_1 r_2}$$

if :

$$L = m\lambda \implies A_r = \max = \frac{r_2 + r_1}{1 + r_1 r_2}; A_t = \min = \frac{t_1 t_2}{1 + r_1 r_2}$$

The light resonates into the cavity if its phase is increased by exactly 2π each two reflections

Main cavity features

Finesse: $F = \frac{FSR}{FWHM} \approx \frac{\pi \sqrt{r_1 r_2}}{1 - r_1 r_2}$
full width @ half max

Free Spectral Range:

where $FSR = \frac{c}{2L} = \Delta \nu_{FSR}$

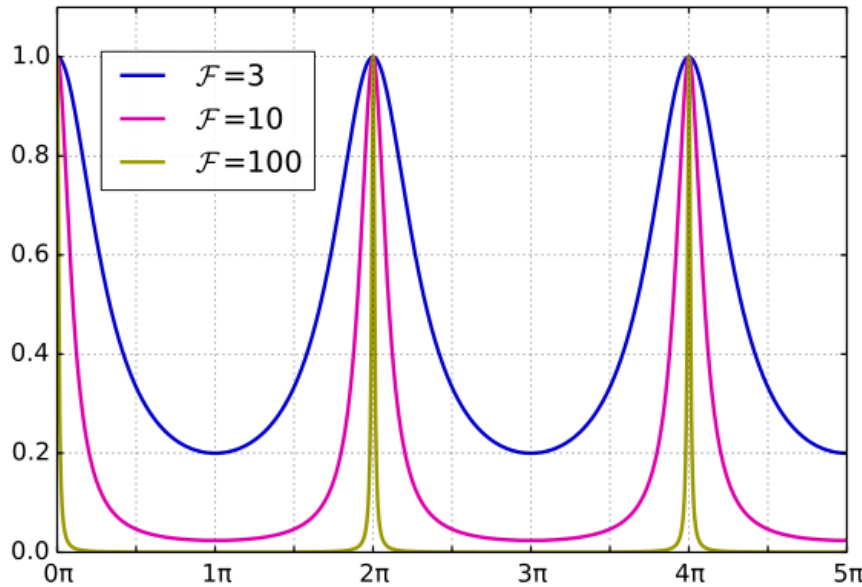
Round trip number:

note $N = \frac{F}{2\pi}$

$$FSR = \Delta \lambda_{FSR} = \frac{\lambda^2}{2L}$$

Cavity cut-off: $\omega_c = 2\pi \frac{FWHM}{2} = \pi \frac{c}{2FL} = \frac{1}{\tau_s}$ *Storage time*

Recycling factor : $FI = \left| \frac{\psi_{3;L=(2m+1)\frac{\lambda}{4}}}{\psi_{in}} \right| = \frac{t_1^2}{(1 - r_1 r_2)^2}$



$\delta\lambda \rightarrow$ bandwidth

$$\text{Finesse} = F = \Delta\lambda_{\text{FSR}} / \delta\lambda$$

$\Delta\lambda \rightarrow$ Free Spectral Range \rightarrow FSR

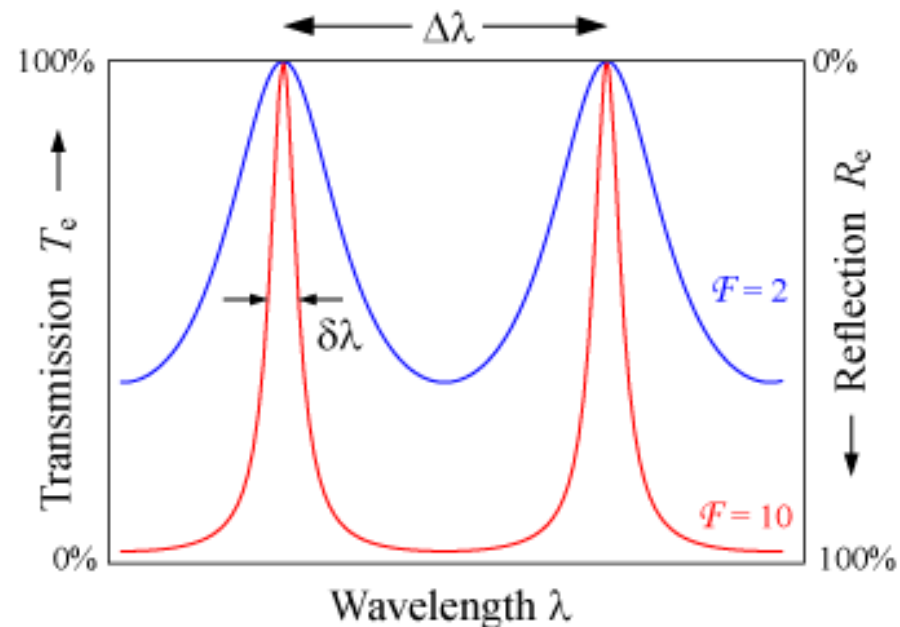
A numeric example:

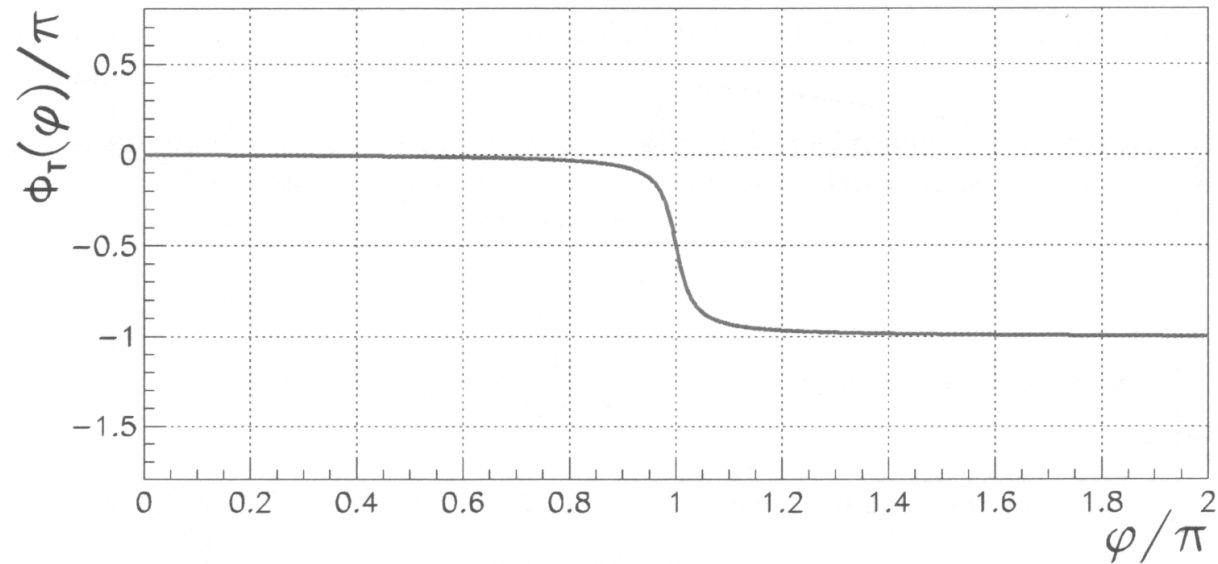
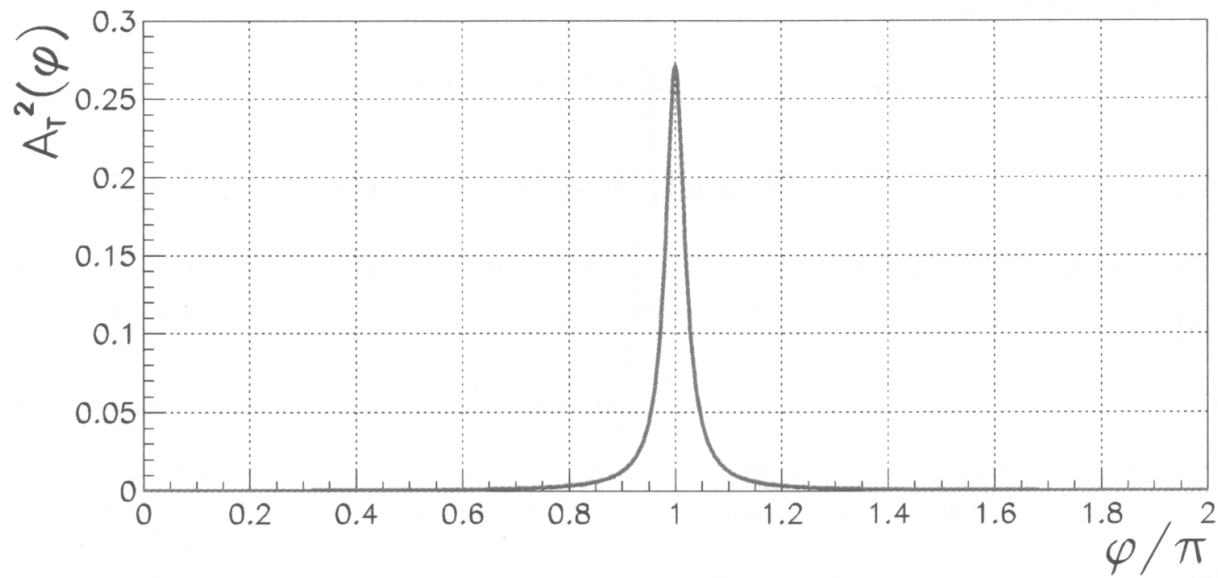
$$\lambda = 1 \text{ } \mu\text{m}$$

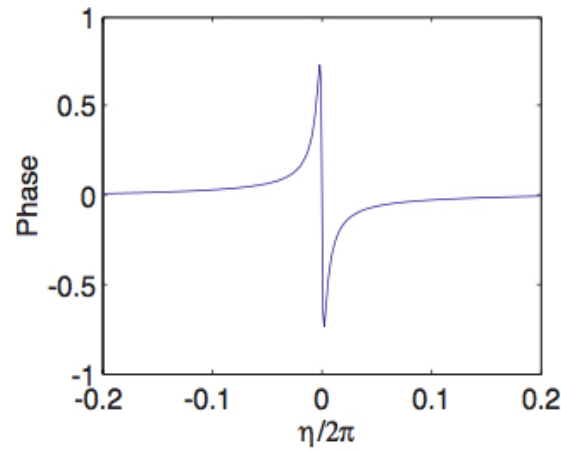
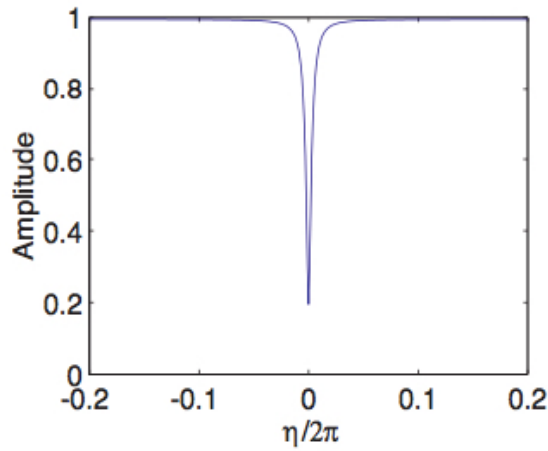
$$\text{FSR} = \lambda^2 / 2L = \frac{1}{2} (10^{-12} / 3 \cdot 10^3) = 1.7 \cdot 10^{-16}$$

$$F=10$$

$$\delta\lambda = \text{FSR} / F \sim 1.7 \cdot 10^{-17} \text{ m}$$

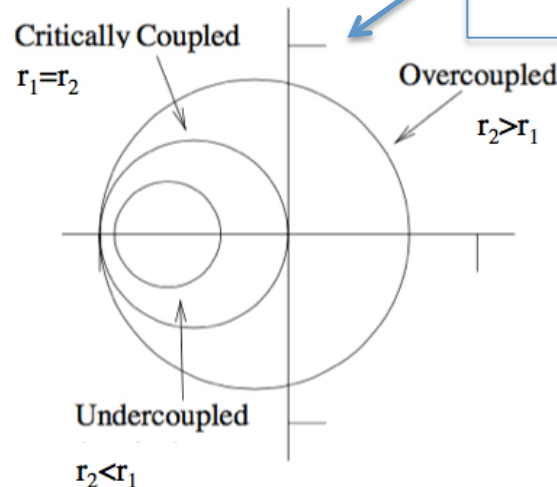
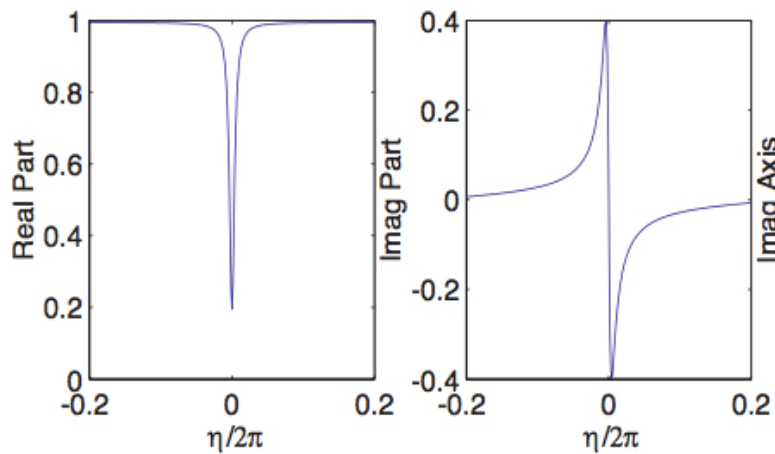






$$\frac{\Psi_r}{\Psi_{in}} = i \frac{r_1 + r_2 e^{-2ikl}}{1 + r_1 r_2 e^{-2ikl}}$$

$$\frac{\Psi_r}{\Psi_{in}} \text{ In a complex plane}$$



Over coupled cavity :almost all the light is reflected back by the cavity

*Optical cavity length
(depends on the frequency!!):*

$$L_{opt} = FL \frac{2}{\pi} \frac{1}{\sqrt{1 + (\omega / \omega_c)^2}}$$

Note: in an interferometer with FP cavities every flat noise into the system (e.g. shot noise) causes a departure linearly dependent on the frequency from the cavity cutoff.

FP-Michelson ITF

The formalism to compute Michelson output field for it is exactly the same of that adopted for the simple Michelson, once the simple mirror reflectivity has been replaced with FP cavity reflectivity:

$$r_1 = A^{(1)}_r \quad r_2 = A^{(2)}_r$$

$$\nabla^2 U(x, y, z) + k^2(x, y, z) = 0$$

Coherent monochromatic wave
emitted by the source

Along z, Hermite-Gauss complete set:

$$U(x, y, z) = \frac{A_{mn}}{w(z)} H_m\left(\sqrt{2} \frac{x}{w(z)}\right) H_n\left(\sqrt{2} \frac{y}{w(z)}\right) \left[-\frac{x^2 + y^2}{w^2(z)} - i \frac{k(x^2 + y^2)}{2R(z)} - i(kz - \phi_{nm}) \right]$$

$$w^2(z) = w_o^2 \left[1 + \left(\frac{\lambda z}{\pi w_o^2} \right)^2 \right]$$

beam divergence θ_∞

$$R(z) = z \left[1 + \left(\frac{\pi w_o^2}{\lambda z} \right)^2 \right]$$

$$\phi = (m + n + 1) \arctan\left(\theta_\infty \frac{z}{w_o}\right)$$

In a mode-matched cavity, if the cavity is stable and $R(z)$ and $w(z)$ are the same after an arbitrary number of reflections,

$$\nu_0 = c/2d$$

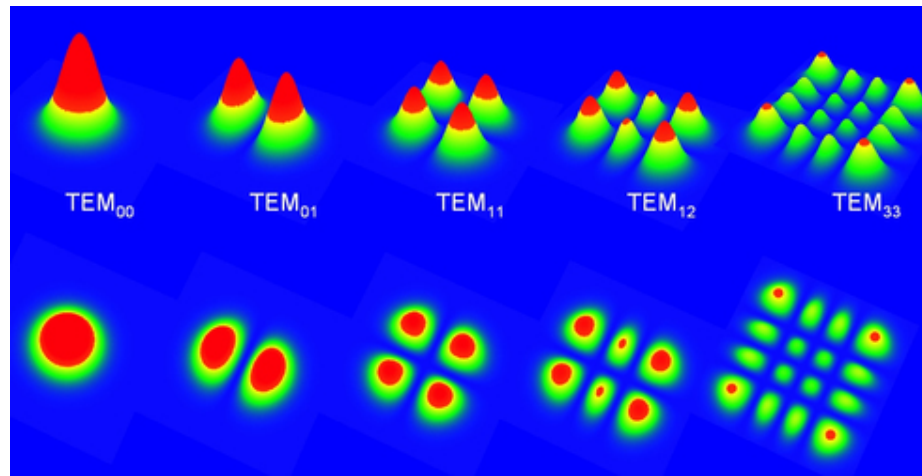
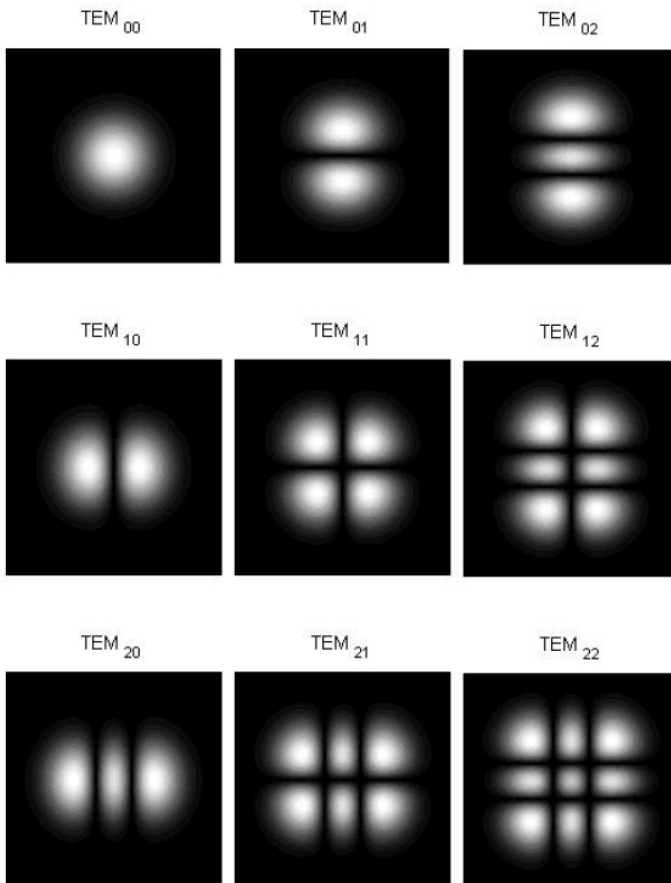
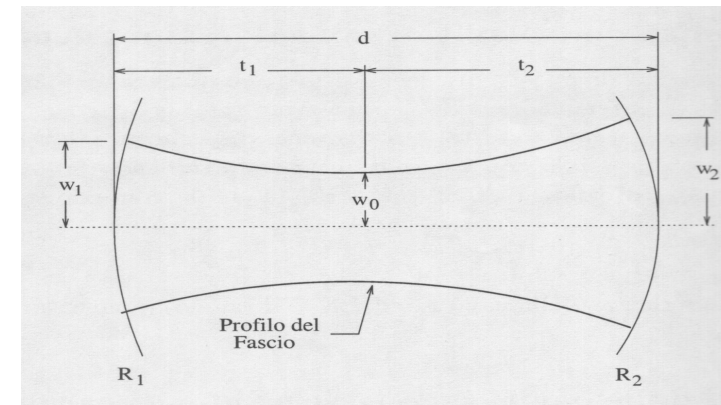
TEM_{mn} resonant condition

$$\nu = \nu_0 \left[(q+1) + \frac{1}{\pi} (m+n+1) \arccos \sqrt{\left(1 - \frac{d}{R_1}\right) \left(1 - \frac{d}{R_2}\right)} \right]$$

arb.integer mode order

Cavity stability

Longitudinal beam profile



Optical Cavity Stability

Mirror Curvatures ρ_1 , ρ_2

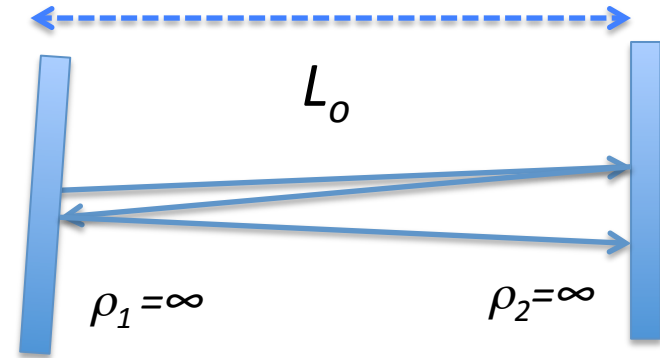
Stability parameter $\rightarrow g_i = 1 - (L / \rho_i)$

with $i=1,2$ $L=nL_o$ and n intra-cavity refraction index

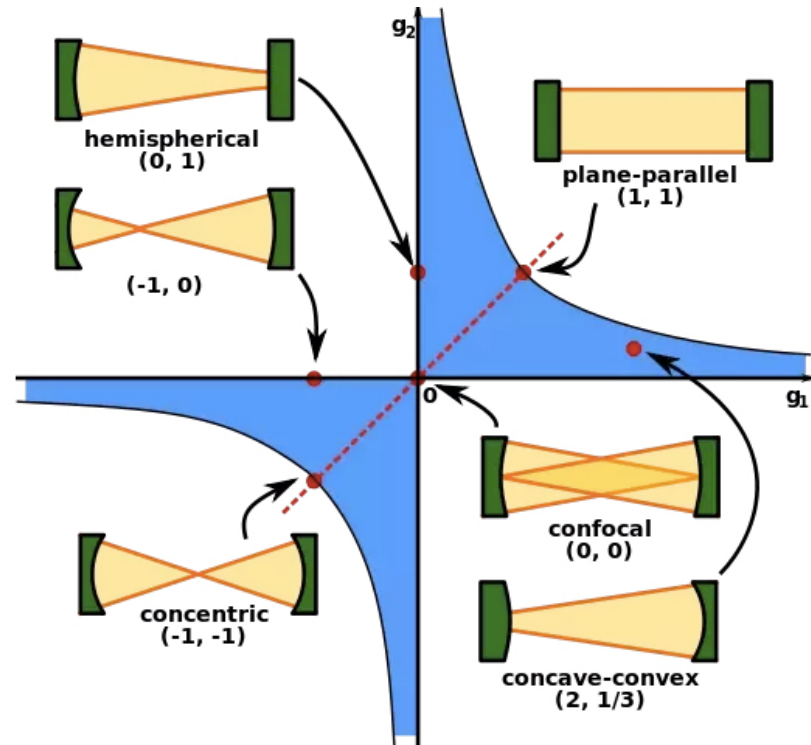
Stability Condition $\rightarrow 0 \leq g_1 g_2 \leq 1$

Marginally stable cavities $g_1 g_2 = 1$

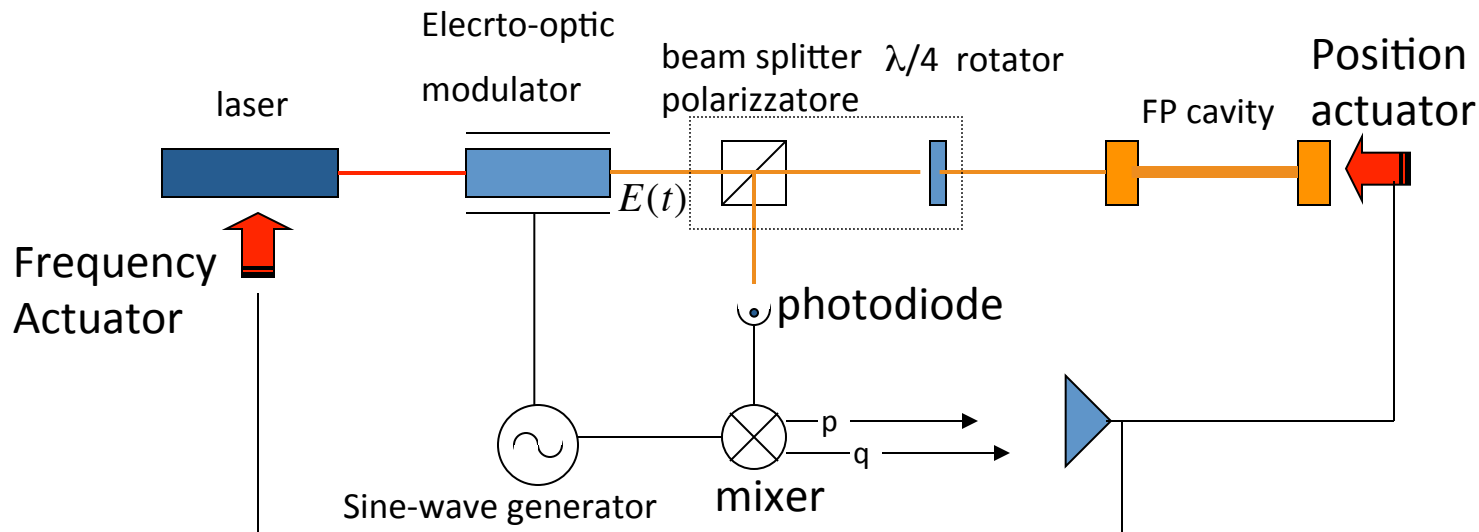
A cavity is stable if a ray launched inside the resonator parallel to the optical axis remains inside the resonator after an infinite number of bounces . If the ray is slightly off axis , it will be reflected in a direction to bring it back toward the center of the bore



*Conditionally stable
i.e. Stable if the mirrors are parallel*



In 1946 R. Pound defined the strategy, widely used in radio physics, that today is the standard method to stabilize the lasers and in general the optical cavities.



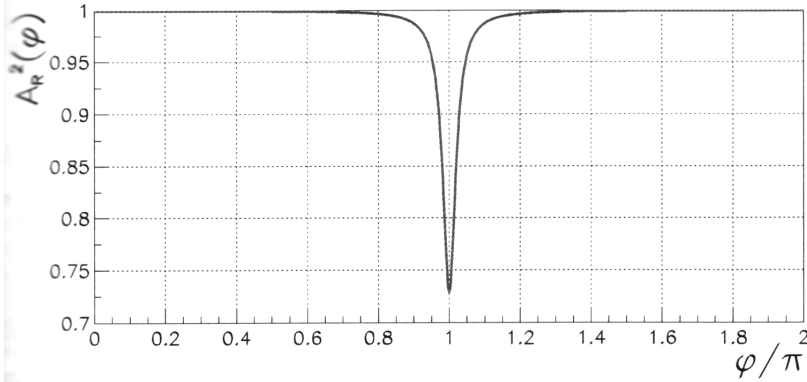
$$E(t) = E_0 \cdot e^{i(\omega_0 t + m \cos \Omega t)} \approx E_0 (J_0(m) e^{i\omega_0 t} + J_1(m) e^{i(\omega_0 + \Omega)t} - J_1(m) e^{i(\omega_0 - \Omega)t})$$

$$J_0(m) = 1; \quad J_1(m) = m/2$$

$$\begin{aligned} \text{Re}\{E(t)\} &= E_0 \cos(\omega_0 t + m \cos \Omega t) \approx E_0 \cdot \cos \omega_0 t - E_0 \sin \omega_0 t \cdot (m \cos \Omega t) \\ &= E_0 \cdot \cos(\omega_0 t) - m \frac{E_0}{2} \sin(\omega_0 + \Omega)t - m \frac{E_0}{2} \sin(\omega_0 - \Omega)t + \dots \end{aligned}$$

$m \ll 1$

Lock of a single F-P cavity



The cavity length is set such that the carrier (ω_0) is resonant

The side bands due to the modulation at Ω are anti-resonant $\{\Omega = (N+1/2) (\pi c/l)\}$

When we have the perturbation of the cavity length z , i.e. $\phi_d = -2 k z$:

Photo-diode output :

$$p = -4P_{in}J_0J_1r_2(1+r_2^2)r_1t_1^2 \frac{\sin 2kz}{1+r_1^4r_2^4 - 2r_1^2r_2^2 \cos 4kz}$$

$$q = 0$$

$$\psi_t = i \frac{r_1 + r_2 e^{-2ikl - 2ikz \pm i2\frac{\Omega}{c}l}}{1 + r_1 r_2 e^{-2ikl - 2ikz \pm i2\frac{\Omega}{c}l}} \psi_{in}$$

$$e^{-2ikl} = -1$$

$$\psi_t = i \frac{r_1 - r_2 e^{-2ikz \pm i2\frac{\Omega}{c}l}}{1 - r_1 r_2 e^{-2ikz \pm i2\frac{\Omega}{c}l}}$$

$$e^{\pm 2\frac{\Omega}{c}l} = e^{\pm i\pi} \quad \psi_t = i \frac{r_1 + r_2 e^{-2ikz}}{1 + r_1 r_2 e^{-2ikz}} \psi_{in}$$

The demodulated signal in phase at Ω contains the information on the length variation

If the FINESSE $F \gg 1$

$$\frac{dp}{dz} \cong J_0(\delta_m) J_1(\delta_m) P_{in} \frac{8F}{\lambda}$$

Higher FINESSE higher sensitive to length variation

Control dynamics

$$\delta z = \frac{1}{10} \frac{FWHM}{2} = \frac{FSR}{10 \cdot 2 \cdot F}$$

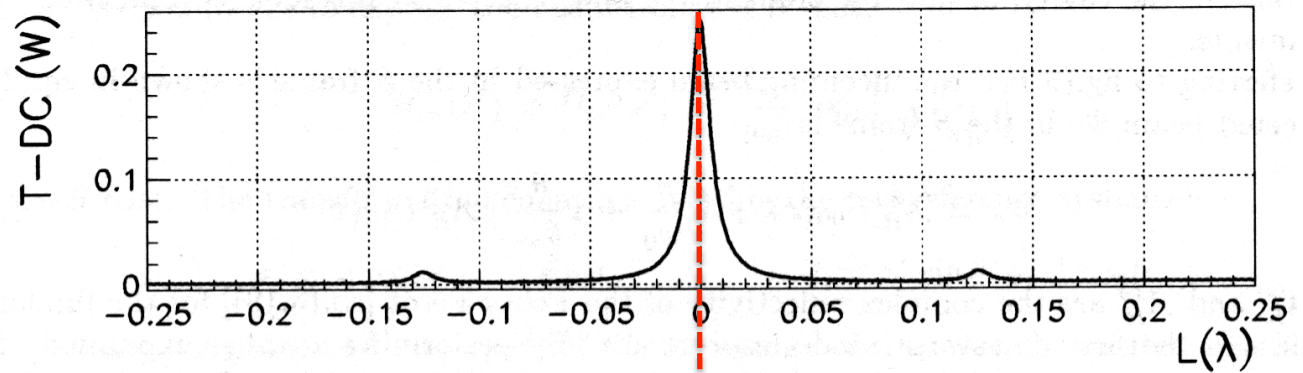
$\lambda / 2$

Linearity Range

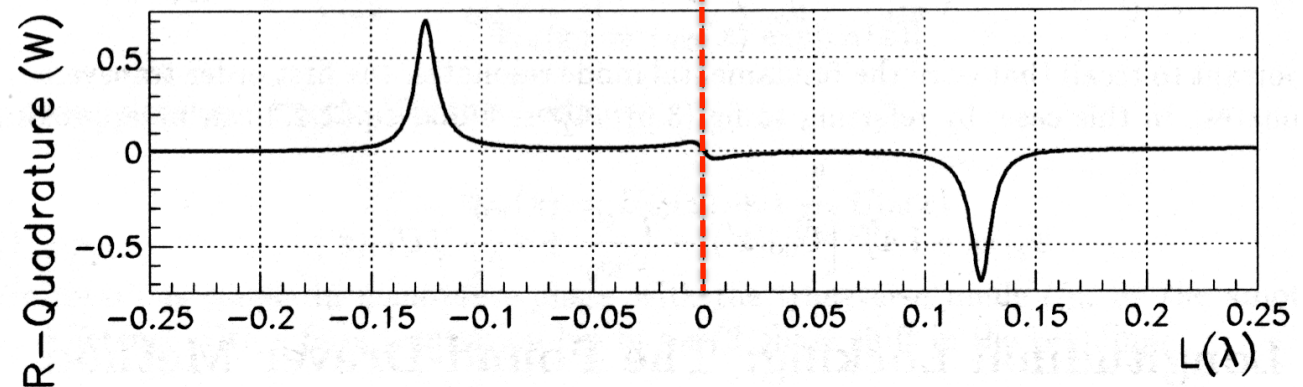
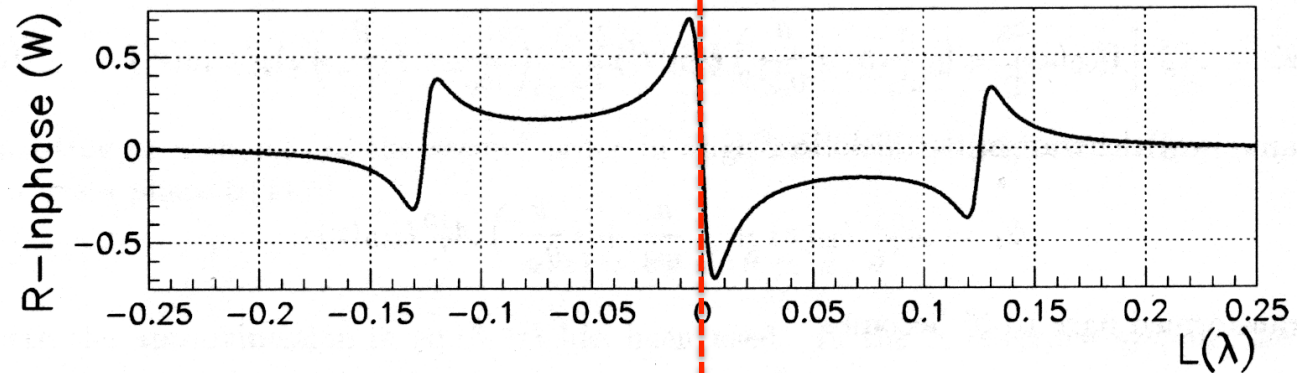
$$FWHM = \frac{\lambda}{2F}$$

Control dynamics and linearity range reduced

Amplitude of the
reflected light



Pound-Drever Signals



- Longer arms yield larger GW signals
- Increase the effective length of the arms by bouncing the light back and forth within them, or *folding* the beam
- *In a FP* kept near resonance, the phase of the output beam is very sensitive to the distance between the mirrors, consequence of the long storage time of the light into the cavity

- Light storage time
$$\tau_s = \frac{2}{\pi} \frac{FL}{c}$$

- Equivalent number of bounces
$$N_{eq} = \frac{F}{2\pi}$$

Michelson+Fabry-Perot

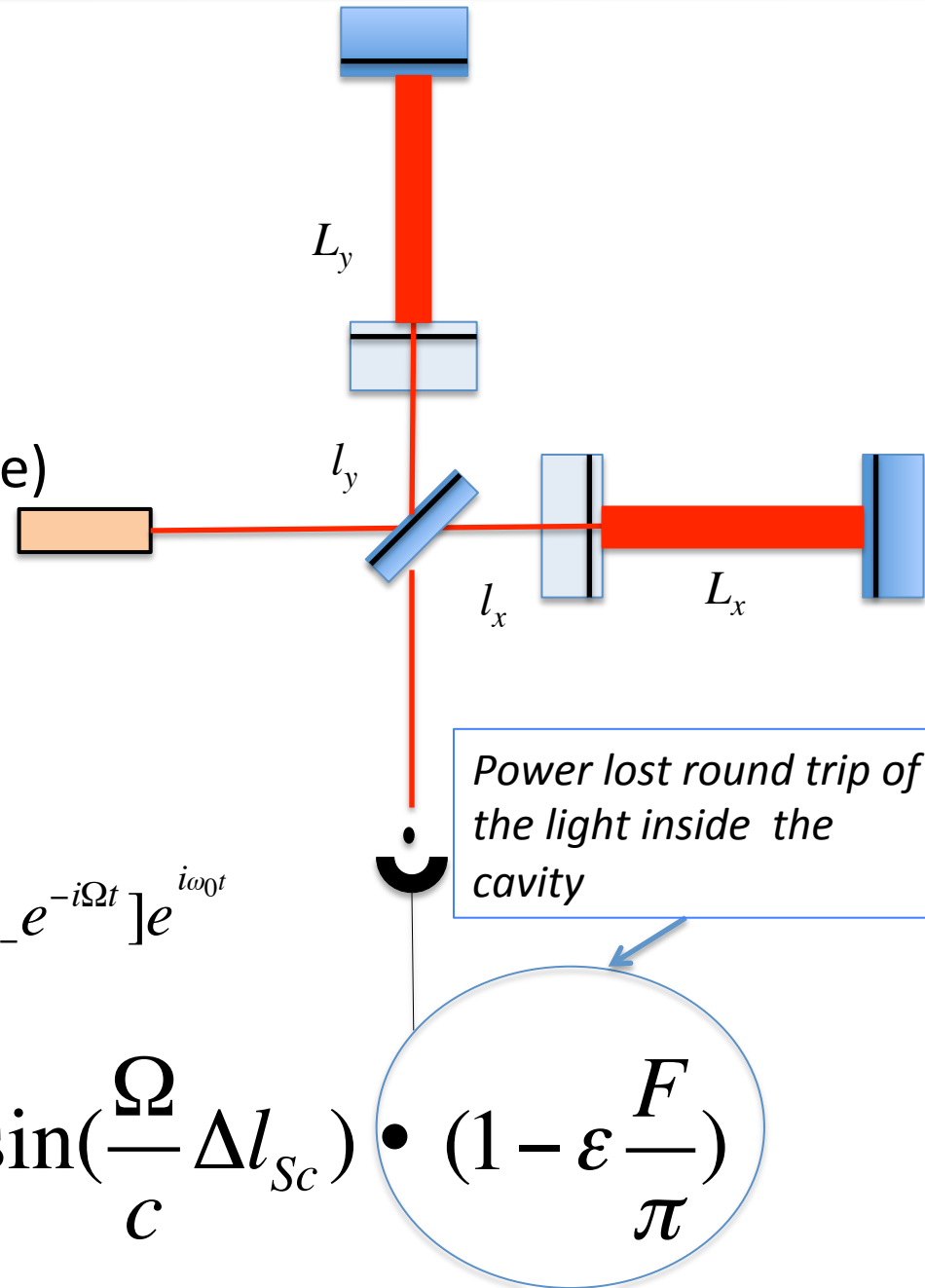
$l_{x,y} \rightarrow$ distance between the beam splitter and the first, or input (m range)

$L_x = L_y = L \rightarrow$ lengths of the Fabry-Perot cavities in the arms (km range)

$\Delta l_{Sc} = l_x - l_y$ Schnupp asymmetry \rightarrow between beamsplitter and arm cavities (tens of cm)

$$\Psi_{out}^m(t) = [A'_c \Psi_0 + A'_+ \Psi_+ e^{i\Omega t} + A'_- \Psi_- e^{-i\Omega t}] e^{i\omega_0 t}$$

$$P_{out} = 2P_{in} J_0 J_1 \frac{F}{\pi} \frac{\omega_o}{c} L h \sin\left(\frac{\Omega}{c} \Delta l_{Sc}\right) \cdot \left(1 - \varepsilon \frac{F}{\pi}\right)$$



Michelson+ Fabry-Perot +Recycling

The most of the carrier gets reflected back towards the laser

Including the recycling mirrors means to add an other FP cavity

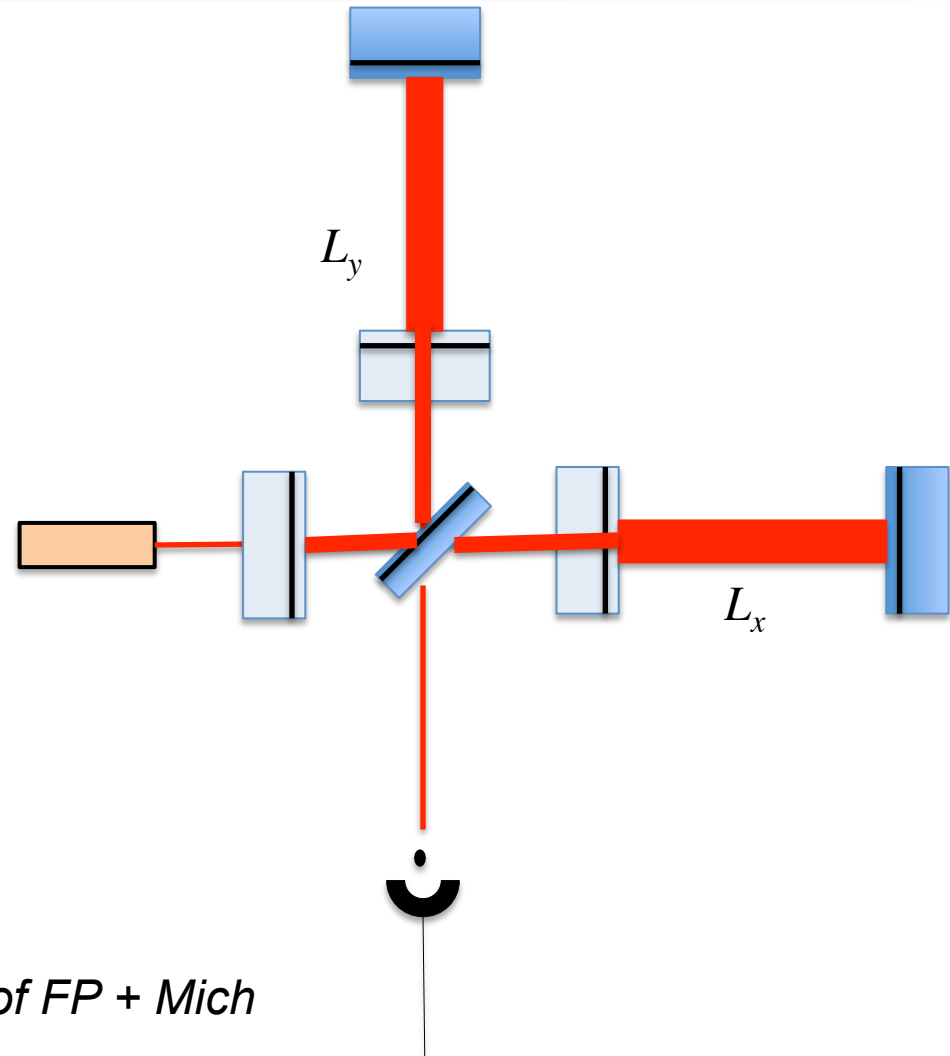
Transmission and reflection coefficients of the interferometer, including power recycling, are those of the cavity with the recycling mirror and the rest of the interferometer acting as the output mirror.

$$t_{rc} = \frac{t_{rm} t_{ifo} e^{i \frac{\omega_o}{c} l_{rc}}}{1 - r_{rm} r_{ifo} e^{i \frac{\omega_o}{c} l_{rc}}}$$

t_{ifo} and r_{ifo} \rightarrow transmission and reflection of FP + Mich

We separate the contribution of carrier and sidebands

$$\Psi_{out}^m(t) = [A_c'' \Psi_0 + A_+'' \Psi_+ e^{i\Omega t} + A_-'' \Psi_- e^{-i\Omega t}] e^{i\omega_0 t}$$



$$r_{\pm}^{ifo} = -e^{\pm i \frac{\Omega}{c} (l_x + l_y)} \cos\left(\frac{\Omega}{c} \Delta l_{Sc}\right)$$

$$t_{\pm}^{ifo} = \mp e^{\pm i \frac{\Omega}{c} (l_x + l_y)} \sin\left(\frac{\Omega}{c} \Delta l_{Sc}\right)$$

These relations, inserted in t_{ifo} and r_{ifo} bring to the conclusion that the resonance condition is

$$e^{i \frac{\Omega}{c} (2l_{rc} + l_x + l_y)} = -1$$

the optimum recycling is obtained for

$$r_{rm} = \cos\left(\frac{\Omega}{c} \Delta l_{Sc}\right)$$

the transmission coefficients for the sidebands is

$$t_{\pm}^{ifo} = \pm i e^{\mp i \frac{\Omega}{c} l_{rc}}$$

$$r_c^{ifo} = e^{i\frac{\omega_o}{c}(l_x+l_y)} \left(1 - \frac{1}{\pi} F_{ac} \varepsilon\right)$$

$$t_c^{ifo} = ie^{i\frac{\omega_o}{c}(l_x+l_y)} 2F \frac{L\omega_o}{\pi c} h \left(1 - \frac{1}{\pi} F_{ac} \varepsilon\right)$$

In this case the resonance condition is

$$e^{i\frac{\omega_o}{c}(2l_{rc}+l_x+l_y)} = 1$$

the optimum coupling for the recycling is obtained for

$$r_{rm} = \left(1 - \frac{1}{\pi} F_{ac} \varepsilon\right)$$

$$t_c^{rc} = ie^{-i\frac{\omega_o}{c}l_{rc}} 2 \frac{F_{ac}}{\pi} \sqrt{\frac{F_{rc}}{\pi}} \frac{\omega_o}{c} Lh$$

with

$$F_{rc} \approx \frac{\pi}{t_i^2}$$

Summary: carrier + sidebands conditions

Recycling reflectivity

$$\cos\left(\frac{\Omega}{c} \Delta l_{sc}\right) = r_{rm} = \left(1 - \frac{1}{\pi} F_{ac} \varepsilon\right)$$

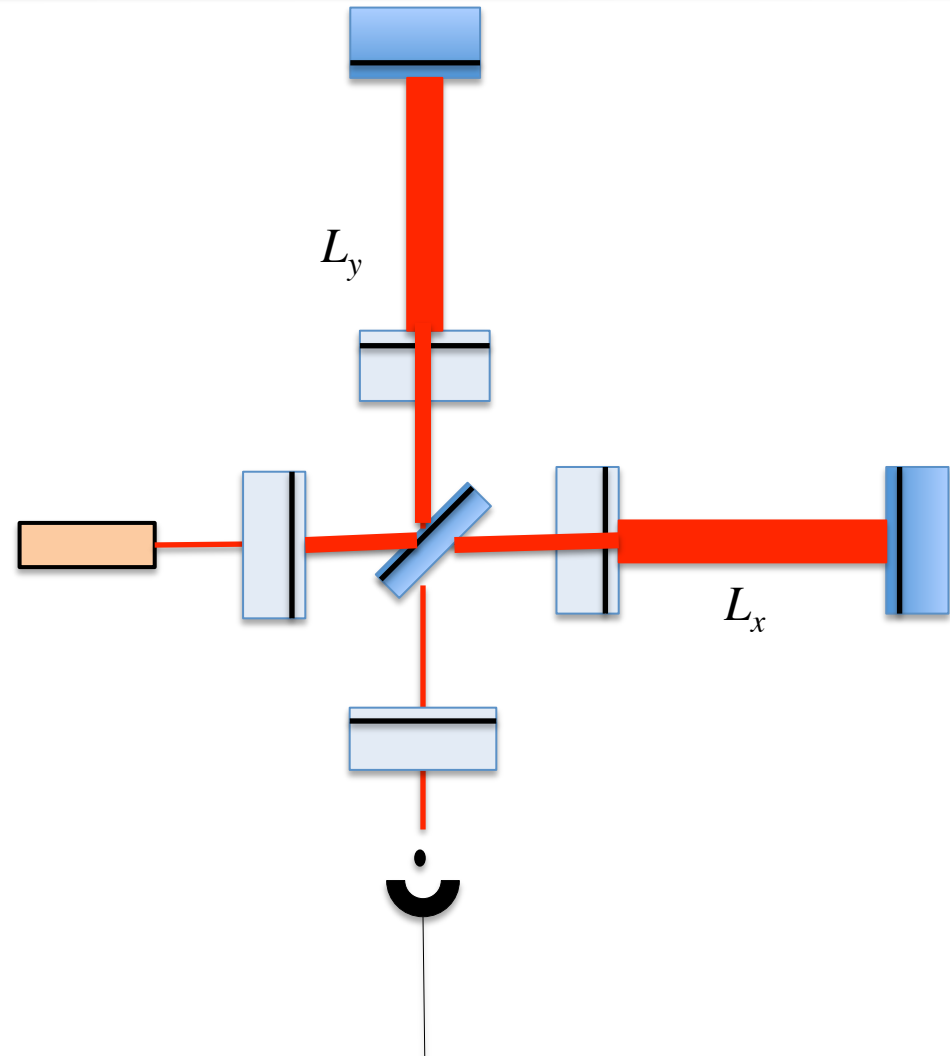
Resonance conditions

$$e^{i\frac{\omega_o}{c}(2l_{rc}+l_x+l_y)} = 1$$

$$e^{i\frac{\Omega}{c}(2l_{rc}+l_x+l_y)} = -1$$

SR mirror reflects back into ITF adding coherently with more signal due to GW.

In practice we store the signal storage making more efficient the transfer of power from carrier to sidebands



Two possible configuration:

Tuned \rightarrow SR cavity resonates at ω_0

Untuned \rightarrow SR cavity does not resonate at ω_0

- The signal sidebands will be sent back into the interferometer and
 - they can be coherently enhanced
 - or
 - used to coherently extract more sideband amplitude from the arm cavities.

In the later case, the *SR mirror is placed at a position where the carrier is resonant in the SR cavity which increases the effective transmissivity of the ITMs*. The subsequent reduction in the finesse increases the bandwidth of the entire detector. This is called resonant sideband extraction.

- Changing the position of the SRM will increase the peak displacement sensitivity in a position dependent specific frequency range but will also reduce the bandwidth of the detector. This is commonly known as *detuned signal recycling or detuned resonant sideband extraction*.

Static output

$$P_{out} = 2P_{in} J_o J_1 \frac{F_{ac}}{\pi} \sqrt{\frac{F_{rc}}{\pi}} \frac{\omega_o}{c} L h$$

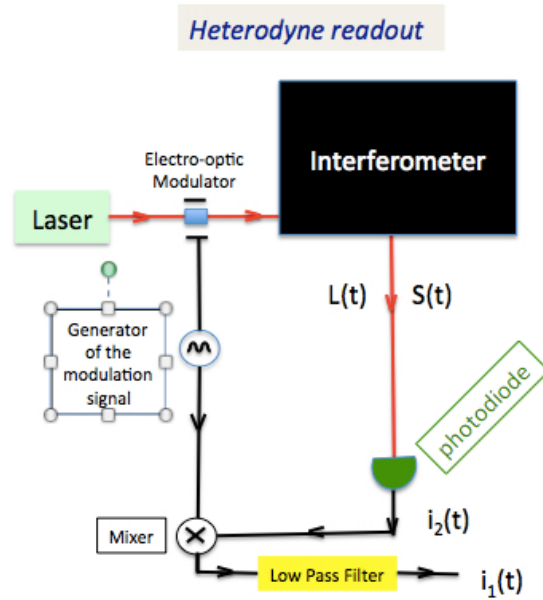
As we have shown before, the finite storage time of the arm cavities introduces a single pole in the response of the interferometer with a cut-off frequency

$$f_o = c / 4L F_{ac}$$

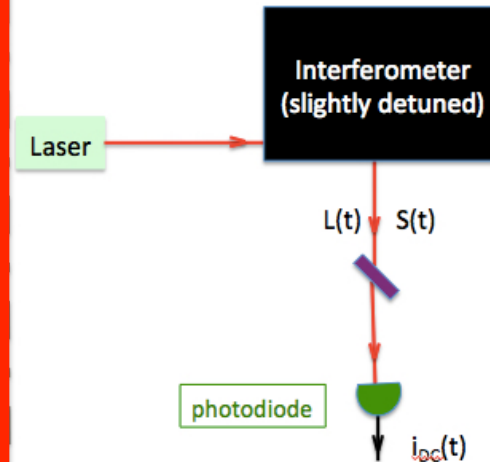
Spectral output

$$P_{out}(f) = 2P_{in} J_o J_1 \frac{F_{ac}}{\pi} \sqrt{\frac{F_{rc}}{\pi}} \frac{\omega_o}{c} L \sqrt{\frac{f_o^2}{f_o^2 + f^2}} h(f)$$

RF sidebands are modulated onto the light at the input of the Michelson and the out is demodulated via mixer driven by the RF signal



DC readout



The local oscillator is the fraction of the carrier light that leaks into the signal port due to the residual interferometer asymmetry of the arms.

- **Advantage:** the DC detection to the respect of the heterodyne readout: the shot noise contribution from frequencies twice the heterodyne frequency does not exist
- **Disadvantage:** an increased coupling of laser power noise

$$S_{\text{hom}}(\Omega) = \frac{h^2}{2K} \frac{1}{t_{rc}^2} \frac{1}{|D_1 \sin \phi_{\text{hom}} + D_1 \cos \phi_{\text{hom}}|^2} \cdot$$

$$\cdot \{[(C_{11} \sin \phi_{\text{hom}} + C_{21} \cos \phi_{\text{hom}})^2 + (C_{21} \sin \phi_{\text{hom}} + C_{22} \cos \phi_{\text{hom}})^2] +$$

$$+ \left[\frac{|D_+|^2 + |D_-|^2}{|D_+ e^{-i\phi_{\text{dem}}} + D_- e^{-i\phi_{\text{dem}}}|^2} \right] \}$$

2 Ω terms

$$S_{\text{hom}}(\Omega) = \frac{h^2}{2K} \frac{1}{t_{rc}^2} \frac{1}{|D_1 \sin \phi_{\text{hom}} + D_1 \cos \phi_{\text{hom}}|^2} \cdot$$

$$\cdot [(C_{11} \sin \phi_{\text{hom}} + C_{21} \cos \phi_{\text{hom}})^2 + (C_{21} \sin \phi_{\text{hom}} + C_{22} \cos \phi_{\text{hom}})^2]$$

Interferometer Control

- **Coherent** (*Global Control*): the operation point has to be locked by using ITF light. Angular control performed coherently with respect to the light into the ITF by means of beam wavefront sensing (i.e. detection of $TEM_{01,10}$ modes due to misalignment)
 - ➔ relative position of mirrors is correlated by light
- **Incoherent** (*Local Control*) : the mirrors are controlled by means of independent ground-based sensors and quasi-inertial actuators.
 - ➔ relative position of mirrors is correlated by the ground

Global Control

Power recycling Interferometer

Feedback control system acting on the mirrors without reintroducing noise in the detection band.

Four independent lengths have to be controlled

$l_{rec} = 2 l_{rc} + l_x + l_y \rightarrow$ Power recycling cavity length

$\Delta l = l_x + l_y \rightarrow$ Asymmetry length to be on dark fringe

$L_x \rightarrow$ Length of the first long arm

Typical lock accuracy to be achieved $10^{-12} \text{ m}_{\text{rms}}$

$L_y \rightarrow$ Length of the second long arm

Signal recycling Interferometer

+

$l_{sec} = 2 l_{sr} + l_x + l_y \rightarrow$ Signal recycling cavity length

In practice the control is based the following physical degrees of freedom

$$L_x - L_y \rightarrow \text{DARM}$$

$$(L_x + L_y)/2 \rightarrow \text{CARM}$$

$$l_x - l_y \rightarrow \text{MICH}$$

$$l_{rec} = l_{rc} + (l_x + l_y)/2 \rightarrow \text{PRLC}$$

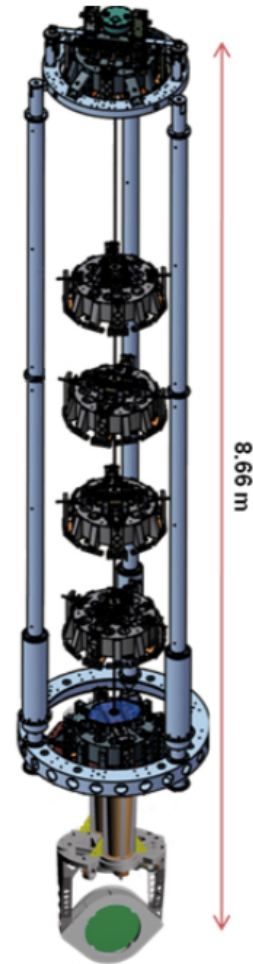
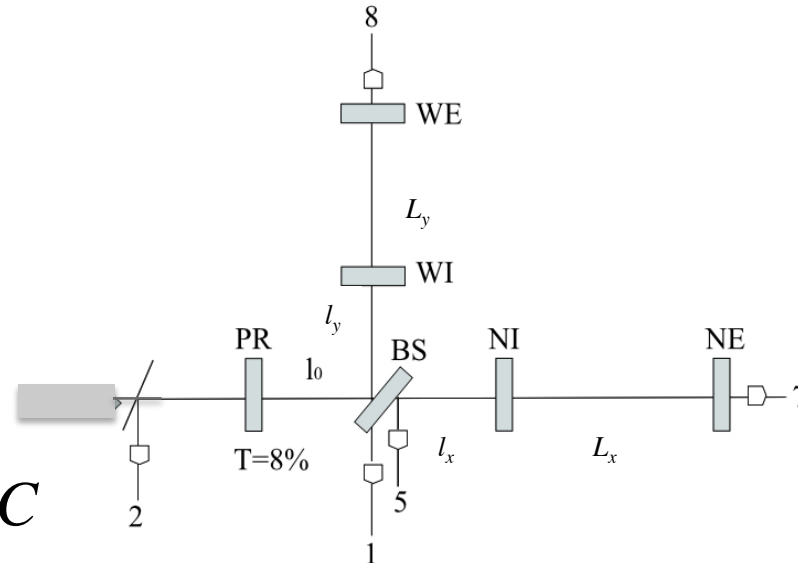
$$l_{sc} = l_{sr} + (l_x + l_y)/2 \rightarrow \text{SRLC}$$

The mirror suspended swings \rightarrow Local control reduction $\rightarrow 1\mu\text{m/s}$, then

Step 0 evaluate the residual

Step 1 actuate to reduce further the residual velocity

Step 2 feedback loop engaged



$B1 \rightarrow$ DARM, the GW signal

$B2 \rightarrow$ ITF Common mode

$B5 \rightarrow$ central cavity

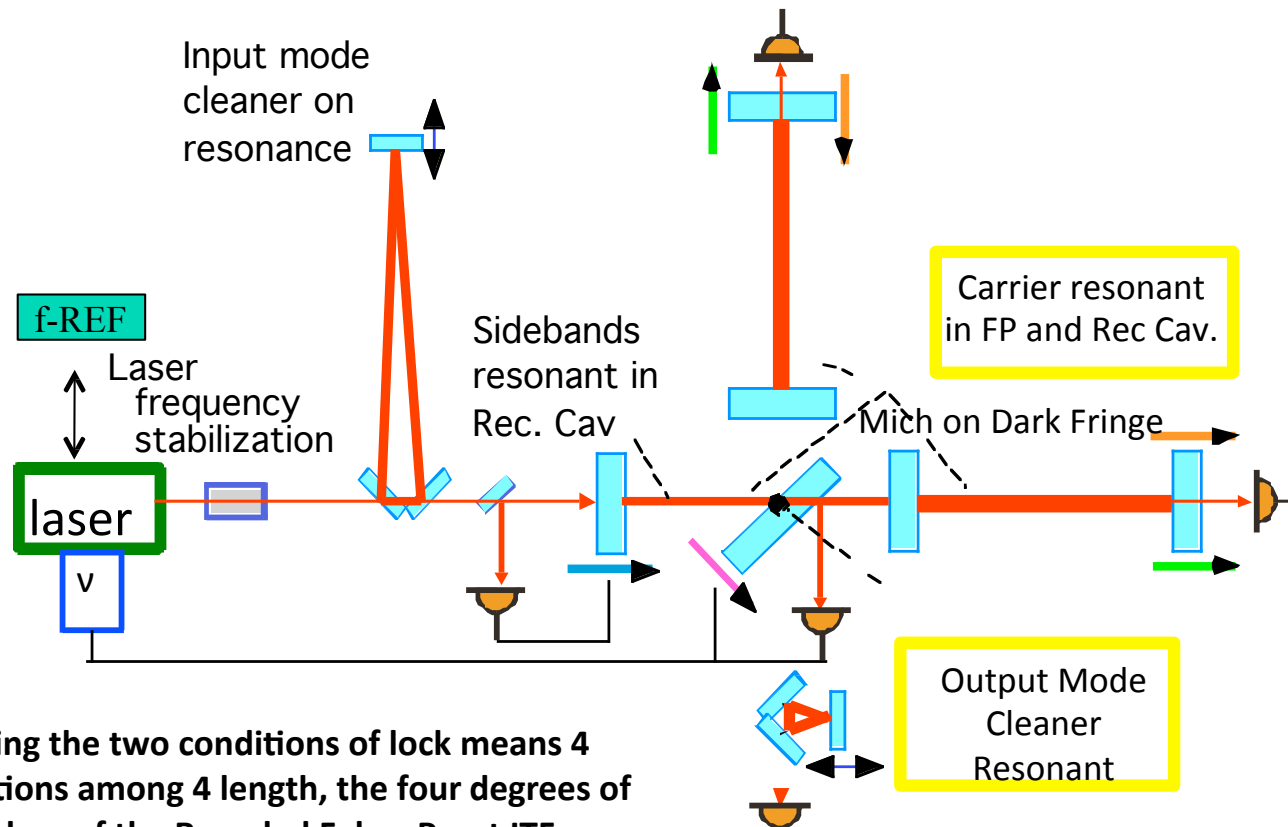
$B7 \approx L_x$

$B8 \approx L_y$

Driving Matrix

Mirror/ DOF	DARM	CARM	MICH	PRCL	SRC
NE	-1/2	-1			
WE	1/2	-1			
BS			$1/\sqrt{2}$		
PR			-1/2	-1	
SR			1/2		-1

Basic interference setup: two conditions

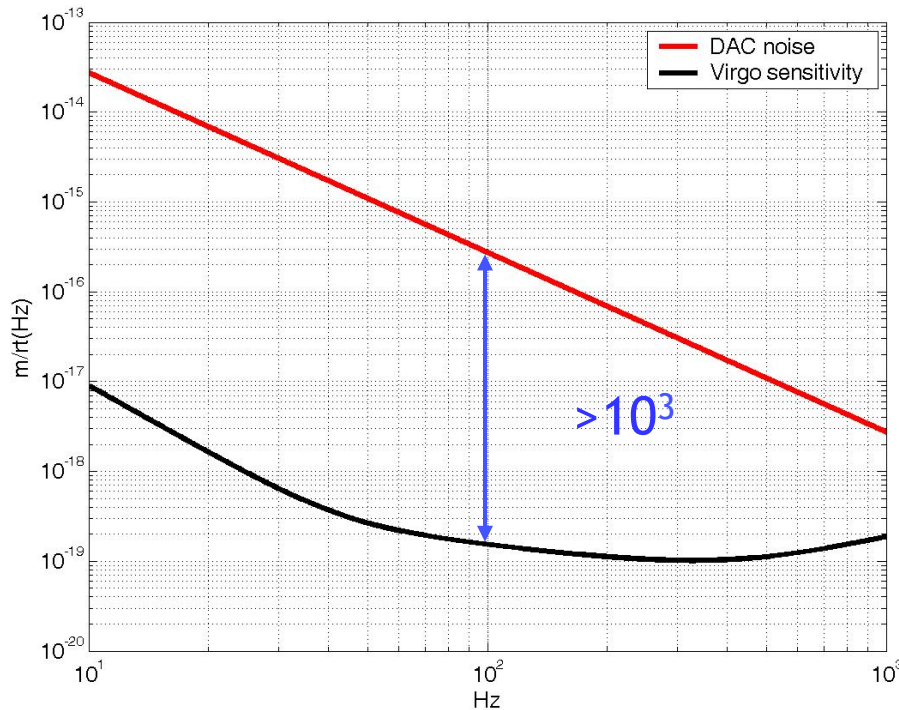


Setting the two conditions of lock means 4 relations among 4 length, the four degrees of freedom of the Recycled Fabry-Perot ITF

- Dark Fringe set point (MICH)
- Differential Arm length (DARM)
- Power Recycling Cavity Length (PRCL)
- Common Arm length (CARM)

- The force needed to acquire the lock is much larger than that needed to keep it. We need momentum to stop the mirror
- Drawback: “Strong actuation” means large electronic noise

Velocity Limits



1) Response of the feedback loop

$$v_{max1} = \pi \lambda B / F$$

with
 $B = \text{loop bandwidth}$

2) Maximum available force
 F_{max}

$$v_{max2} = (\lambda F_{max} / 2mF)^{1/2}$$

3) time to cross resonance > light storage time

$$v_{max3} = \pi \lambda c / (4 F^2 L)$$

with
 $B = \text{loop bandwidth}$

Main peculiarity: the finesse of the recycling cavity changes during the lock acquisition sequence

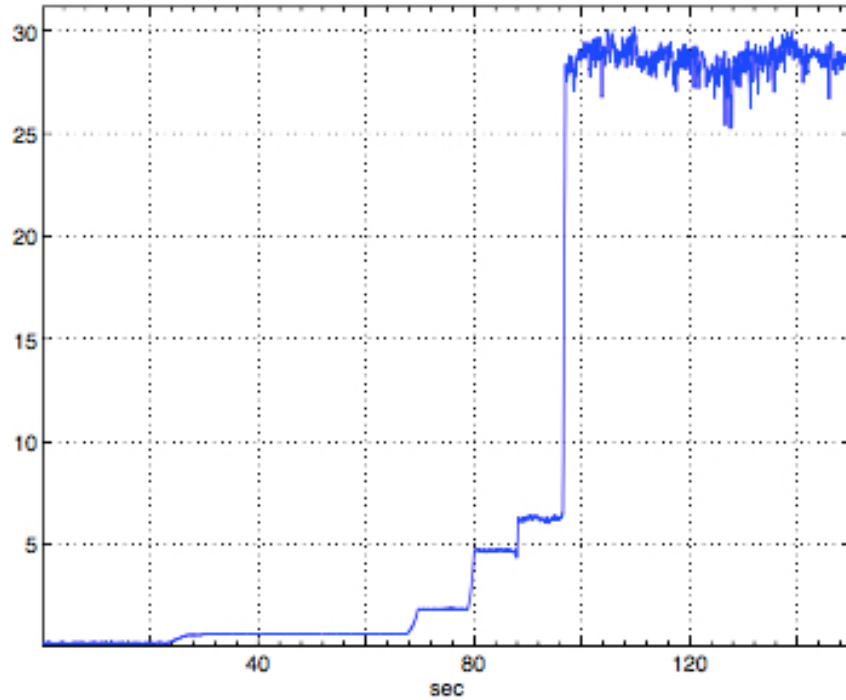
Rational: operate with all the degrees of freedom are weakly coupled and independently controllable, making the control scheme much easier.

All four longitudinal degrees of freedom of the ITF are stably locked in a configuration that is easy to lock: a recycled ITF with a very low recycling gain

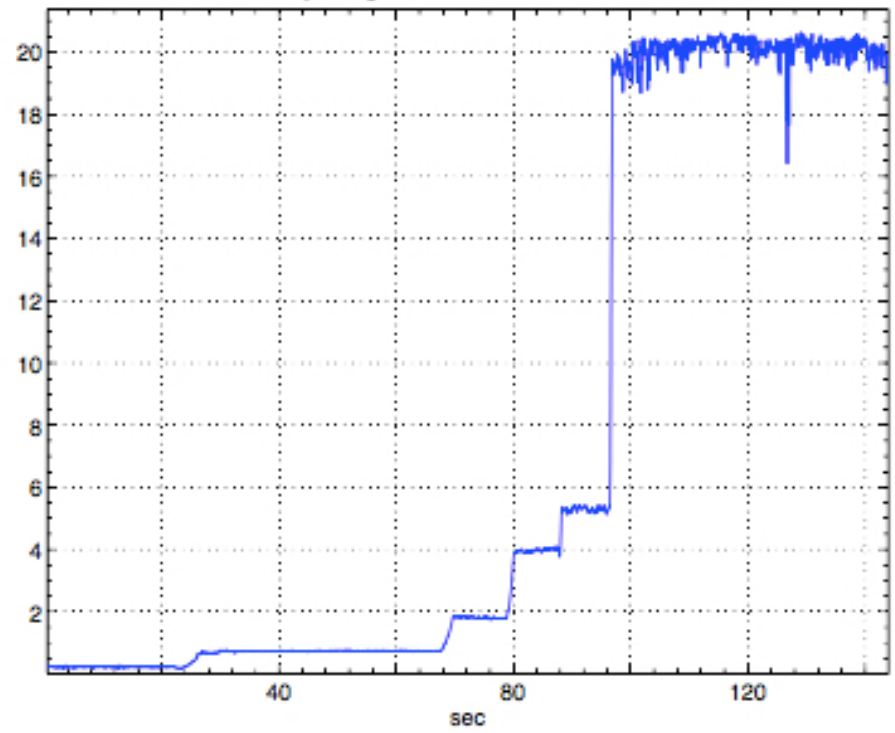
Step (a): locking it on the half fringe, so that a large fraction of light escapes through the antisymmetric port and the power build-up inside the recycling cavity is extremely low.

Step (b): ITF adiabatically brought to the operating point with the Michelson on the dark fringe.

Recycling Gain of the Carrier



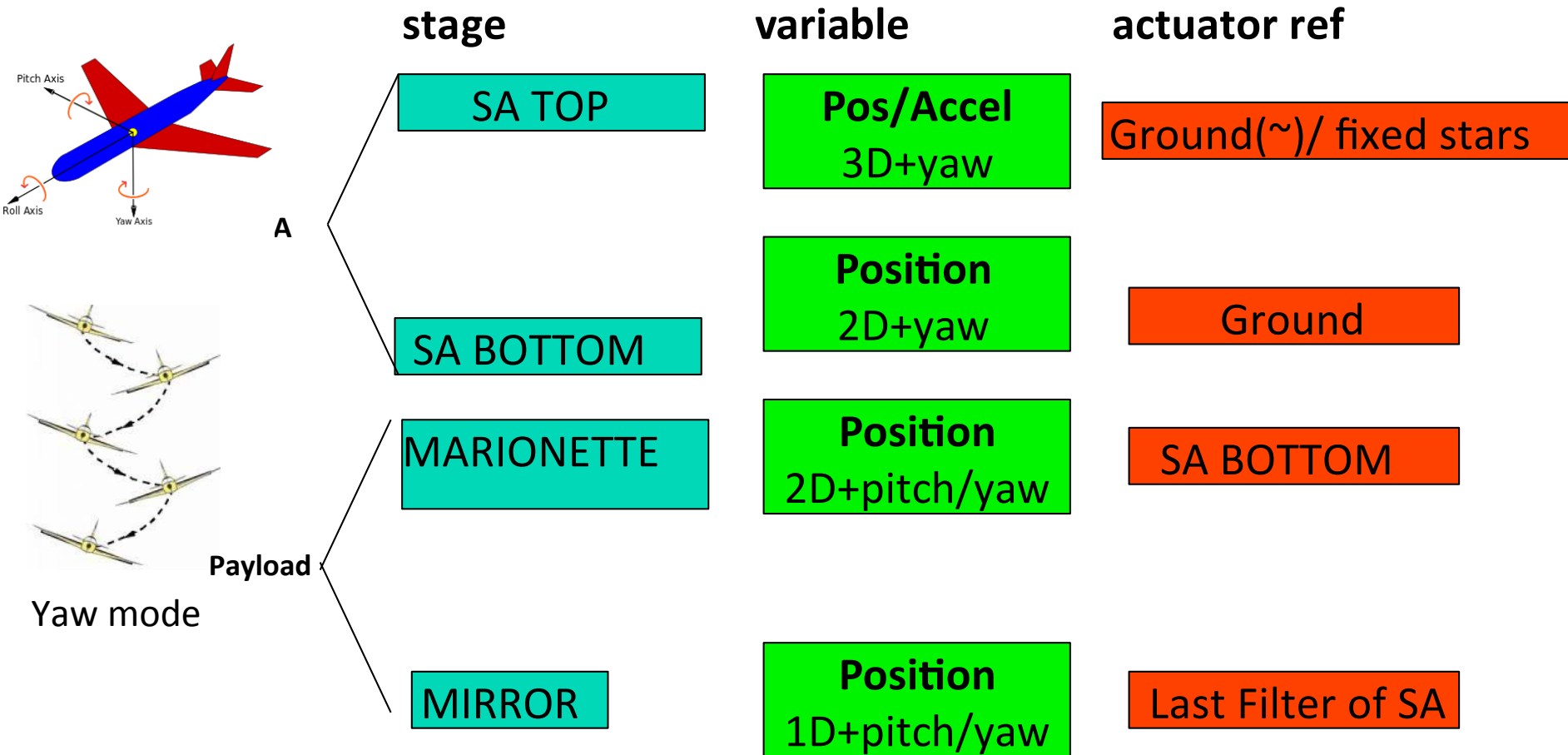
Recycling Gain of the Sidebands



Local Control

The strategy

1) single point suspension, 2) DOF separation, 3) inertial damping, 4) hierarchical control
(small forces close to the test mass)



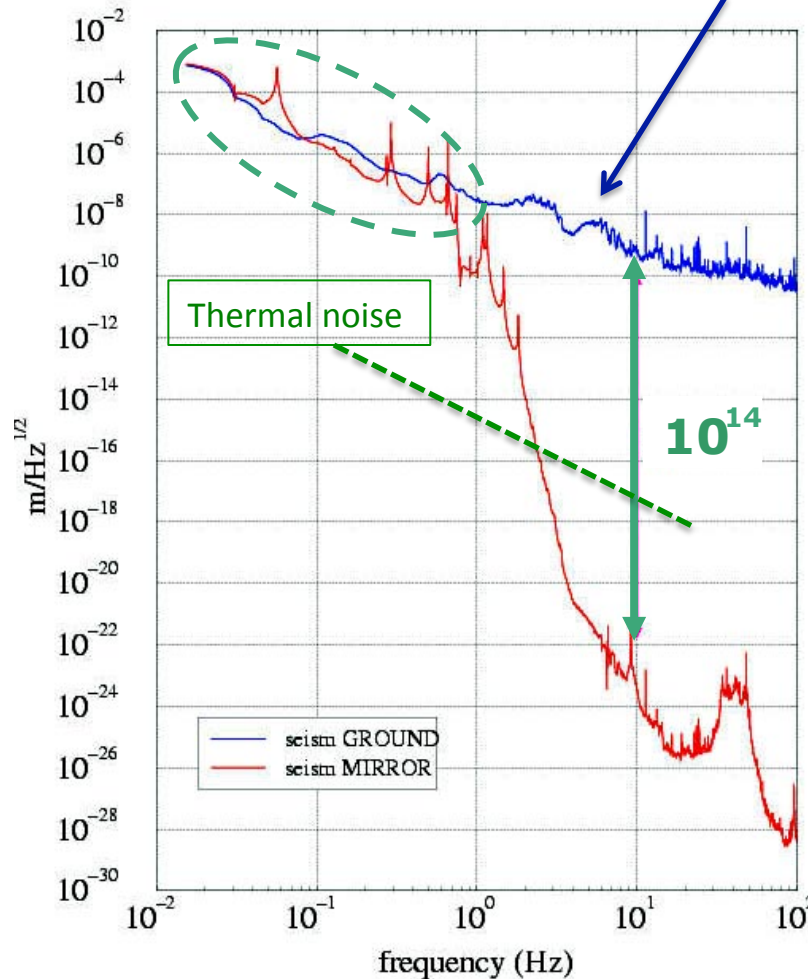
Basic requirements: **sensing and actuation diagonalization**

+

hiearchical control

→ The Superattenuator is a multi-stage pendulum, with **passive attenuation**:

$$S^{1/2}(f) \approx \frac{10^{-7}}{f^2} m / \sqrt{\text{Hz}} \quad (f > 1 \text{ Hz})$$



At lower frequencies the noise is instead totally transferred to the mirror, even amplified by the pendulum resonances



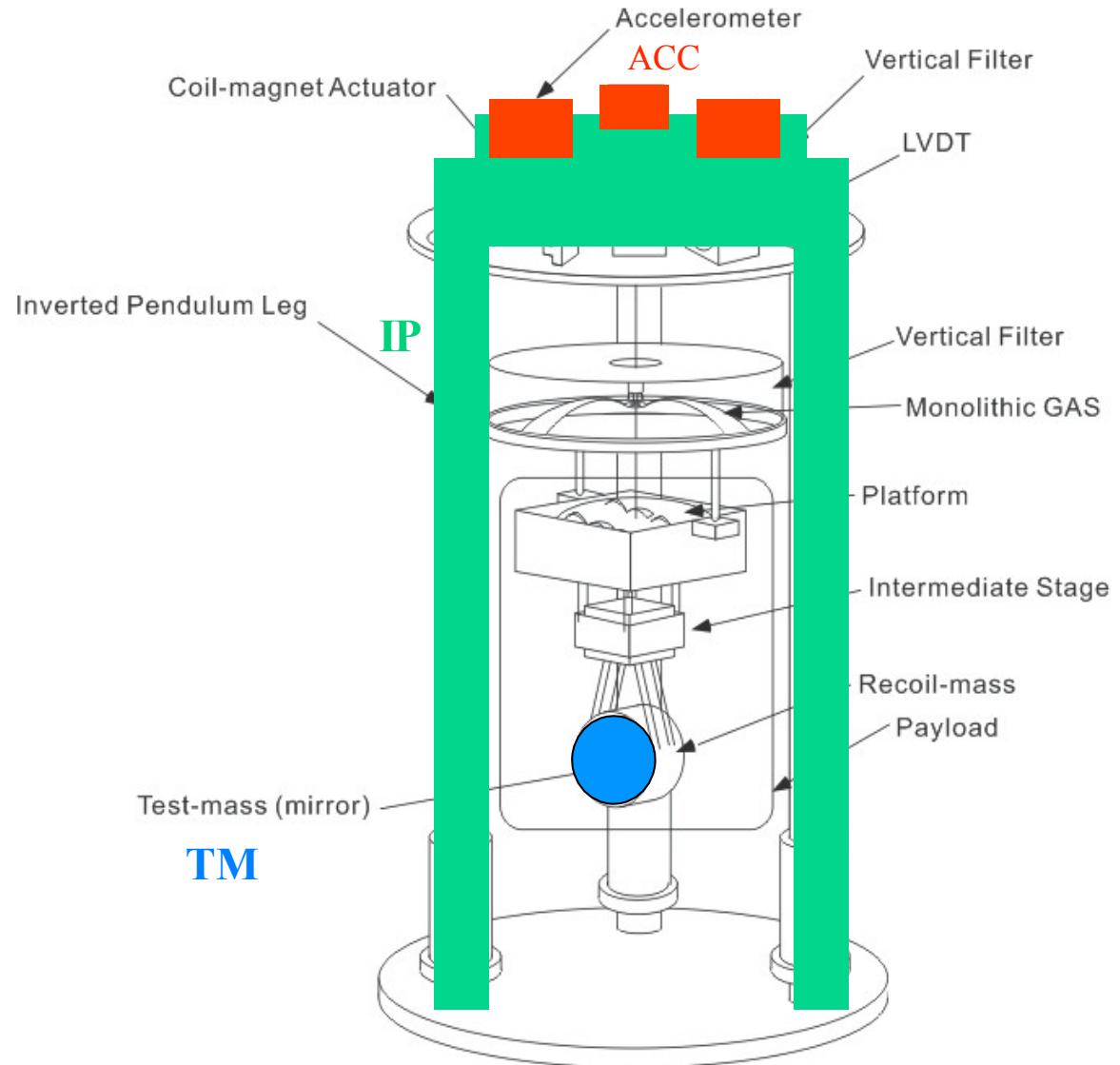
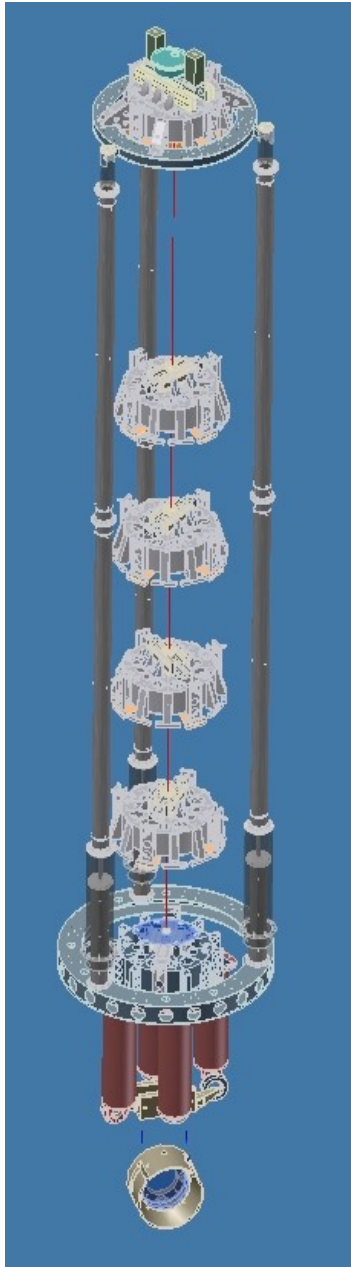
Local active control of the Superattenuator reduces mirror motion below a few Hz



Residual longitudinal motion of the mirror
 $\delta L \sim \mathbf{10^{-6} \text{ m RMS}}$



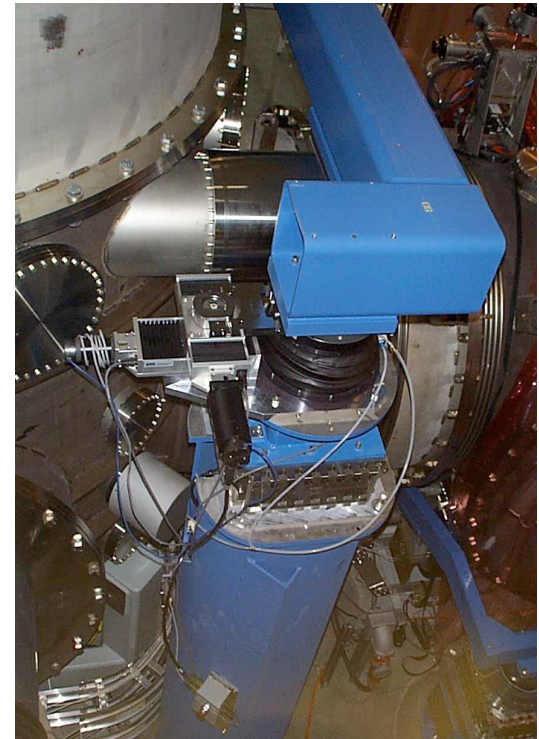
Similar approach in TAMA, GEO and KAGRA



LIGO



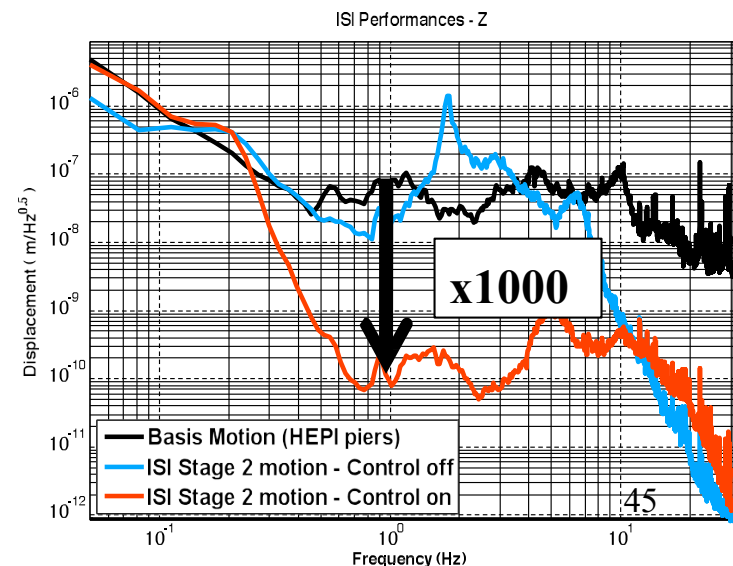
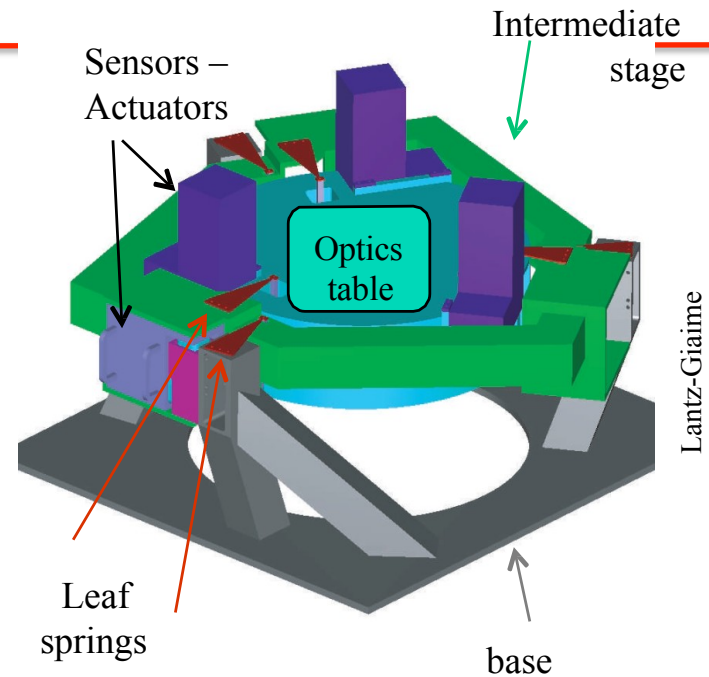
Passive (to reduce noise in sensitive freq. band)

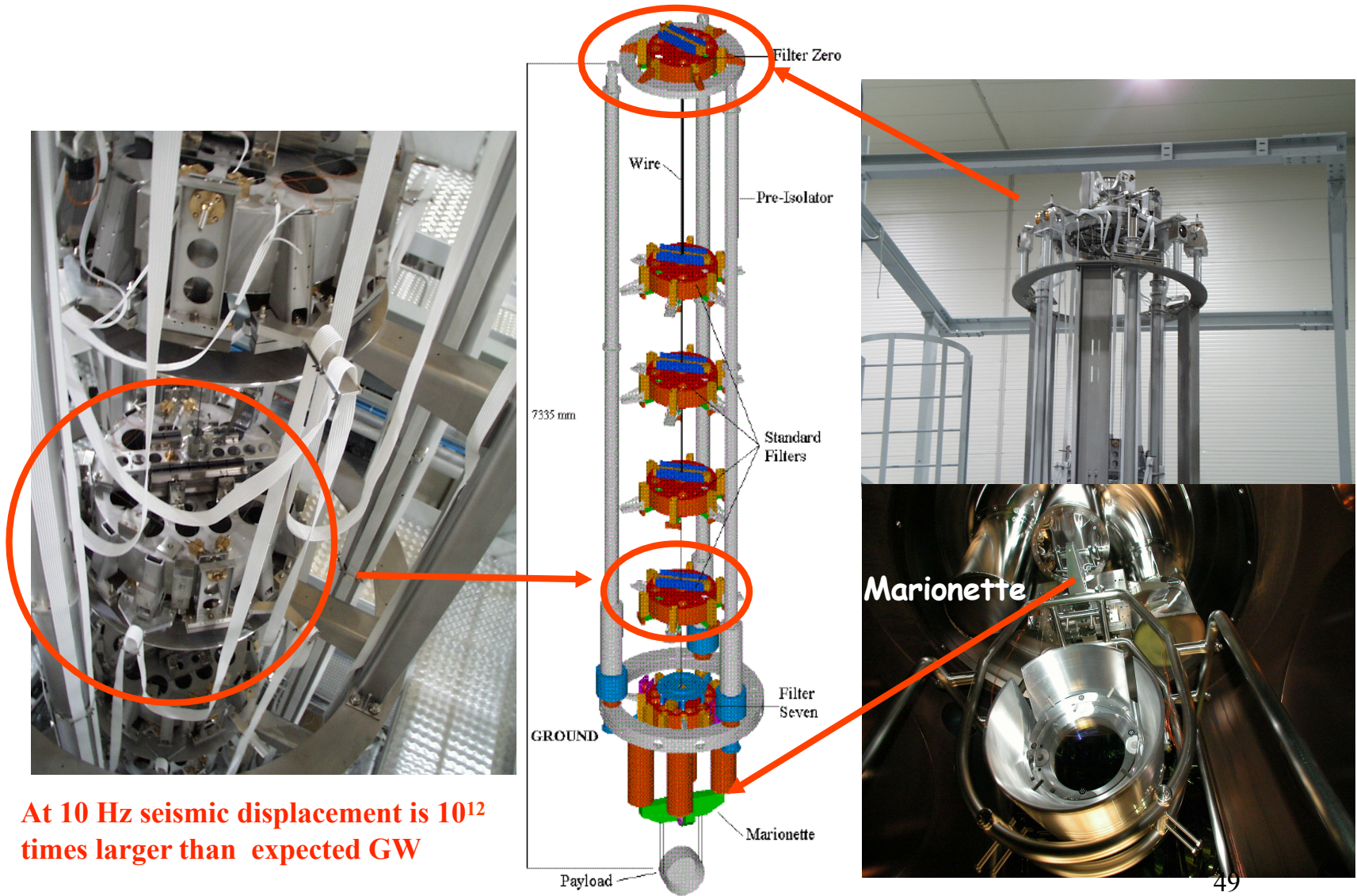


Active (to improve lock acquisition/maintenance)

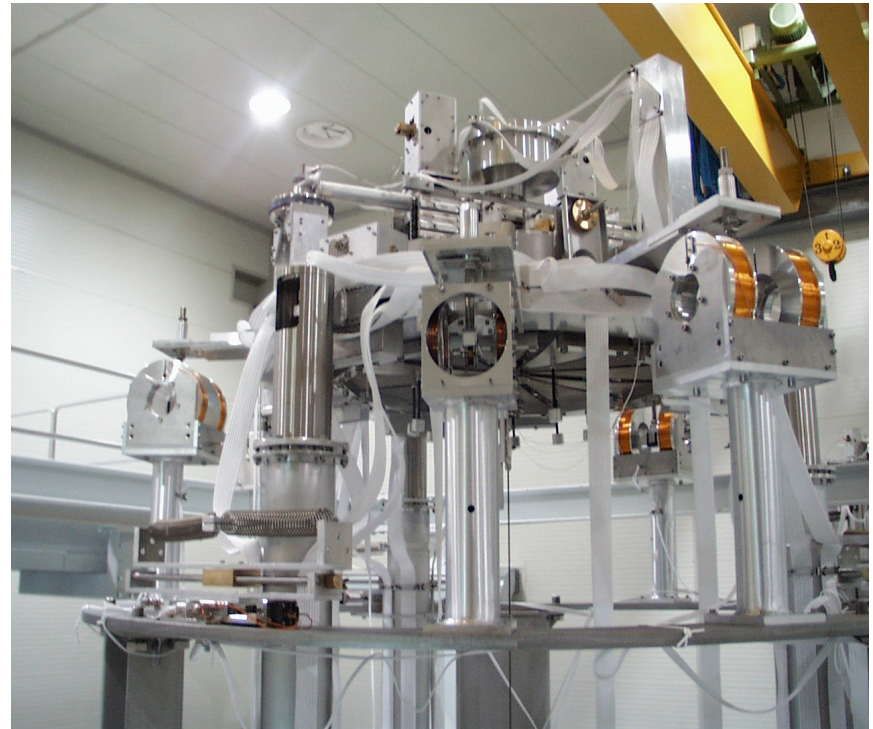
AdV-LIGO seismic isolation: different approach based on active suppression

- One new approach is 'active' suppression, used in Advanced LIGO
 - Low-noise seismometers on payload detect motion in all six degrees-of-freedom
 - Actuators push on payload to eliminate perceived motion
 - Multiple 6-DOF stages to achieve desired suppression, allocation of control
- Challenges in structural resonances, sensor performance
- Nice to have a quiet table to mount lots of stuff on
- Enhanced LIGO using this approach for one chamber per interferometer
 - Meets requirements, will remain in place for AdvLIGO



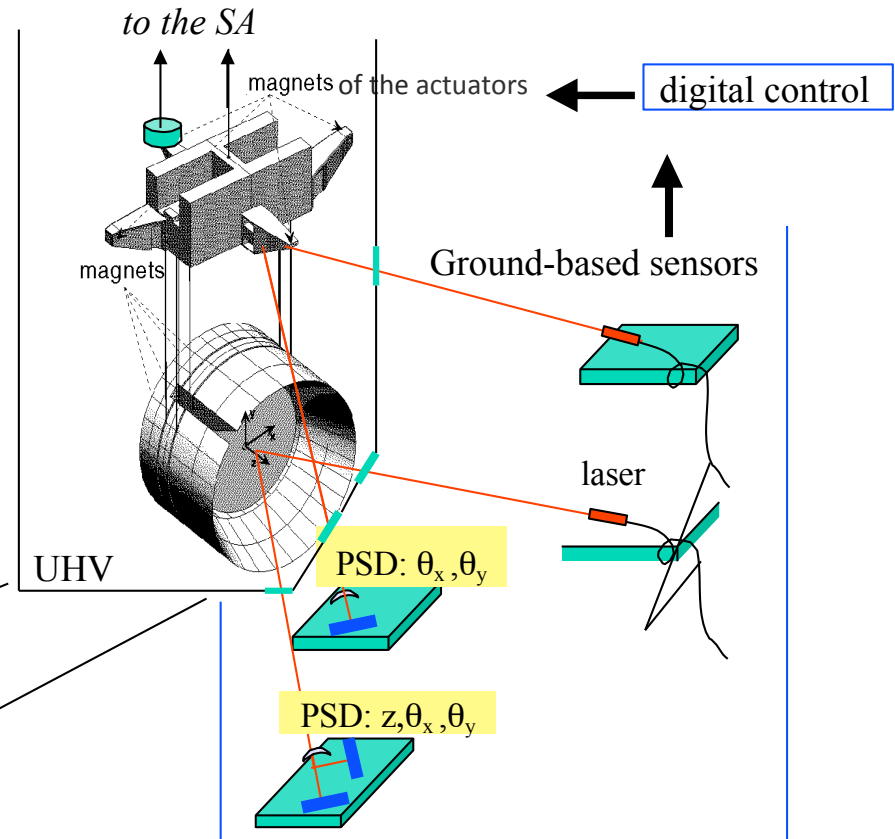
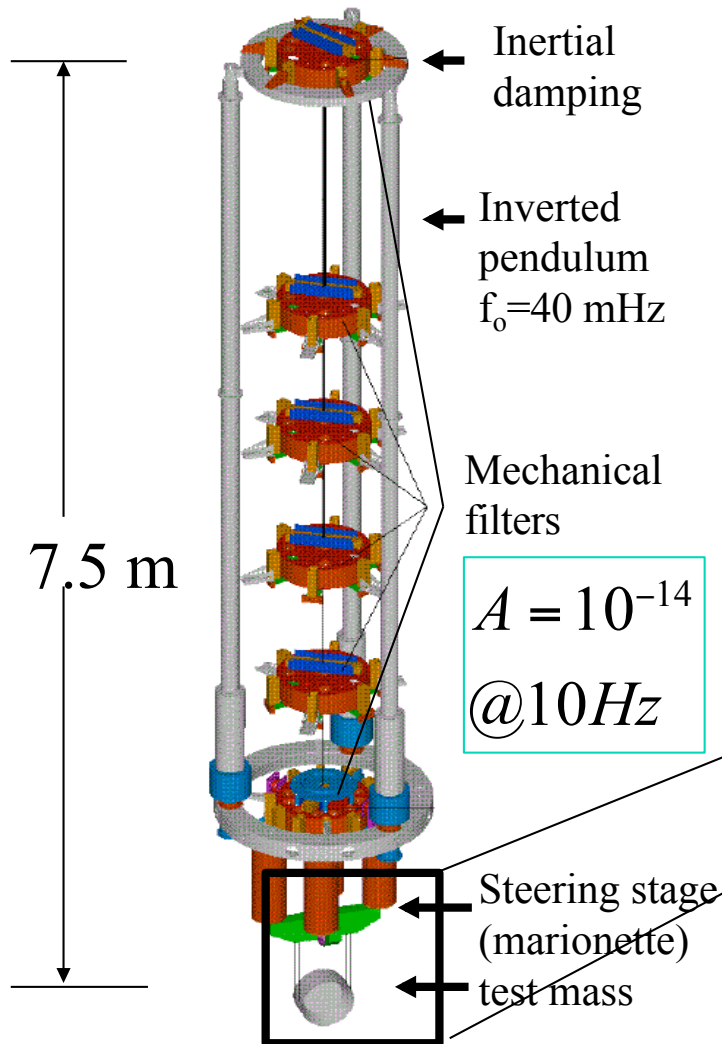


- Inertial sensors (accelerometers):
 - DC-100 Hz bandwidth
 - Equivalent displacement sensitivity: 10^{-11} m/sqrt(Hz)
- Displacement sensors LVDT-like:
 - Used for DC-0.1 Hz control
 - Sensitivity: 10^{-8} m/sqrt(Hz)
 - Linear range: ± 2 cm
- Coil magnet actuators:
 - Linear range: ± 2 cm
 - 0.5 N for 1 cm displacement
- Loop unity gain frequency:
 - 5 Hz
- Sampling rate:
 - 10 kHz



Soft isolator concept:

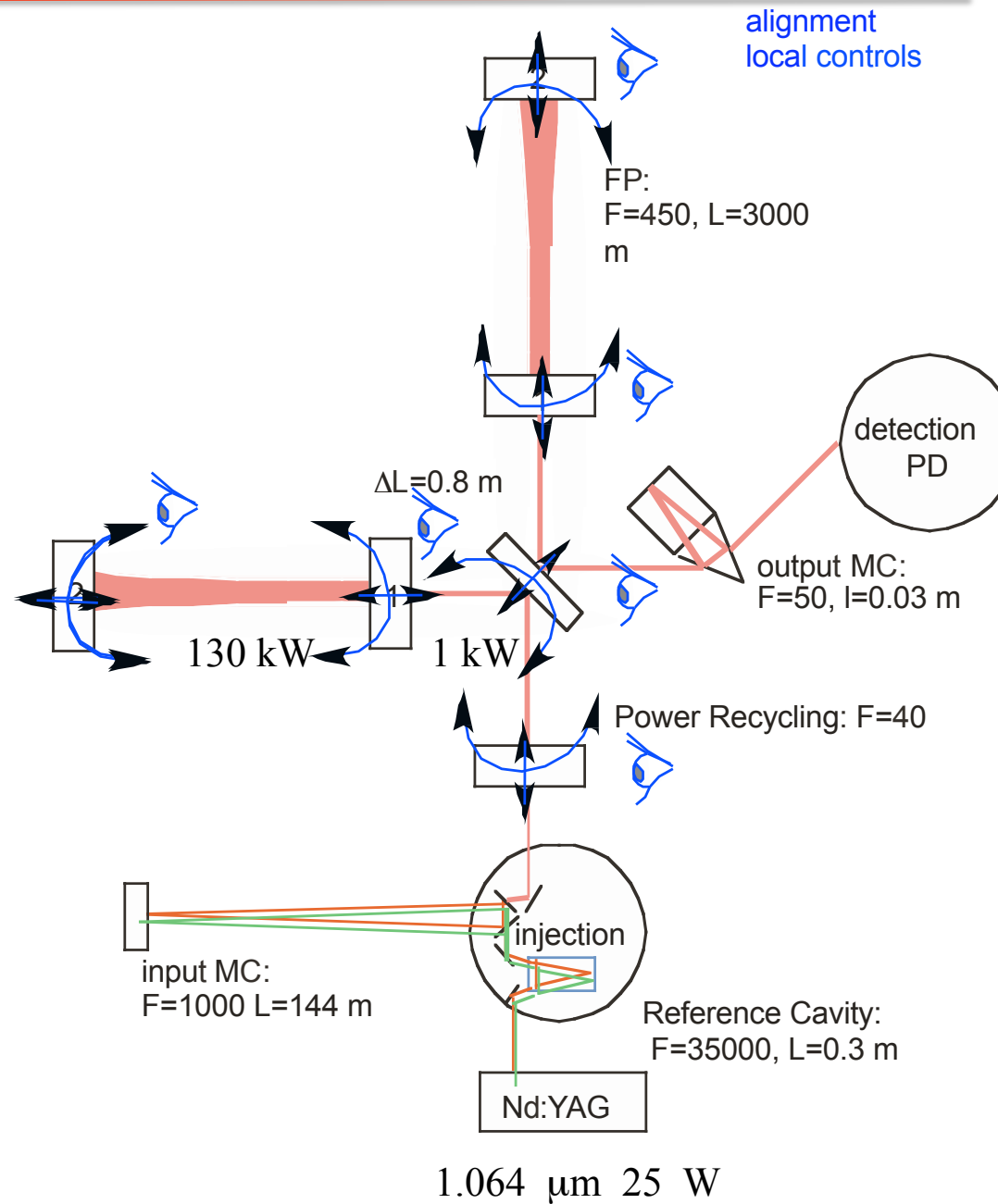
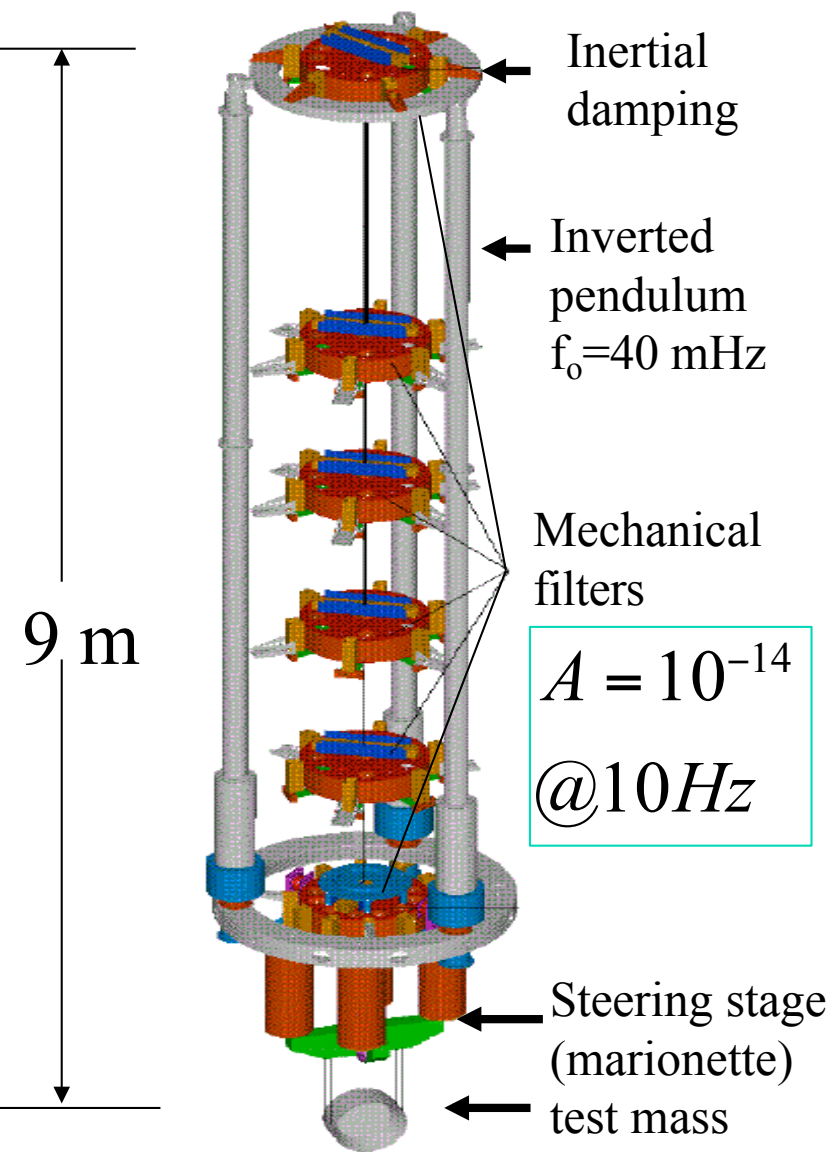
1. very efficient passive attenuation
2. active controls for suspension mode damping



$$\delta\theta_x = \delta\theta_y < 0.1 \mu\text{rad RMS}, \delta z = 0.5 \mu\text{m RMS}$$

Drift (1 h) $\sim 1 \mu\text{rad}$

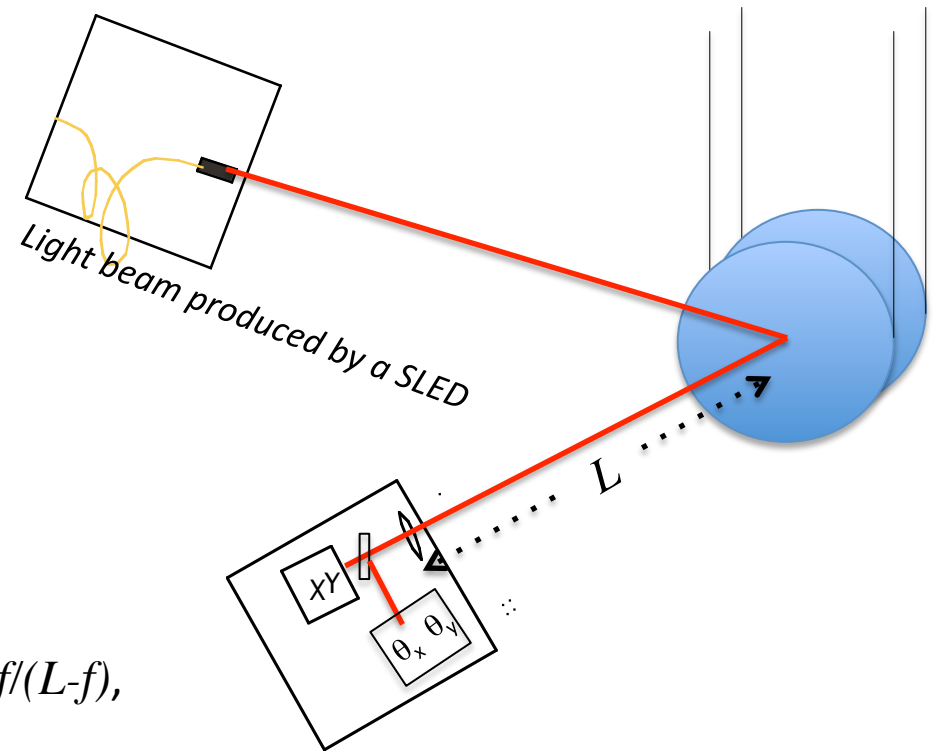
Range: $5 \times 10^4 - 5 \times 10^{-2} \mu\text{rad}$, $10^4 - 0.1 \mu\text{m}$



PSD device on the lens focal plane, $D=f$,
sensitive to mirror rotation

$$x_{PSD} = 2z \sin i \left(1 - \frac{D}{f}\right) + 2\delta i \left[L \left(1 - \frac{D}{f}\right)\right]$$

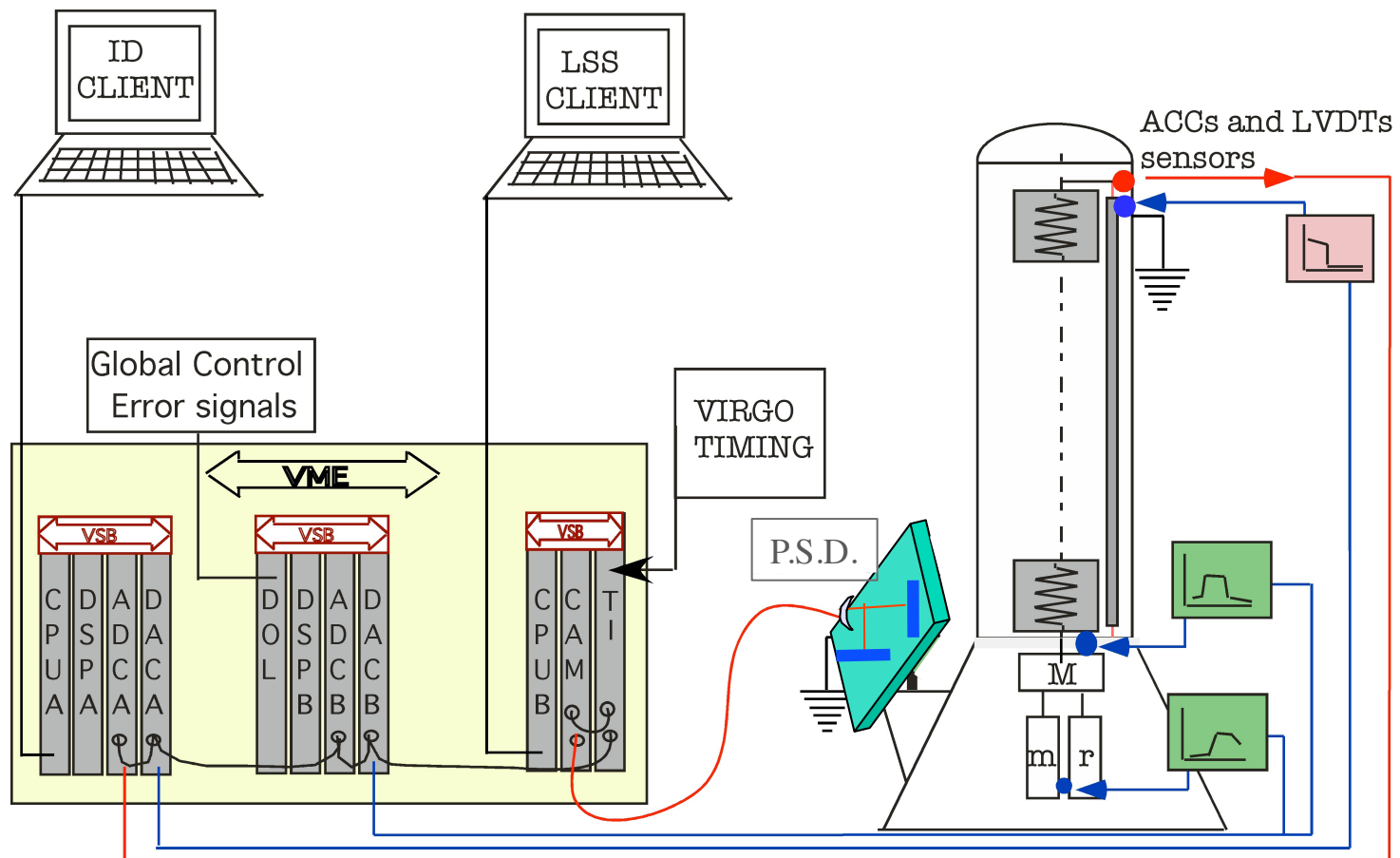
PSD device on the lens image plane, $D=Lf/(L-f)$,
sensitive to mirror translation



- Optical diagonalization of optical levers allow to reduce the coupling among mirror d.o.f. to <1%
- Large dynamics: 50 mrad < 80 nrad thanks to hybrid sensor system

I) Local controls apply corrections to mirror position using local sensors: **swinging interference**

II) Local controls receive error signals from global sensors.
ITF Locked, resonant light





The sensitivity curve

$x_n(t) \rightarrow$ a stochastic variable with gaussian statistic and associated to a stationary process

For a stationary and ergodic process (statistical invariance to the respect of the time translation) the time dependence is limited to the time difference $\tau = t_1 - t_2$.

Auto correlation

$$R_{xx}(\tau) = E[x(t+\tau)x(t)] = \lim_{T \rightarrow \infty} \frac{1}{2T} \int_{-T}^{+T} x(t+\tau)x(t)dt$$

Properties

$$R_{xx}(0) = E[x^2(t)] = \eta^2 + \sigma^2$$

$$R_{xx}(0) \geq R_{xx}(\tau) \quad \forall \quad \tau$$

$$R_{xx}(\tau) = R_{xx}(-\tau)$$

The amplitude of a GW is a strain, a **dimensionless** quantity h_s .
This gives a fractional change in length, or equivalently light travel time, across a detector.

The detector sensitivity is a measure of the capability to extract the signal h_s from the detector noise

This is usually done by optimizing the Signal to Noise Ratio (SNR) by the matched filter (Wiener filter in the case of δ signal),

$$SNR = \rho^2 = \int_{-\infty}^{+\infty} d(f) \frac{H_s^2(f)}{S_n(f)}$$

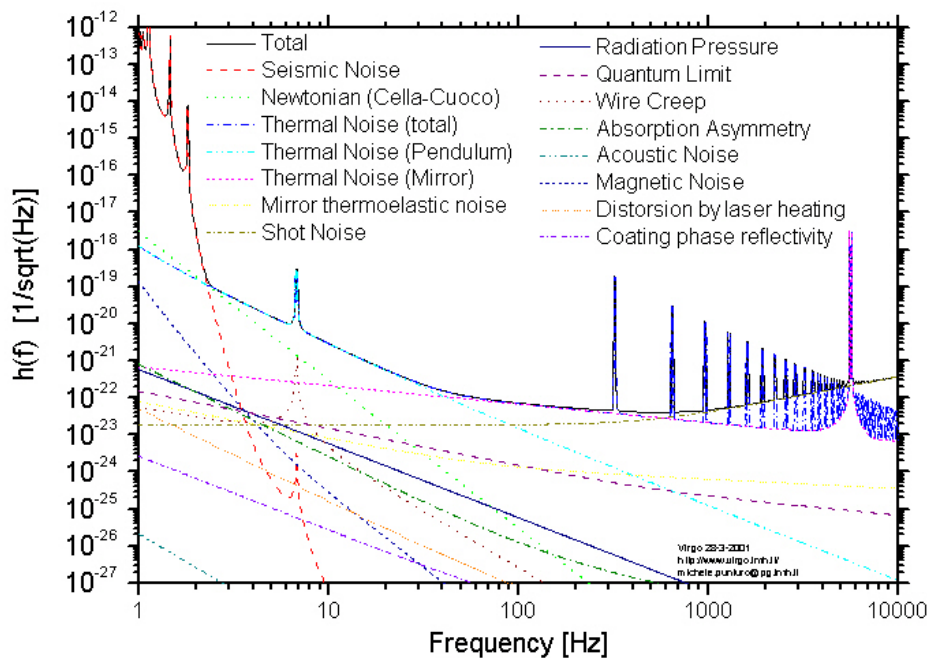
$H_s(f)$ \rightarrow Fourier transform of h_s with dimensions $(Hz)^{-1}$

$$H_s(f) = \int_{-\infty}^{+\infty} h_s(t) \exp(-i2\pi f t) dt \quad S_n(f) = \int_{-\infty}^{+\infty} R_{hh}^{(n)}(\tau) \exp(-i2\pi f \tau) d\tau$$

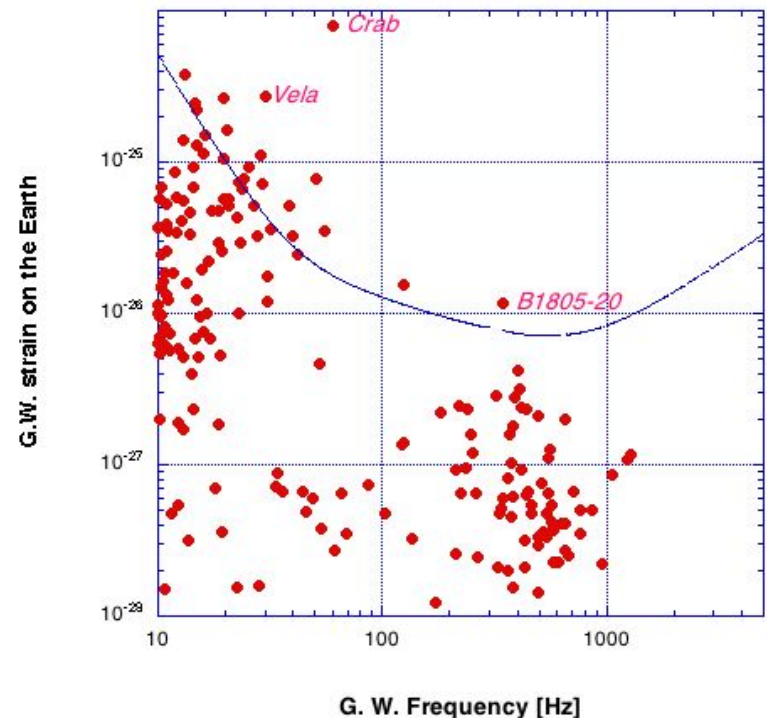
$$h_{rss}^2 = \int [|h_+(t)|^2 + |h_x(t)|^2] dt$$

For a linearly polarized GW, with $H(f)$ constant across the bandwidth Δf

$$h_{rss} \approx H(f) (\Delta f)^{1/2}$$

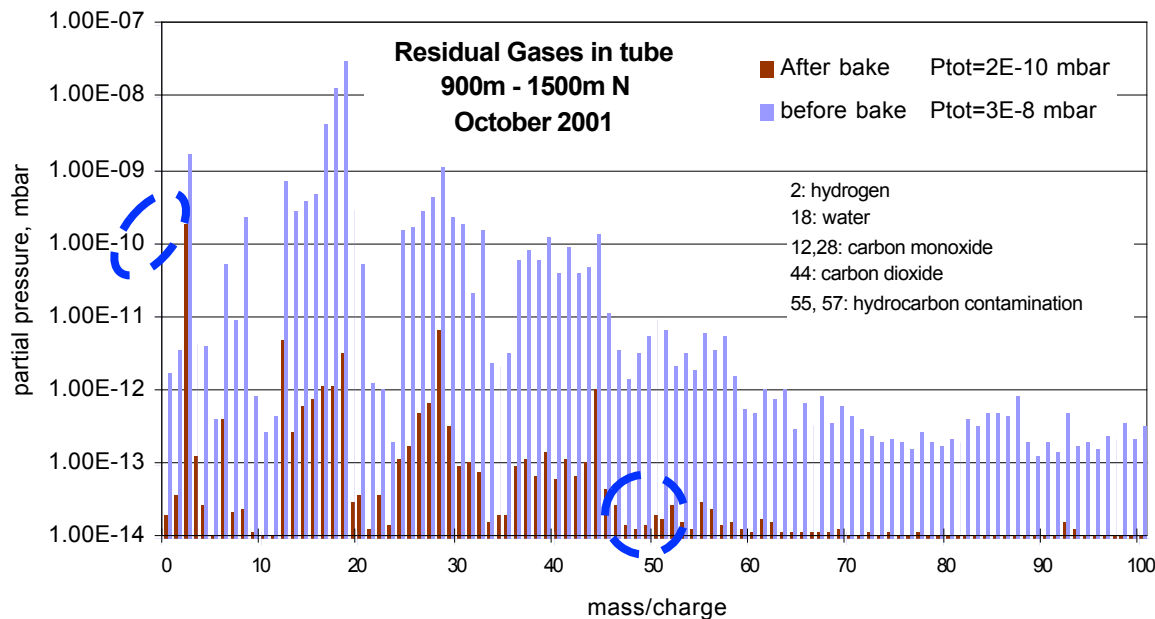


VIRGO sensitivity curve after 1 year of integration time and
 Australian Catalog of Pulsars



The detector

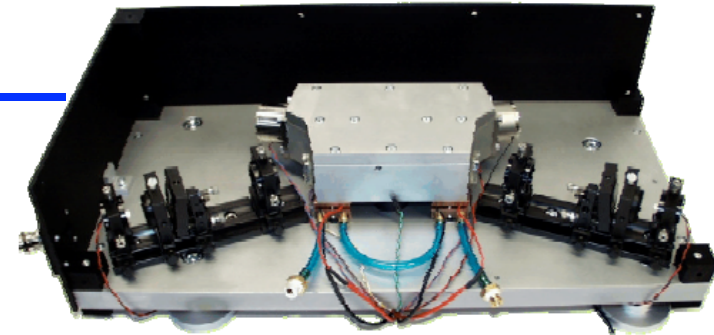
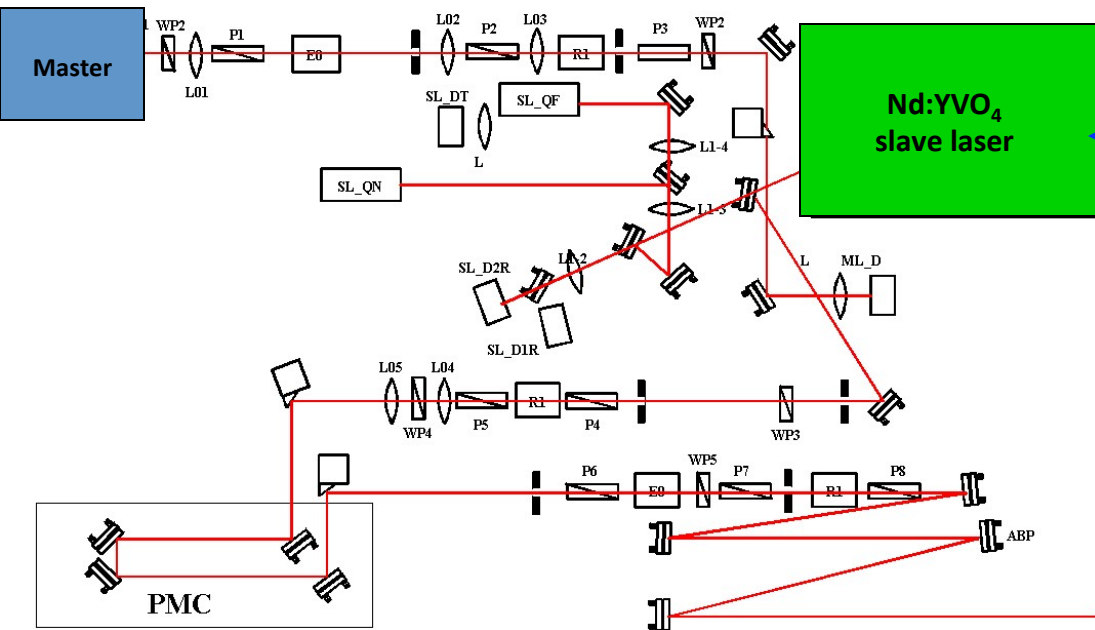
- Requirements:
 - 10^{-9} mbar for H_2
 - 10^{-14} mbar for hydrocarbons
- Status:
 - Whole tube welded, leak-tested and baked and evacuated
 - 160 steel baffles for reflected light shielding installed
 - Vacuum performance within requirements



- High quality fused silica mirrors
 - 35 cm diameter, 20 cm thickness
 - Substrate losses: ~ 1 ppm
 - Coating losses: < 5 ppm
 - Surface deformation: $\lambda/100$ (rms on 150mm)
 - ΔR : $< 10^{-4}$

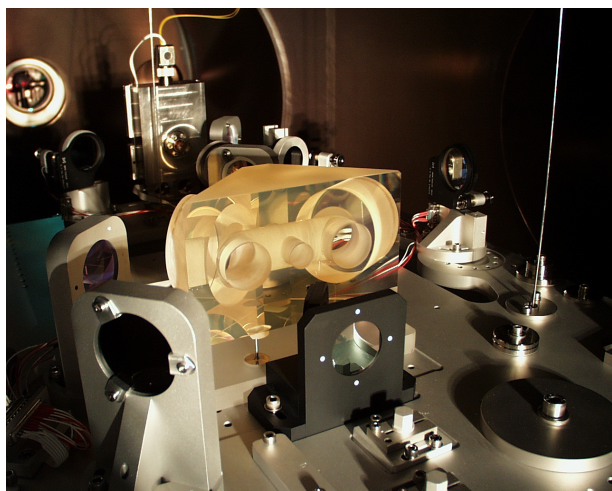


IPN - Lyon

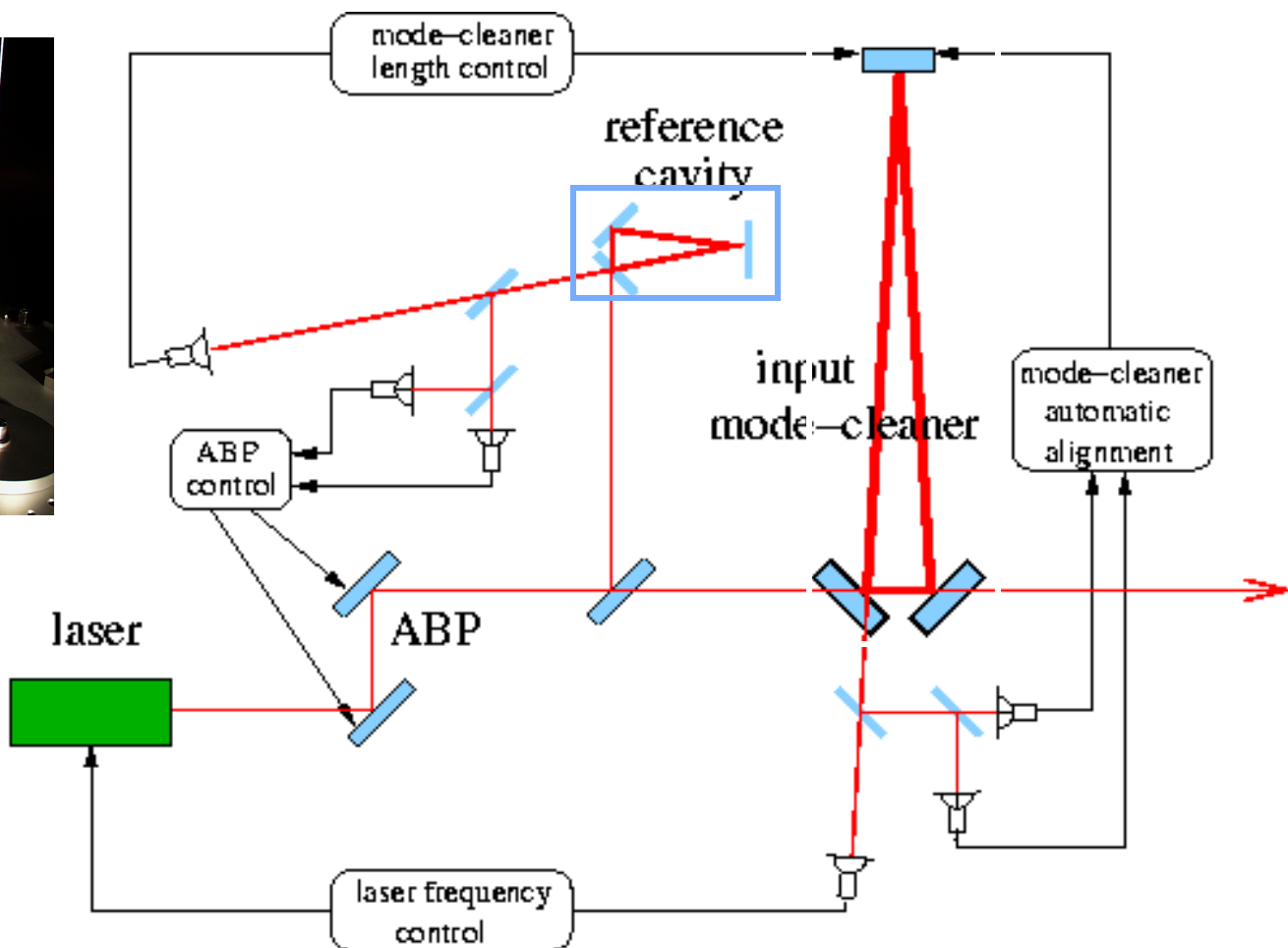


- Nd:YAG master commercial CW single mode (700 mW) @1064 nm
- Phase locked to a Nd:YVO₄ slave (monolithic ring cavity)
- Pumped by two laser diodes at 806 nm (40 W power)
- Output power: 25 W

- Laser frequency locked to the Input Mode cleaner
- Input Mode Cleaner length locked to the reference cavity (monolithic)
- Final stage: ITF common mode as frequency reference

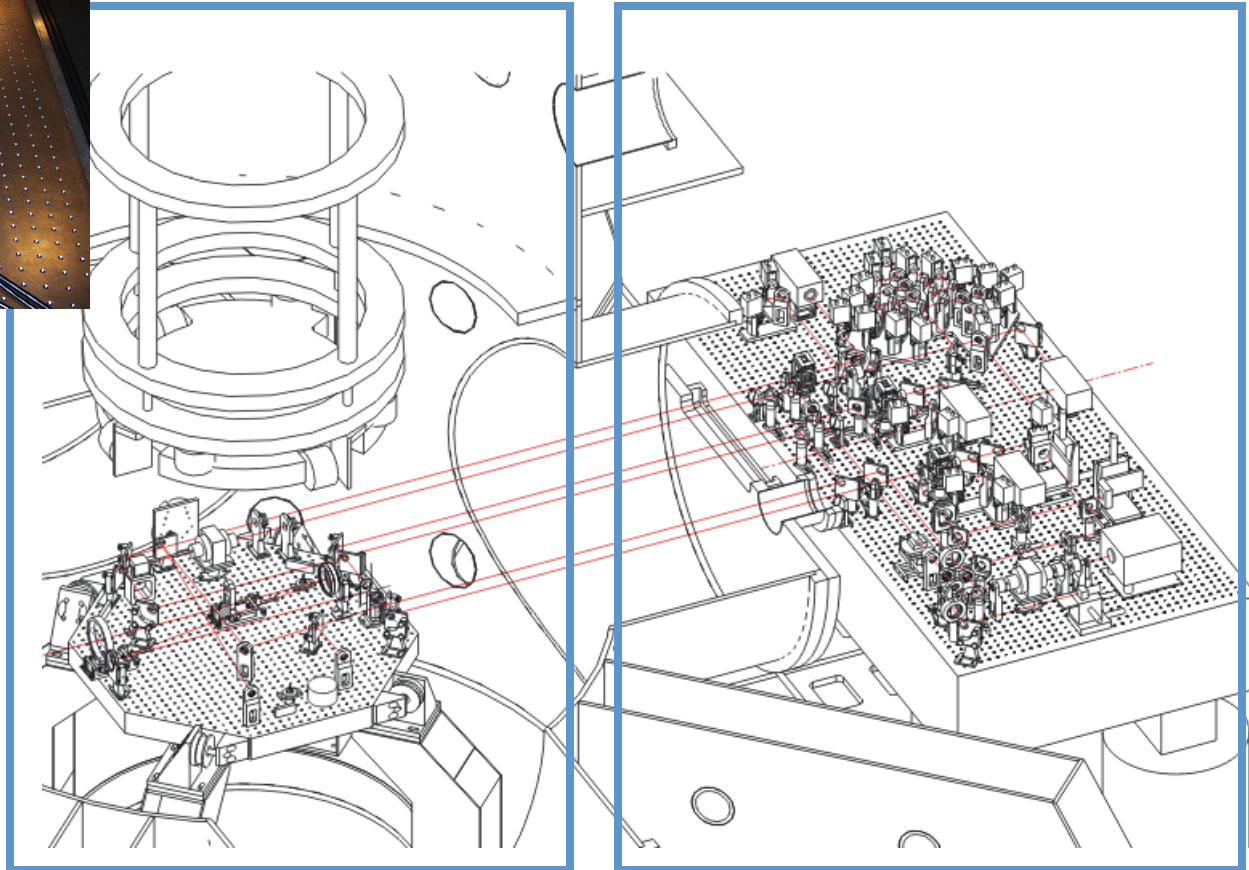


Frequency
stabilization
requirements:
up to 10^{-4} Hz/ $\sqrt{\text{Hz}}$
at 1 Hz



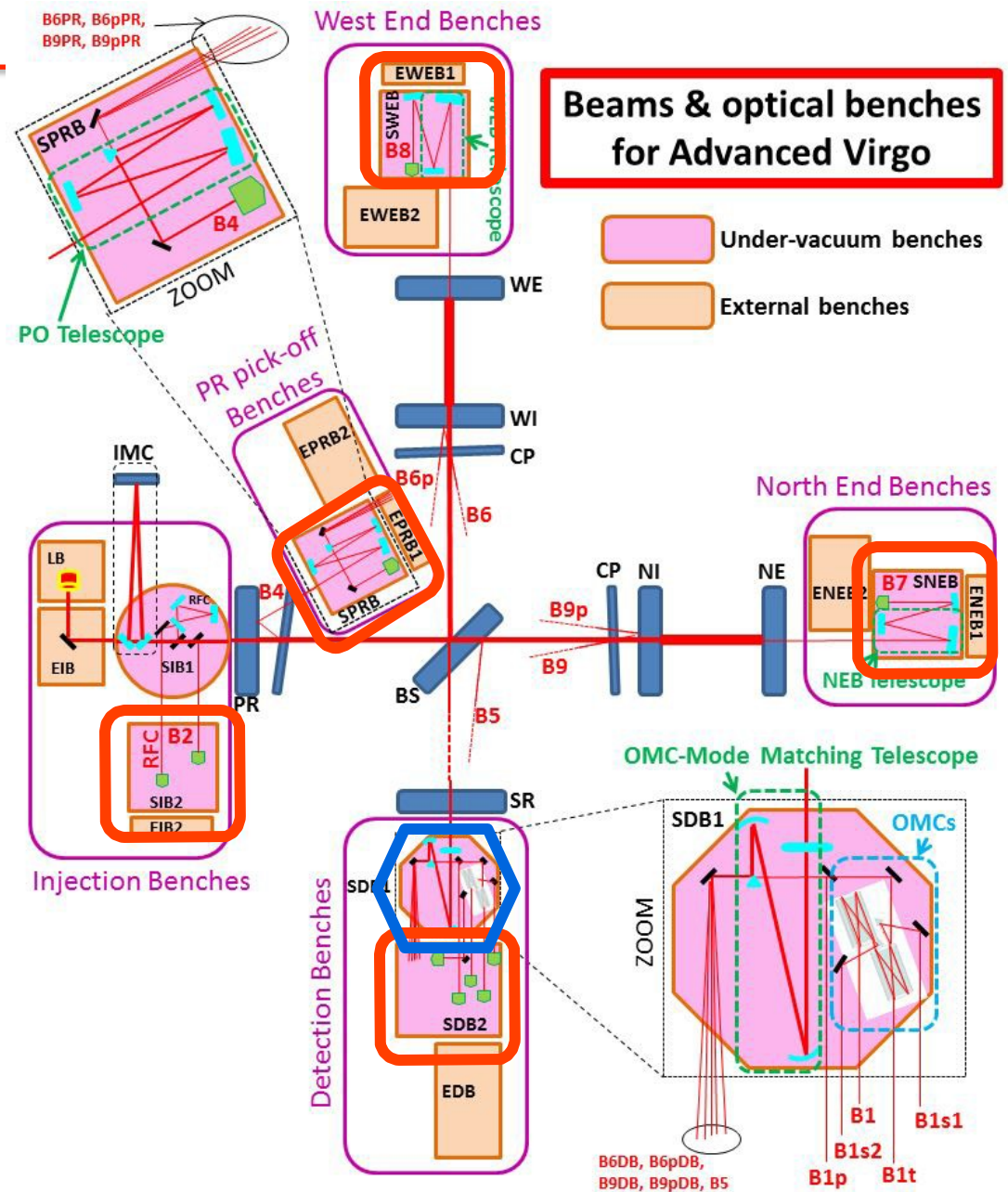
- In-vacuum output mode cleaner (spatial filtering of fake light)

- Detection, amplification and demodulation in air on the external bench (InGaAs photodiodes array)

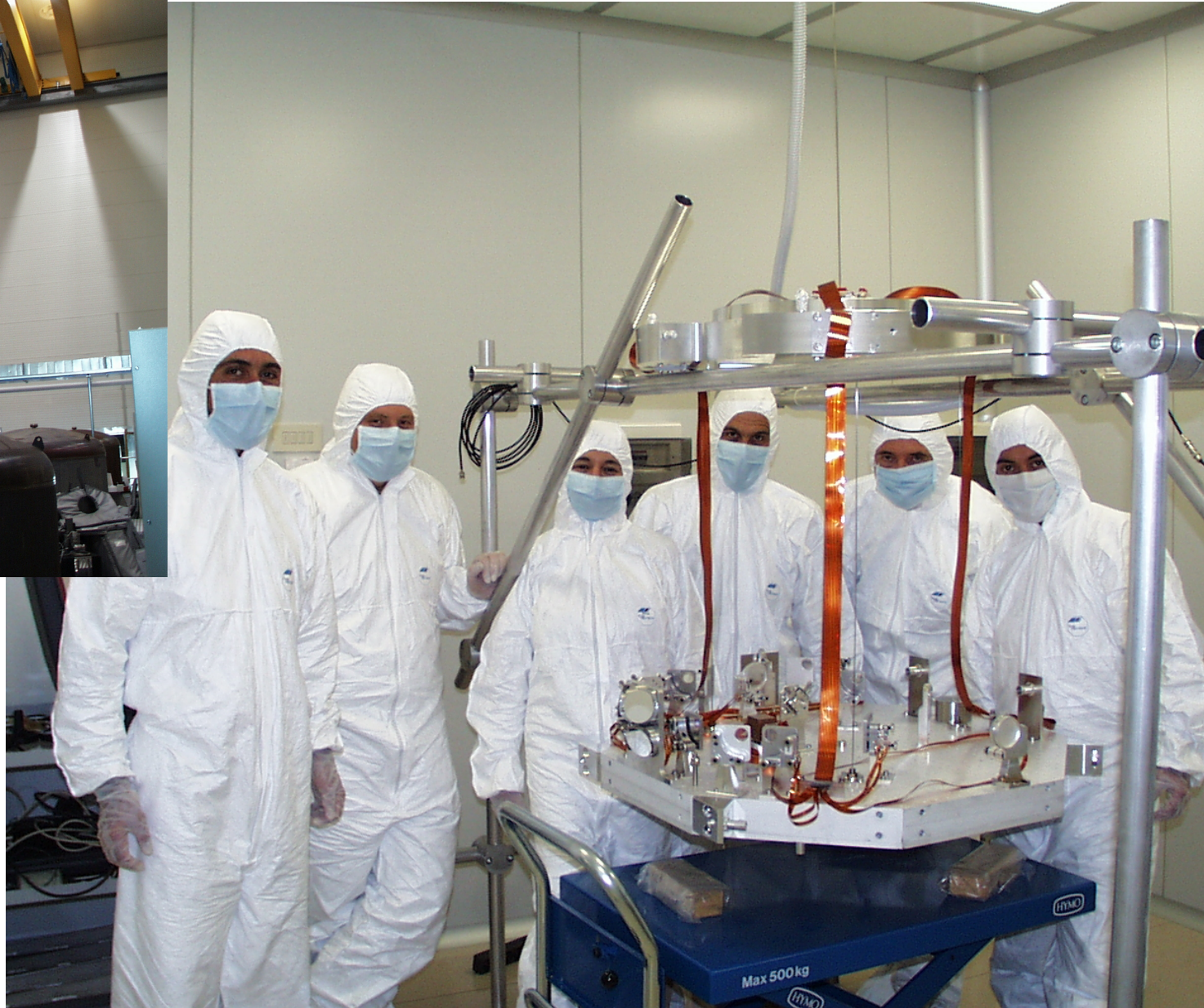


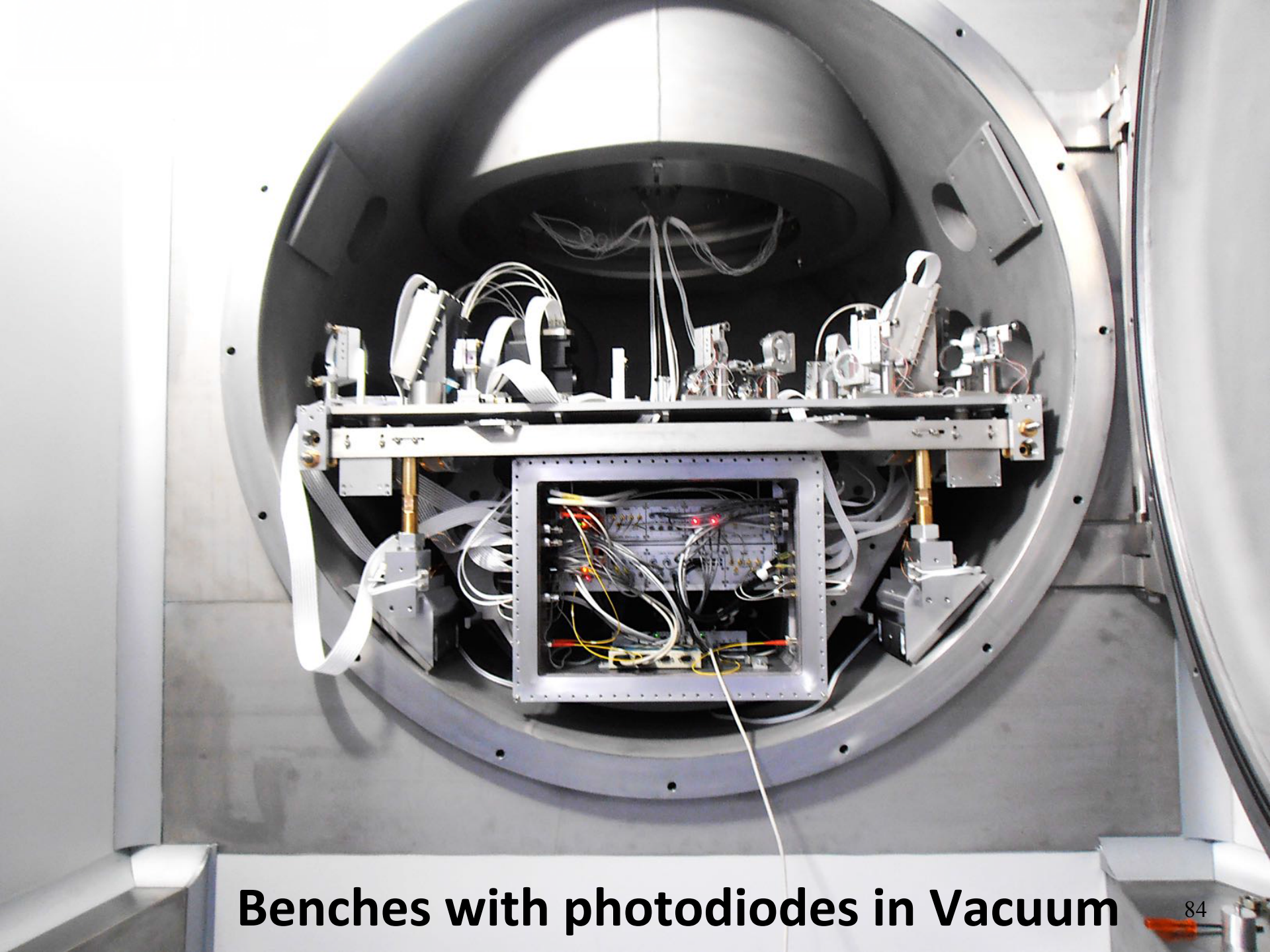
Detection System

- ❑ 6 benches placed in vacuum chamber
- ❑ 2 (SDB1 and SNEB) suspended and controlled
- ❑ SNEB in vacuum

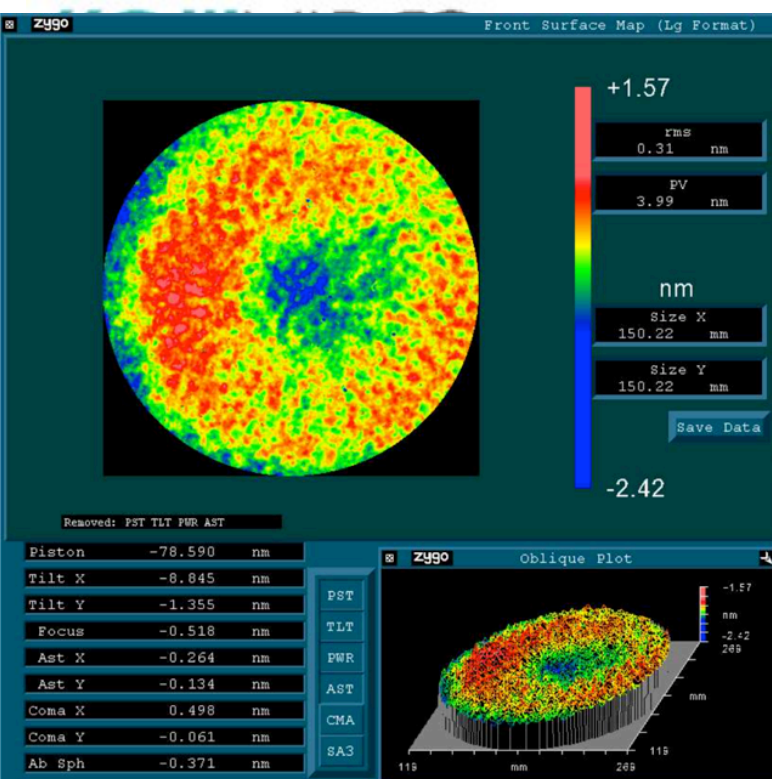


Suspended benches





Benches with photodiodes in Vacuum

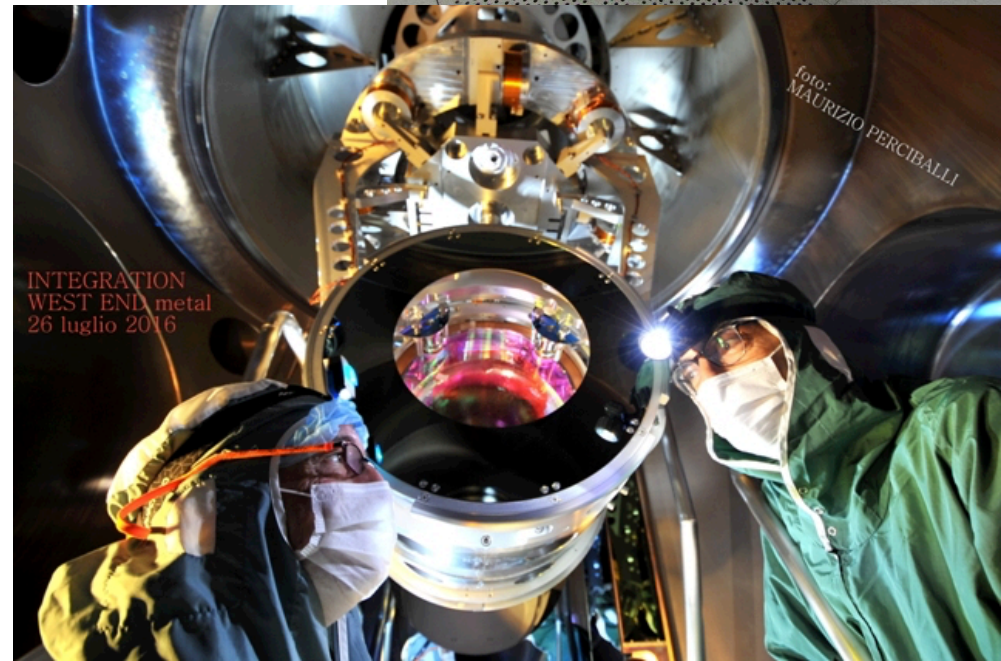


Payloads

Absorption < 0.5 ppm
Flatness \varnothing 150 cm < 0.5 nm

Test mass I ROC 1424 m
Test Mass E ROC 1690 m
Test Mass 42 kg
Diameter 350 mm
Thickness 200 mm

Beam Splitter 50%
Weight 43 kg
Diameter 500 mm
Thickness 100 mm
BS flatness \varnothing 220 cm \leq 1 nm

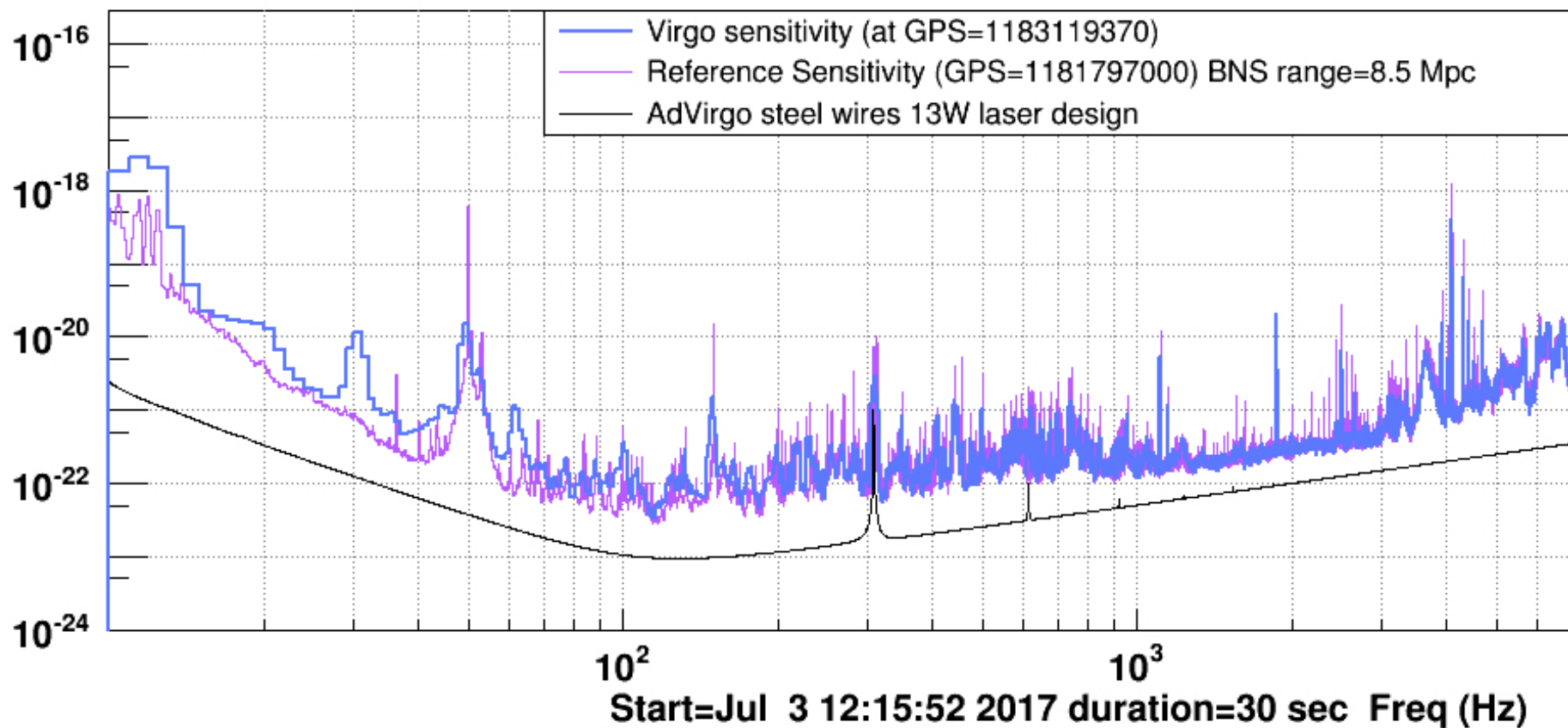


AdV Overview, Part I

Subsystem and Parameters	AdV design (TDR)	Initial Virgo
Sensitivity		
Binary Neutron Star Inspiral Range	134 Mpc	12 Mpc
Anticipated Max Strain Sensitivity	$3.5 \cdot 10^{-24} / \sqrt{\text{Hz}}$	$4 \cdot 10^{-23} / \sqrt{\text{Hz}}$
Instrument Topology		
Interferometer	Michelson	Michelson
Power Enhancement	Arm cavities and Power Recycling	Arm cavities and Power Recycling
Signal Enhancement	Signal Recycling	n.a.
Laser and Optical Powers		
Laser Wavelength	1064 nm	1064 nm
Optical Power at Laser Output	>175 TEM ₀₀ W	20 W
Optical Power at Interferometer Input	125 W	8 W
Optical Power at Test Masses	650 kW	6 kW
Optical Power on Beam Splitter	4.9 kW	0.3 kW
Test Masses		
Mirror Material	Fused Silica	Fused Silica
Main Test Mass Diameter	35 cm	35 cm
Main Test Mass Weight	42 kg	21 kg
Beam Splitter Diameter	55 cm	23 cm
Test Mass Surfaces and Coatings		
Coating Material	Ti doped Ta ₂ O ₅	Ta ₂ O ₅
Roughness*	< 0.1 nm	< 0.05 nm
Flatness	0.5 nm RMS	< 8 nm RMS
Losses per Surface	37.5 ppm	250 ppm (measured)
Test Mass RoC	Input Mirror: 1420 m End Mirror: 1683 m	Input Mirror: flat End Mirror: 3600 m
Beam Radius at Input Mirror	48.7 mm	21 mm
Beam Radius at End Mirror	58 mm	52.5 mm
Finesse	443	50
Thermal Compensation		
Thermal Actuators	CO ₂ Lasers and Ring Heater	CO ₂ Lasers
Actuation points	Compensation plates and directly on mirrors	Directly on mirrors
Sensors	Hartmann sensors and phase cameras	n.a.

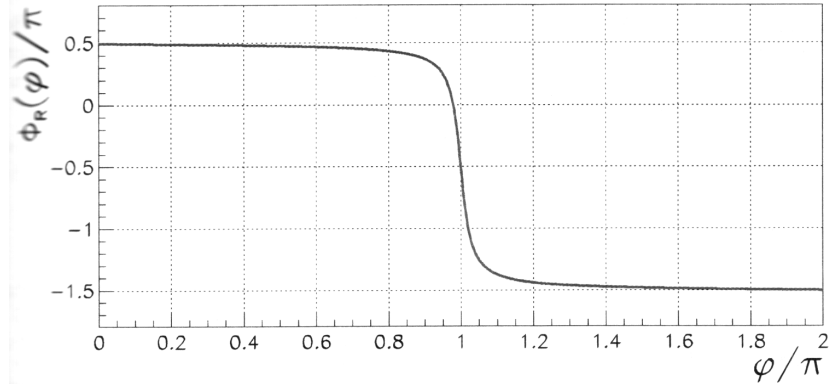
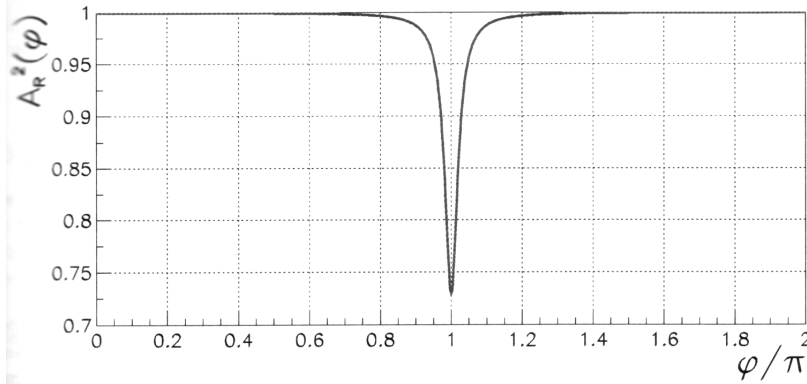
AdV Overview, Part II

Subsystem and Parameters	AdV design (TDR)	Initial Virgo
Suspension		
Seismic Isolation System	Superattenuator	Superattenuator
Degrees of Freedom of Inverted Pendulum Inertial Control	6	4
Test mass suspensions	Fused Silica Fibres (optimized geometry)	Steel Wires
Vacuum System		
Pressure	10 ⁻⁹ mbar	10 ⁻⁷ mbar
Injection System		
Input mode cleaner throughput	>96%	85% (meas.)
Detection System		
GW Signal Readout	DC-Readout	Heterodyne (RF)
Output Mode Cleaner Suppression	RF Sidebands and Higher Order Modes	Higher Order Modes
Main Photo Diode Environment	in Vacuum	in Air
Lengths		
Arm Cavity Length	3 km	3 km
Input Mode Cleaner	143.424 m	143.574 m
Power Recycling Cavity	11.952 m	12.053 m
Signal Recycling Cavity	11.952 m	n.a.
Interferometric Sensing and Control		
Lock Acquisition Strategy	Auxiliary Lasers (different wavelength)	Main Laser
Number of RF Modulations	3	1
Schnupp Asymmetry	23 cm	85 cm
Signal Recycling Parameter		
Signal Recycling Mirror Transmittance	20 %	n.a.
Signal Recycling Tuning	0.35 rad	n.a.



- EXTRA SLIDES

Lock of a single F-P cavity(II)



- The cavity length is such that the carrier (ω_0) is resonant
- The phase of the reflected light is $\arctan(\text{Im } \psi_r / \text{Re } \psi_r) = \pi$ at the resonance

$$i \frac{r_1 + r_2 e^{-2ikl}}{1 + r_1 r_2 e^{-2ikl}} \psi_{in}$$

$$e^{-2ikl} = -1$$

$$i \frac{r_1 - r_2}{1 - r_1 r_2}$$

At ω_0 the light reflected by the cavity has a phase shift of -180° to the respect of the incident light
and $\pm 90^\circ$ for the side-bands (for $\Omega \pm \omega_0$ far from ω_0)

$$E_r(t) = -E_0 \cdot \cos(\omega_0 t) - mE_0 \cos(\Omega t) \cdot \sin(\omega_0 t)$$

When we have a perturbation of the cavity length :

$$E_r(t) = -E_0 \cdot \cos(\omega_0 t + \phi_d) - mE_0 \cos(\Omega t) \cdot \sin(\omega_0 t + \phi_d)$$

Photo-diode output : $|E_r(t)|^2 \propto m \cdot \sin \phi_d \cos \Omega t$

The demodulated signal in phase at Ω contains the information on the length variation